

Astronomy 330



This class (Lecture 23):

Lifetime 2

Matthew Zettinger

Next Class:

Future of Civilization

Adam Flanders

Stefanie Pansch

Papers due today

HW10 is due Thursday.

Music: *Supermassive Black Hole*– Muse

Presentations



- Matthew Zettinger

[Alien Abduction](#)

Outline



- Lifetime of alien civilizations.
 - End of the Sun (Solar System)
 - End of the Universe
- What is L?

Drake Equation



Frank Drake

That's 0.96 Communicating life/year



$$N = R_* \times f_p \times n_e \times f_l \times f_i \times f_c \times L$$

# of advanced civilizations we can contact in our Galaxy today	Star formation rate	Fraction of stars with planets	# of Earthlike planets per system	Fraction on which life arises	Fraction that evolve intelligence	Fraction that communicate	Lifetime of advanced civilizations
20	stars/yr	0.8 systems/star	2 x 0.11 = 0.22 planets/system	0.775 life/planet	0.505 intel./life	0.7 comm./intel.	yrs/comm.

Massive Impacts = Extinctions?



- Asteroids and comets have hit the Earth.
- A major impact is only a matter of time: not IF but WHEN.
- Major impacts are very rare: For an extinction level event, you have to wait millions of years.
- But! For an event that causes major *damage*, you only have to wait only roughly tens to hundreds of years.



So How to Mitigate?



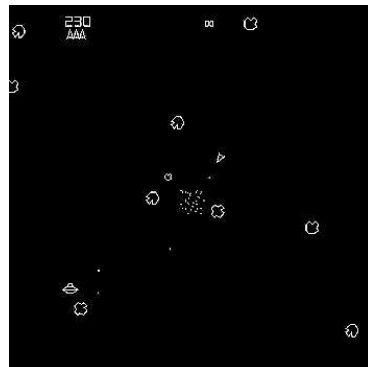
Two main options:

- Destroy
 - Can be problematic
 - Fragment into many pieces (all in the same orbit)..
Have to track hundreds or thousands of objects now!
- Delay
 - Earth is moving 30 km/s, or 1 Earth diameter every 7 minutes.

Blow the Mother Up!



- Typical option discussed is nuclear missiles.
- Might work, vaporizes or at least reduce mass.
- But, need to make sure not to fragment into many still dangerous pieces.
- Imagine twenty-five 50m pieces in the same orbit, would be hard to stop!



Blow-Up Job



- Other option is to blow up a nuclear weapon near the asteroid.
- But not too near to fragment it.
- Imparted energy could be enough to change orbit.
- Neutron bomb (nuclear blast where large fraction of energy is in neutrons) is thought to be most efficient, biggest transfer of energy maybe only chance for last minute threats.



<http://www.projectrho.com/rocket/rocket3x.html>

http://www.youtube.com/watch?v=XPS-m_sI7_k

Kinetic Energy Deflection



- Impact the asteroid or attach rockets.
- May still fragment, but most have impacts, so less likely
- Actually an ESA mission to test this is occurring in 2013 or 2015!
- The aptly-named Don Quijote mission



http://www.esa.int/SPECIALS/NEO/SEMZRZNVGJE_1.html

Don Quijote



Two components:

- Sancho: orbits and accurately measures position
 - Plus the Autonomous Surface Package Deployment Engineering eXperiment, which checks out the impact site
- Hidalgo: impactor (10km/s)

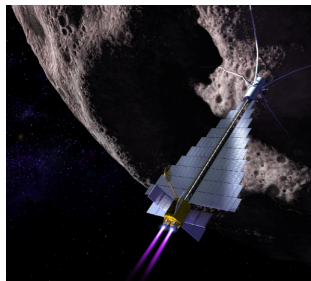


http://www.esa.int/SPECIALS/NEO/SEMZRZNVGJE_1.html

The Ole' Space Tug



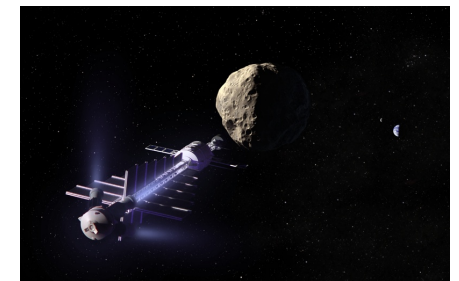
- Put a rocket on the asteroid!
- This can eventually move the rock, but
 - Rockets don't provide too much thrust
 - Will likely need many steerable rockets.
 - Remember that asteroids are rotating!
 - How to attach to a tumbling, rotating asteroid that may only be a big pile of rubble?



Gravity Tractor



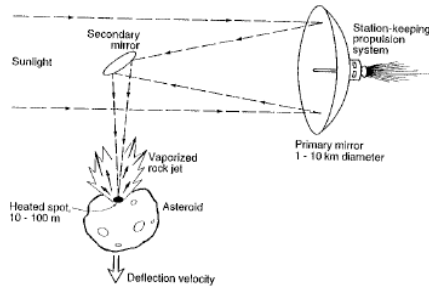
- Put an object near the asteroid!
- Using gravity, the asteroid is attracted to spacecraft.
- Spacecraft uses rockets to keep away, so slow pull.
- Would take ~10 years for moderate mass asteroid
- Works no matter the composition– rubble piles not fragmented.



Focus the Sun on it!



- Use the Sun to melt the asteroid surface.
- This removes material and creates a jet.
- <http://www.youtube.com/watch?v=dcqFy1zjdys>

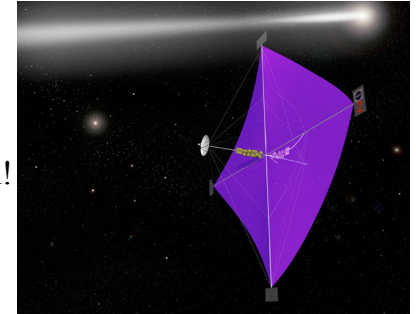


<http://www.lpl.arizona.edu/~jmelosh/HazardsDeflect.pdf>

Other Propulsion: Light Sails



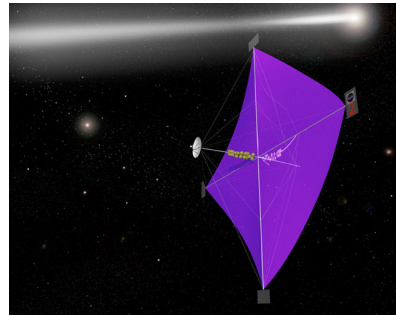
- Imagine a space sailboat but with photons of light hitting the sails and pushing it forward.
- Photons have energy but no rest mass.
- But, they do carry momentum!
 - It is related to the energy such that $p = E / c$
- So, such a craft is not propelled by solar winds!
- But by light bouncing off, like a mirror.



Other Propulsion: Light Sails



- Attach to an asteroid?
- It can work, but it would be slow.
- How to attach to a tumbling, rotating asteroid that may only be a big pile of rubble?



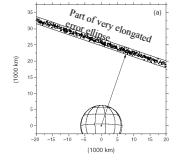
Other Ideas



- Paint it or wrap it in a reflective coating.
 - Hard to do, tumbling asteroid again, plus only pushes in direction opposite to Sun.
- Opposite idea is to sprinkle with soot to reduce Sun pressure (also Yarkovsky effect).
- Asteroid braking. Perhaps a cloud of steam in front of the asteroid to slow it some.
 - Steam? Nuke a comet.

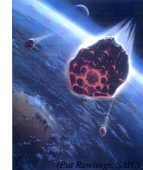


Common Misperceptions

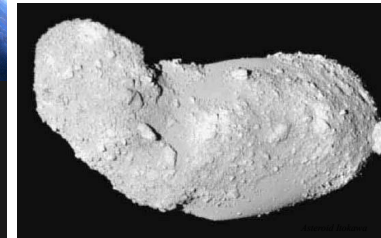


- Long waiting time until next impact
 - Instead, we should think of *chances* of disaster and our responsibilities “on our watch”
- Judging consequences quantitatively
 - Civilization-ending impact vs. K/T mass-extinction
 - “one death” vs. 100 deaths/yr vs. 3000 9/11 dead vs. we will *all* die in next 100 years (what are our values?)
 - Shoemaker-Levy 9 Jupiter impacts overshadowed the Rwanda genocide in the news (July 1994)
- “Blow it up” on the way in
 - Movies misrepresent reality of decades lead-time
- NEA is “on an impact course with Earth”
 - NEA discovery process, error ellipses, NEA orbits the Sun many times before impact: *not intuitive!*

Asteroids are Not Likely to Destroy our World...



- ...but we can contemplate the NEO hazard as the most extreme environmental disaster, and put the lesser, more likely ones into context...
- ...and distinguish between societal issues like global warming and true, sudden catastrophes.
- Many threats to society and our lives (flu, war, famine... global warming) are here today.
- Asteroids *are* in our future...as places to travel to, as fuel stations for a spacefaring civilization ...let's hope they don't come to us first!



4. Natural Catastrophes



3. Stellar Evolution

- The Sun is halfway through its lifetime on the main sequence.

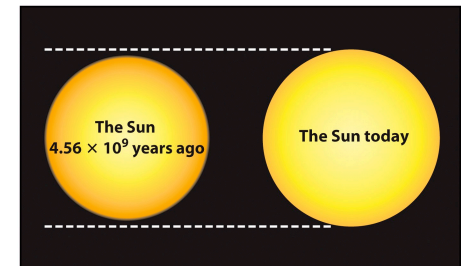


<http://www.astroimages.net/Media/SolarSys/AR03.html>

Life of a Low Mass (Sun-like) Star



- Most of its life is spent in the happy pursuit of burning $H \Rightarrow He$
- With time, luminosity and temperature evolve gradually in response
 - Stays on the Main Sequence, but still evolves..
- The Sun is now 40% brighter and 6% bigger than zero age MS.

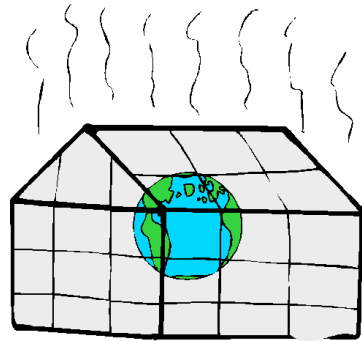


http://wings.avkids.com/Book/Myth/Images/ocean_sun.gif

Life of Our Sun



- Over the next billions of years, our Sun will continue to increase in luminosity.
- So in 1+ billion years, our Sun will be 10% more luminous.
- This will cause a “moist” greenhouse effect adding 10 degrees F to the average temp.



http://www.solcomhouse.com/Greenhouse_Effect.gif

Life of Our Sun



- This increase in total energy will have a major impact on the Earth!
 - Ice caps melt
 - Coastal regions flood
 - Equator becomes uninhabitable
 - Antarctica becomes warm



<http://changeyourways.wordpress.com/2009/06/12/what-on-earth/>

Life of Our Sun



- Increased temperature means that the lighter elements, like water molecules in the air, will have enough speed to escape Earth completely.
- The water of Earth begins to pack up and leave!
- In 1.1 billion years, the continents will be deserts and the oceans are beginning to evaporate.



<http://www.esquire.com/cm/esquire/images/Gd/desert-1108-lg.jpg>

Life of Our Sun



- As the Sun, uses up the hydrogen in the core, the Sun increases by 40% in brightness in 3.5 billion years.
- By that time, all of the oceans are gone!
- The baking sediments at the bottom of the oceans, release CO₂
- Earth will become Venus-like!
- Then the heat makes even those heavier molecules leave the Earth.
- The Earth will be a barren rock in about 4 billion years!



http://wings.avkids.com/Book/Myth/Images/ocean_sun.gif

Mitigation



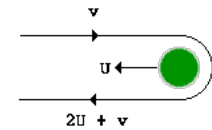
1. Move the population
 - I hear that Mars could be a nice place to live.
 - Need to terraform Mars, which could take a while.



Mitigation



2. Move the Earth
 - There is no place like home, so move it to a nicer place, farther away from the Sun.
 - Use gravity assist or the sling shot technique.

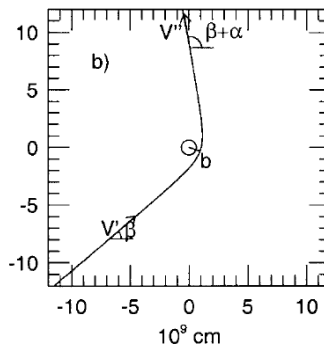


http://upload.wikimedia.org/wikipedia/commons/8/8e/Grav_slingshot_simple_2.gif

Mitigation



2. Move the Earth
 - Asteroids to the rescue?
 - Move many large asteroids in front of the Earth, sends them toward the Sun and the Earth outwards.
 - Need to do this every 6000 years to make Earth survive until the Sun hits the Red Giant phase.

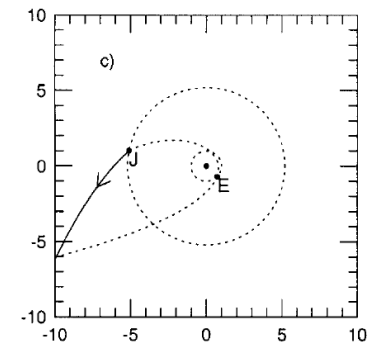


Korycansky et al. 2001

Mitigation



2. Move the Earth
 - For billions of years!
 - We don't have enough large asteroids.
 - We'll have to recycle.
 - The idea is to transfer energy from Jupiter's orbit to Earth's orbit.

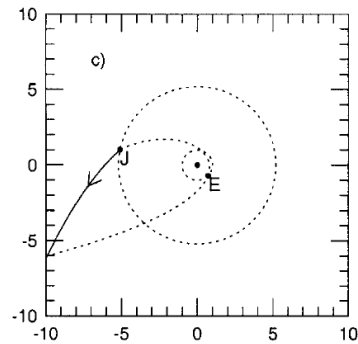


Korycansky et al. 2001

Mitigation



2. Move the Earth
 - Could keep us safe for a good 6 billion years!



Korycansky et al. 2001

How much Gas do we have left?



- Even if we could save the Earth during this time of increased brightness, eventually the Sun runs out of fuel.
- Total energy available is easily calculated by mass of hydrogen in Sun and energy released by each hydrogen conversion.
- We only have about 6 billion years left!

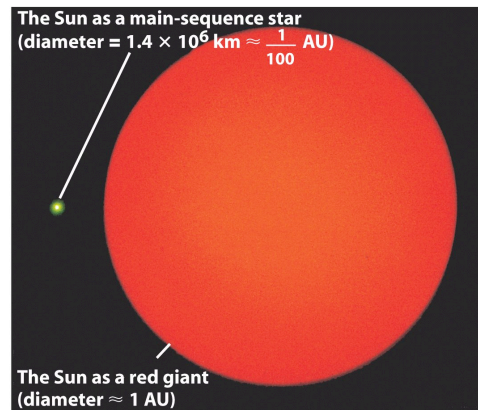


<http://skeptically.org/sitebuildercontent/sitebuilderpictures/pond/sun-econ-gas-pump.jpg.w300h294.jpg>

In 6-7 Billion years



- The Sun will expand to 100-250 times bigger than it is now!
- The same mass but now it's bigger.



The Sun today and as a red giant

In 6-7 Billion years



- The surface gravity decreases and the Sun has more luminosity.
- The solar wind turns into a stellar wind, and it loses material as it expands, about 10^7 times more than now.
- It's blowing it all away!

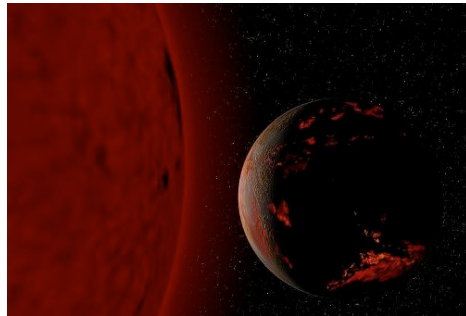


http://www.astropix.com/wp/wp-content/uploads/2006/12/2006_02.JPG

In 6-7 Billion years



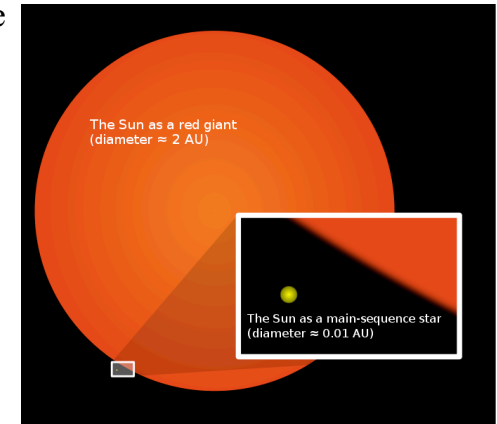
- During the time it expands the Sun loses a significant fraction of mass.
- So, the planets move outward.
- Planets race away as the Sun expands.
- Who wins?
- We aren't yet sure.



In 6-7 Billion years



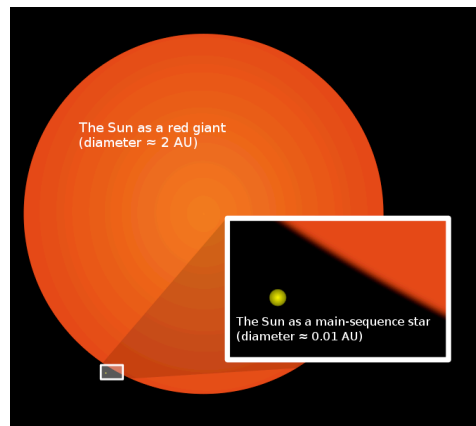
- We use to think that the Sun would gobble the Earth.
 - Mercury gone
 - Venus probably gone
 - Earth?
- BUT even if not, with the Earth's oceans and atmosphere gone, crust still melts.
- Not good...



In 6-7 Billion years



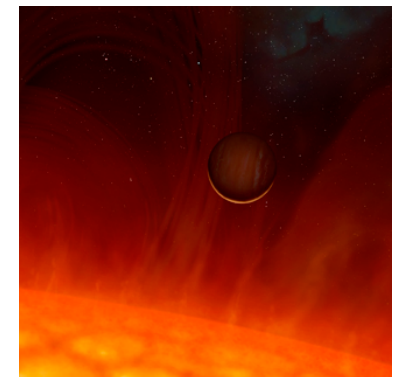
- Mars?
 - For sure too hot.
- Jupiter's Moons?
 - Still too hot
 - Europa's water vaporizes
- Even the moons of Uranus and Neptune may be too hot.



Mitigation



- We would have to move the Earth out to Pluto or further!
- Probably not possible.
 - Interactions with Jupiter may eject us from Solar System
- Even then, Sun no longer in equilibrium, may oscillate in size or brightness.
- BUT, we got billions of years to figure it out!



Natural Catastrophes



3. Stellar Evolution

- Advanced civilization can likely find solutions.
- Eventually, we would have to leave the Earth, move the Earth, or move to Mars.



http://www.boulder.swri.edu/~terrell/dtart_old.htm

Question



In 5 billion years, our Sun will begin to turn into a red giant, on its way to a white dwarf. But never fear,

- a) An advanced civ can stop the Sun from evolving.
- b) We can always move to Venus.
- c) An asteroid will probably hit and destroy the Sun first.
- d) The Moon will be fine.
- e) The Earth's oceans will evaporate before then.

4. Natural Catastrophes



4. Killer Supernovae!

- Death of a nearby massive star would be bad news.
- Explosion within 30 ly would destroy ozone layer.
- Right now, no candidates.
- Unlikely to happen in time scales of less than 2 billion years.
- A supernova event ~2 Myrs ago may account for an extinction event.



Question



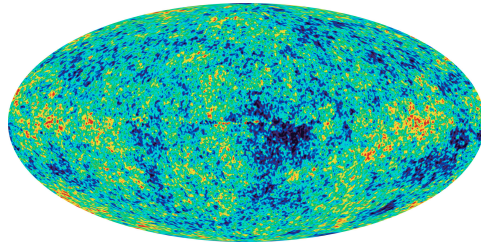
A nearby killer supernova

- a) is the most scary, as we won't know about it.
- b) would have to be very close, but it could destroy the ozone layer .
- c) is a supernova in our Galaxy.
- d) would not cause any real damage, no matter how close it was.
- e) will evaporate the oceans.

4. Natural Catastrophes



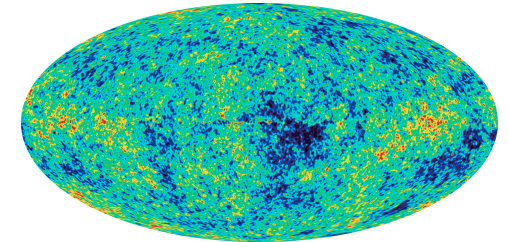
5. Ultimate limit to L!
 - Fate of the Universe.



4. Natural Catastrophes



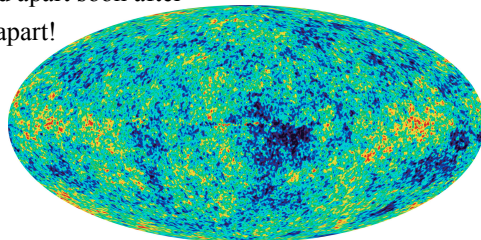
5. Ultimate limit to L!
 - A Big Crunch: 10^{12} years (a trillion years)
 - But, WMAP results from the cosmic microwave background suggest that we are in a flat universe.
 - Which does include dark energy



4. Natural Catastrophes



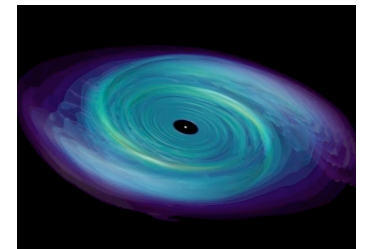
5. Ultimate limit to L!
 - The Big Rip?
 - If repulsive force increases– Brooklyn may expand too.
 - Gravity/E&M forces can not hold Galaxies rip apart
 - Could rip the MilkyWay apart in ~1 billion years
 - Earth gets ripped apart soon after
 - You get ripped apart!



4. Natural Catastrophes



5. Ultimate limit to L!
 - Big Rip seems unlikely
 - We'll know soon.
 - If we are just in a flat Universe, then it is a matter of energy.

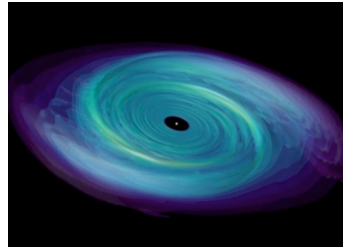


<http://www.youtube.com/watch?v=oGVYg0ce1Ps>

4. Natural Catastrophes



5. Ultimate limit to L!
- The end of the Universe... the death of the dark.



http://homepages.wmich.edu/~korista/web-images/accretion_ncstate.jpg

Stelliferous Age: 10^8 to 10^{15} years



- Last stars to form will happen in a few hundred billion years.
- Stars age and die
- In about trillion years all Sun-like stars are gone from the Universe forever.
- Only stars left are low-mass red dwarfs (~0.1 solar masses), which can live for trillions of years
 - Lots of these stars and they get brighter with age, so Galaxy brightness doesn't change too much



Stelliferous Age: 10^8 to 10^{15} years



- In 7-8 trillion years, in our Galaxy (Milkomeda), the last red dwarf stops fusing, becoming a white dwarf.
- These tiny white dwarfs will stay hot for quite some time.
- Wait another few trillion years and they fade.
- So when the Universe is 100 trillion years old, the Universe goes dark.



Really Dark



- If the Universe keeps expanding, it get worse for astronomers.
- The Galaxies we can see now, far away galaxies move out of our view.. Too far to see given the age of the Universe... out of our horizon.
 - The observable Universe is less and less
- The one giant elliptical galaxy (all that is left from our local group) is all that can be seen.
- The Universe appears empty!



Humans?



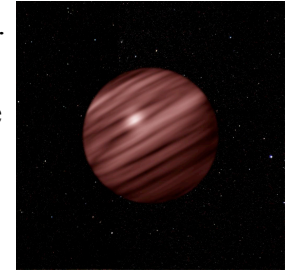
- We have 100 trillion years!
- Maybe longer, by smashing stars together to make fusion last longer.
- Won't last too long.
- When the Universe is slightly older than 100 trillion years old, the human race is out of fuel, out of stars, and out of luck.
- But the Universe isn't done!



The Degenerate Era: 10^{15} to 10^{40} years



- Stellar corpses are all around the Galaxy.
- Every once in a while, a black hole will accrete a compact object, creating light again.
- Corpses may collide (remember we are talking 100 trillion years of time not the measly 13.7 billion of the Universe so far), and create new stars.
- Brown dwarfs, which did not have enough mass to fuse, can collide, making new stars.
- New life? Different Universe..



The Degenerate Era: 10^{15} to 10^{40} years



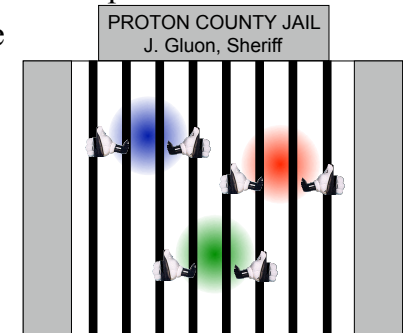
- But after a trillions, then quadrillions, and then quintillions of years, everything that can ever burn has happened.
- The Galaxy starts to lose weight.
 - Interactions with the stellar corpses, cause all the low-mass objects to be ejected from Galaxy.
 - High-mass objects fall to the center.
 - Supermassive Black Hole feeds!
- If the Earth still orbited the dead Sun (white dwarf) it is likely kicked out of the Sun and the Galaxy– a frozen dead planet in intergalactic space.



Proton Decay



- Remember when quarks were imprisoned?
- We think that protons are radioactive.
- Except that they decay with a half-life of about 10^{31} years.
- Time is all that is left.

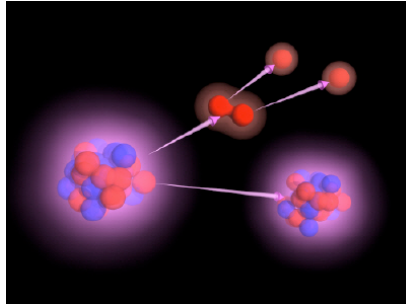


10^{31} years to life
Little chance of parole

Proton Decay



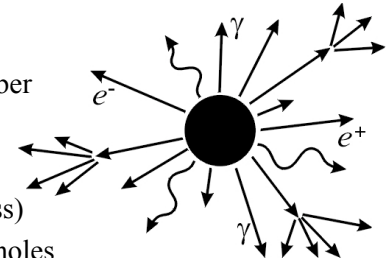
- This proton decay creates heat again, feeble heat.
- What does non-proton life do?
- White dwarfs will evaporate
 - At -454 F, they are the hottest thing around!



The Black Hole Era: 10^{40} - 10^{92} years



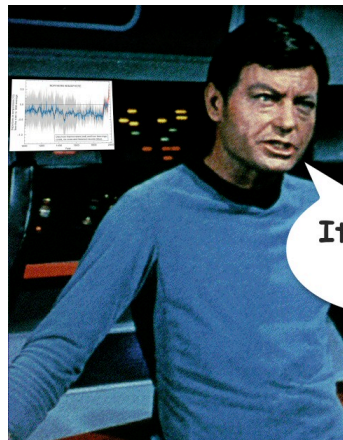
- Black Holes survive.
 - Not made from matter, remember
- Galaxy is
 - The Supermassive Black Hole (1-10% of original Galaxy mass)
 - Trillions of stellar mass black holes
 - Lower mass stuff that was thrown out, so very far away.
- Hawking radiation is slow, but it will begin to evaporate the black holes
 - Slow, but lots and lots of time



The Dark Era: 10^{92} - Infinity



- 10^{92} is crazy!
- I mean really, really crazy!
- The weight of a single proton to the rest of the Universe is only 10^{79} !
- Still, at this point, the Universe is dead!
- Dead Jim!



It's dead
Jim

The Dark Era: 10^{92} - Infinity



- Beyond this, two particles will once in a great while interact, but nothing will really happen.
- Universe is dead, randomized, and silent.
- Nothing really will ever happen again..
- Or will it?



The Dark Era: 10⁹²- Infinity



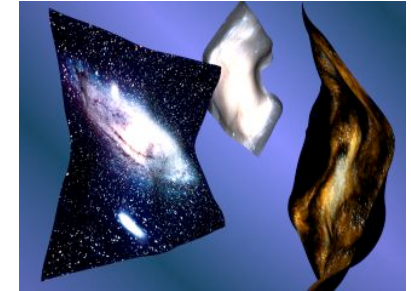
- Rebirth?
- We don't know what caused the Big Bang.
- Maybe it happens again?
- Maybe it already has?



Branes, Branes!



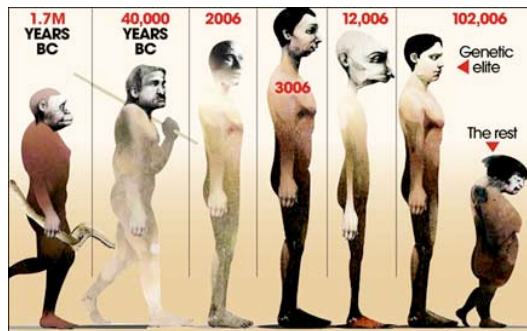
- One idea is that the Universe has 11 dimensions
 - Our 4 dimensional Universe floats around in this space
 - Other universes float there too (called branes, short for membranes)
 - Sometimes they collide
 - Violently disturbed, energy/matter heat up, expanding space
 - Sounds familiar..



Mitigation



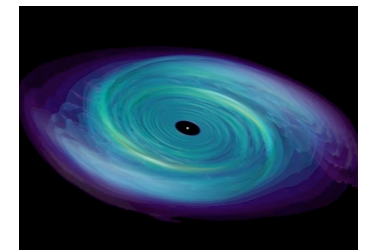
- Are you kidding me?
- If humans live this long, they won't be anything we'd recognize as human.



4. Natural Catastrophes



5. Ultimate limit to L!
 - Bottom line is that the maximum age is speculative.



Top 10 Ways Astronomy Can Kill Humans



1. Impacts
2. Sun Evolution/Coronal Mass Ejections
3. Supernovae
4. Gamma-Ray Bursts
5. Rogue Black holes
6. Rogue White Dwarfs
7. Galaxy Collisions
8. Cosmology
9. Quasars
10. Aliens

L-ing it



- We are talking about the amount of time that an advanced civilization (averaged over time) can communicate.
 - They may not want to for long periods of time
 - They may give up
 - They may be killed off
 - They may run out of resources
- Solving our energy problem (cheap energy) will give the largest lifetimes.

What is L?



- How long on **average** can an advanced civilization exist?
- Again, we only have a sample of 1 from which to discuss. What is our civilization's lifetime?
 - Short Term (100-1000 yrs)
 - Give up on communication due to budgets.
 - Depletion of resources.
 - Population.
 - War.
 - Long Term (10^5 to 5×10^9 yrs– age of galaxy is 10^{10} yrs and we took half of that to evolve)
 - Stellar Evolution.
 - Don't forget the random volcano, asteroid, or supernova.
 - Still in many cases an advanced civilization may be prepared for many of the issues!