

# Astronomy 330



This class (Lecture 10):

Exoplanets 2

*Braden Anderson*

*Jennifer Bora*

Next Class:

Origin of the Moon

*Eric Gobst*

*Suharsh Sivakumar*

**HW 4 is due Thursday!**

Music: *3<sup>rd</sup> Planet* – Modest Mouse

# HW 2



- Ryan Ruddell  
<http://www.cropcirclesearch.com/index2.html>
- Neel Lawande  
<http://www.paranormalhaze.com/5-pieces-of-evidence-that-suggests-intelligent-alien-life-exists/>

# Presentations



- Braden Anderson  
[Roswell](#)
- Jennifer Bora  
[Alien Language](#)

# Outline



- Exoplanets – they are all over the place.
- Radial Velocity studies (wobble method)
- Transit Studies (Kepler)
- Estimate  $f_p$ ?

# Finding Planets



1. Transit Method:      Occultation
2. Radial Velocity:    Stars will wobble
3. Direct Detection:      Direct imaging
4. Astrometry:            See the stars move

**Only a few planets have been detected directly in the optical and IR. Remember that planets in our Solar System seem bright because they reflect light from the Sun in the visible.**

# We have imaged some Exoplanets



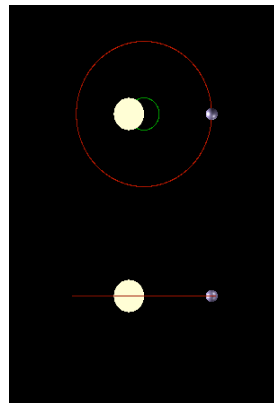
Okay, so imaging is possible (I couldn't say that a few years ago), but difficult-- only working on the planets that are big and stars that are not too bright. Is there a better way to find planets?

Yes, the wobble or radial velocity method and now the transit method.

# Wobbling Stars



- Planet's gravity "tugs" the parent star
- As planet orbits the star, the tugs make the star "wobble"



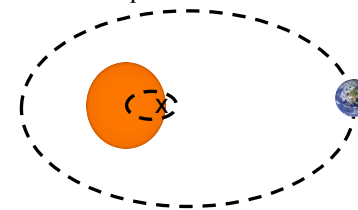
Star and planet each orbit around their mutual center of mass

# Star Wobble: Radial Velocity

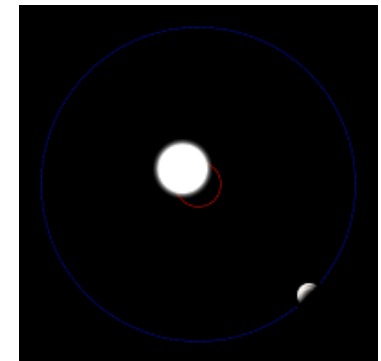


Newton's 3<sup>rd</sup> Law:

- Both planet and star move
- Both orbits fixed around the "center of gravity"
- Star's period? Place your bets...
  - Same as planet



Greatly exaggerated

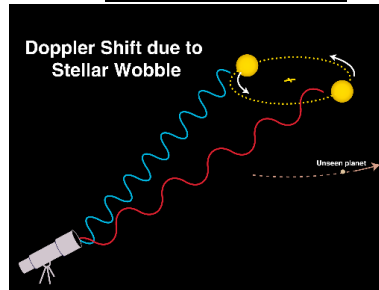
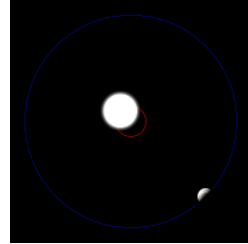


[http://en.wikipedia.org/wiki/File:Planet\\_reflex\\_200.gif](http://en.wikipedia.org/wiki/File:Planet_reflex_200.gif)

## Star Wobble: Radial Velocity



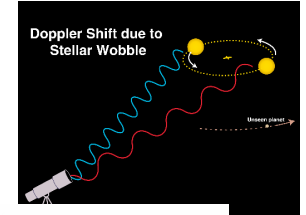
- Star movement too small to see
  - Moves in small, tight circle
  - But “wobble” in star speed detected!
  - The stellar spectrum is shifted red and blue as it moves towards us and away from us.



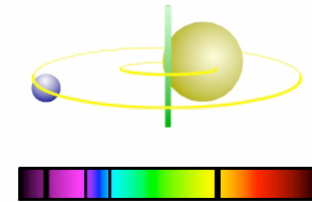
## Star Wobble: Radial Velocity



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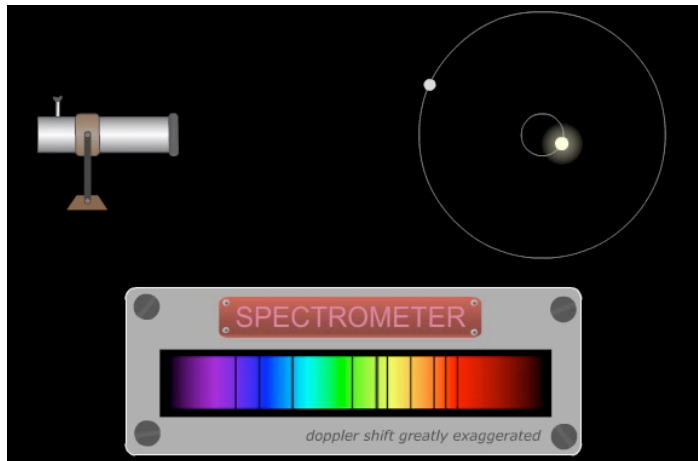
How Planet Hunting Works



## Star Wobble: Radial Velocity



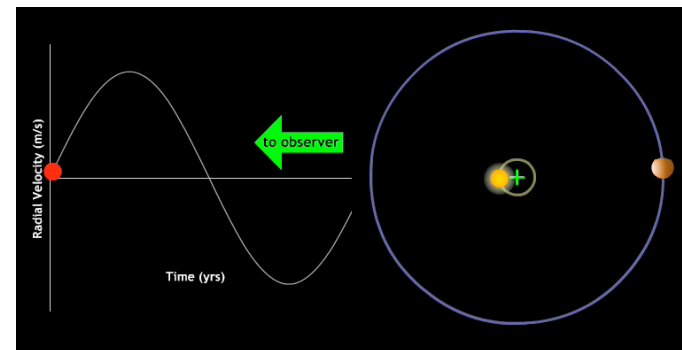
- Need very sensitive spectrometers to measure the wobble.



## Star Wobble: Radial Velocity



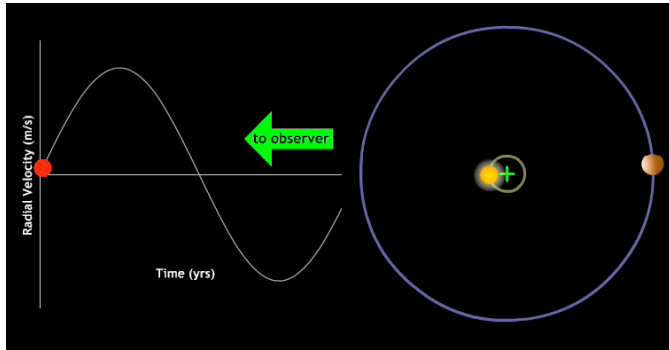
- Observed wave form as the planet orbits the star.
  - Period depends on what?
  - Amplitude depends on what?



## Star Wobble: Radial Velocity



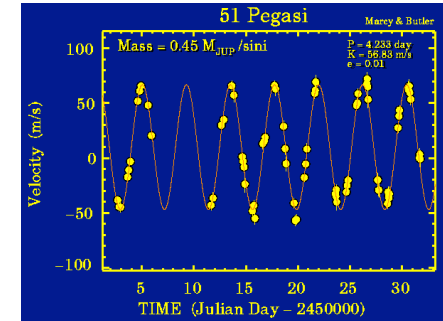
- Observed wave form as the planet orbits the star.
  - Period depends on what? Orbital distance from star.
  - Amplitude depends on what? Mass of planet.



## First extrasolar planet around a Sun-like star



- Discovered in 1995 orbiting 51 Pegasi
  - Doppler shifts reveal a planet with 4.23 day orbital period
  - $0.5 M_{Jup}$  at 0.05 AU from its star!
- As of Feb 2012, 760 exoplanets are listed in the Extrasolar Planets Encyclopedia.



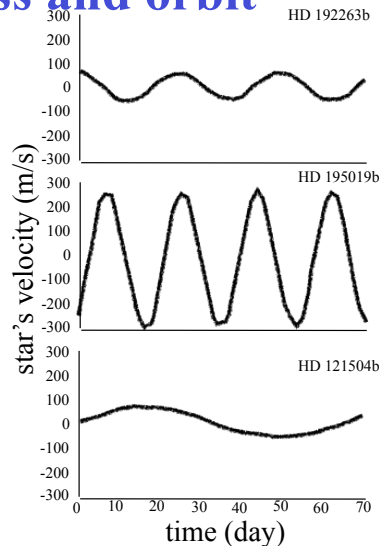
51 Pegasi b orbits its star every 4.23 days!

<http://exoplanet.eu/catalog.php>

## Doppler shift tells us about a planet's mass and orbit



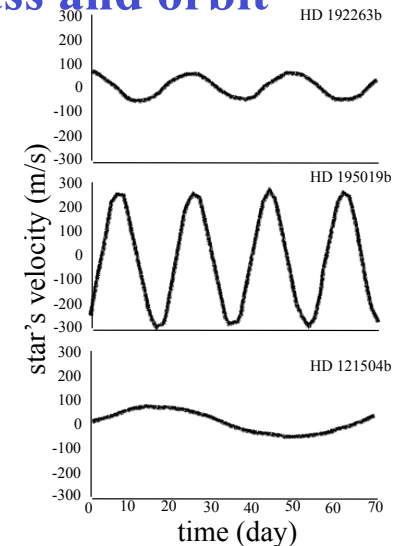
- Top graph shows a  $0.5 M_{Jup}$  planet in an orbit around a Sun-like star at 0.05 AU
- Middle graph - more massive planet ( $2 M_{Jup}$ ) in a 0.05 AU orbit
- Bottom graph -  $0.5 M_{Jup}$  planet in more distant orbit



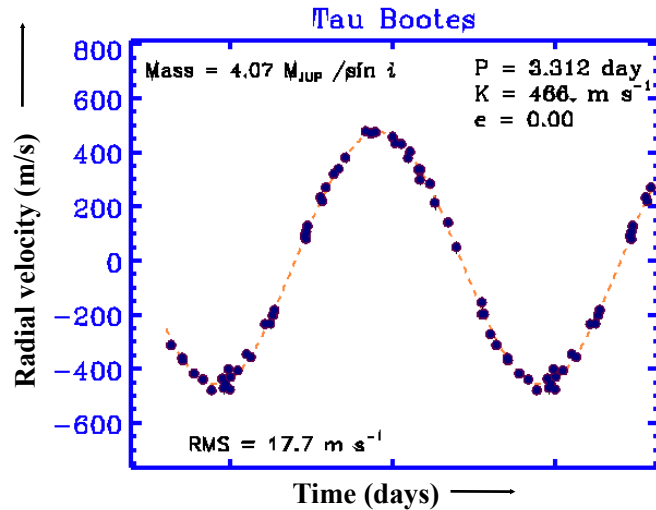
## Doppler shift tells us about a planet's mass and orbit



- The amplitude of the wobble tells us the planet's mass
- More massive planet in same orbit produces greater Doppler shift with same period around the star
- The period of the wobble tells us the radius of its orbit
- Same mass planet in a more distant orbit produces a smaller Doppler shift with a longer period.



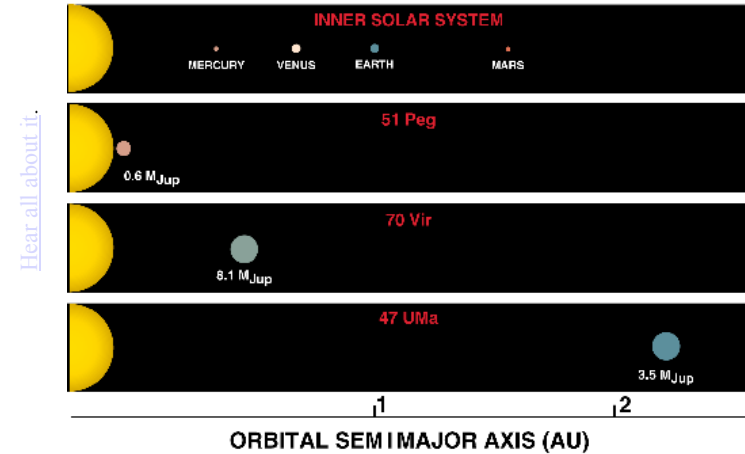
# Radial Velocity Shifts: Planets around other Stars?



# Early Discovery-- 1996



## PLANETS AROUND NORMAL STARS



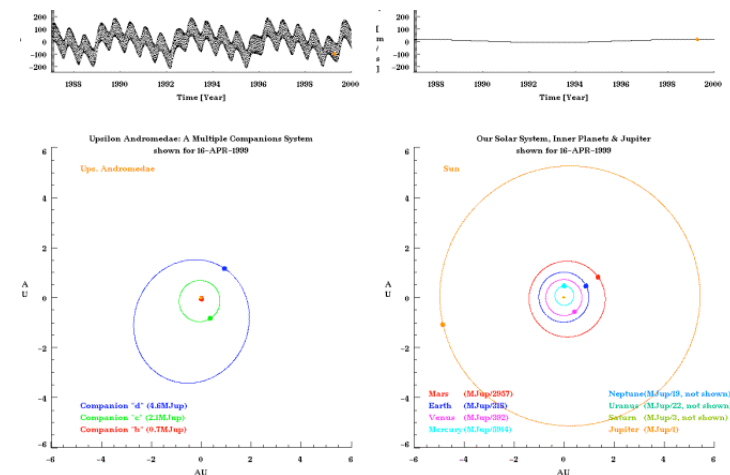
Hear all about it.

# Selection Effect



- Big planets make big wobbles
  - Bigger Doppler shift
  - Easier to detect
- Close planets make fast wobbles
  - Shorter period of Doppler shift
  - Less time to observe a full orbital period
- Think about it...
  - Jupiter's orbital period ~12 years
  - Saturn's orbital period ~30 years
  - Detecting extrasolar planets for 15 years
  - Enough time for 1 Jupiter orbit, but not Saturn

# Detecting Our Wobble

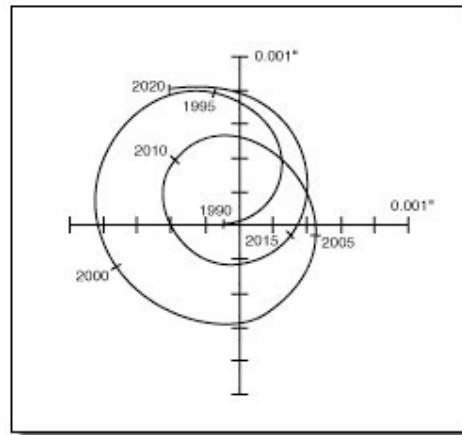


## The Sun's Wobble



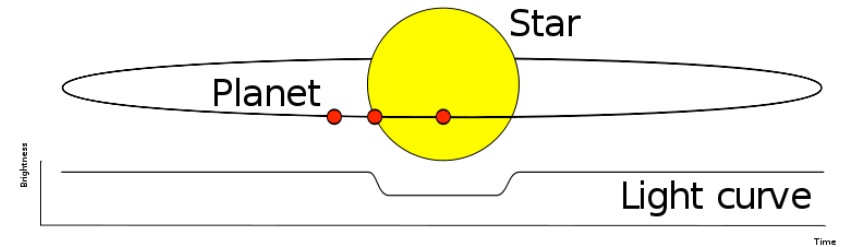
*Astrometric displacement of the Sun due to Jupiter (and other planets) as it would be observed from 10 parsecs, or about 33 light-years.*

*If we could observe this, we could derive the planetary systems— also called astrometry.*



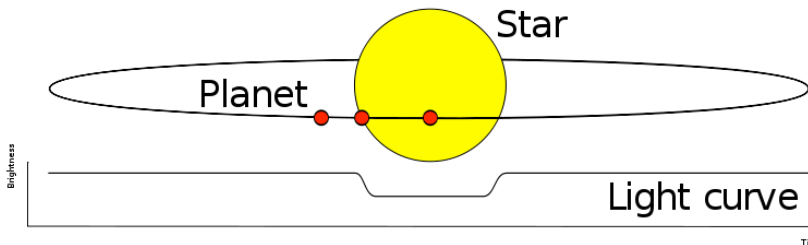
[http://planetquest.jpl.nasa.gov/Keck/astro\\_tech.html](http://planetquest.jpl.nasa.gov/Keck/astro_tech.html)

## Detecting Planets via Transits



The effect on the observed brightness when a planet passes in front of the parent star

## Detecting Planets via Transits

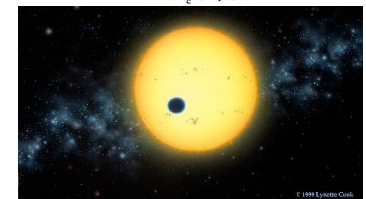
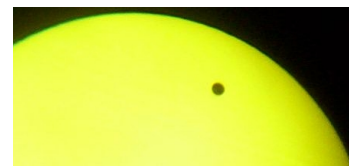
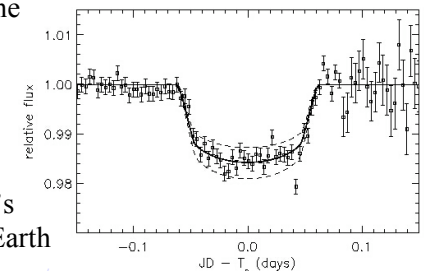


Look for regular dips in the brightness of a star as a large planet passes, or transits, in front of it. Requires the extrasolar planet's orbital plane to be pointed right at Earth.

## Transits



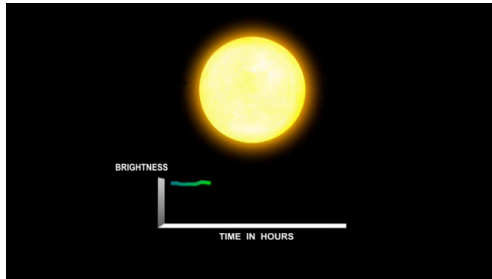
- The planet passes in front of the star— like Venus 2004.
- Can find planet radius
- Best chance of finding Earth-like planets
- Requires the extrasolar planet's orbital plane to be pointed at Earth
- <http://science.howstuffworks.com/planet-hunting1.htm>



## Transits are the best chance to find Earth-like planets



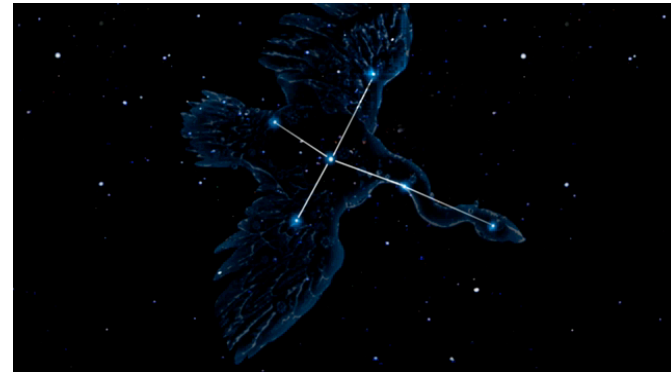
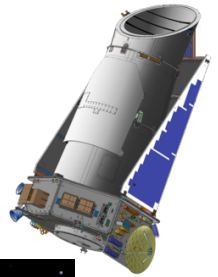
- Fraction of starlight blocked tells us planet's size
- Time between transits gives us orbit period
- A Jupiter-sized planet transiting a Sun-like star would cause a 1% brightness drop
- Best method to find Earth-like planets



Earth-sized planet would cause a 0.01% drop in brightness (1 in 10,000), but best instruments can measure 1 in 50,000 drop in brightness!

## Kepler Mission

- Launched March 7, 2009
- Probing planet transits toward 145,000 main sequence stars (10 square degs)



## Kepler Mission

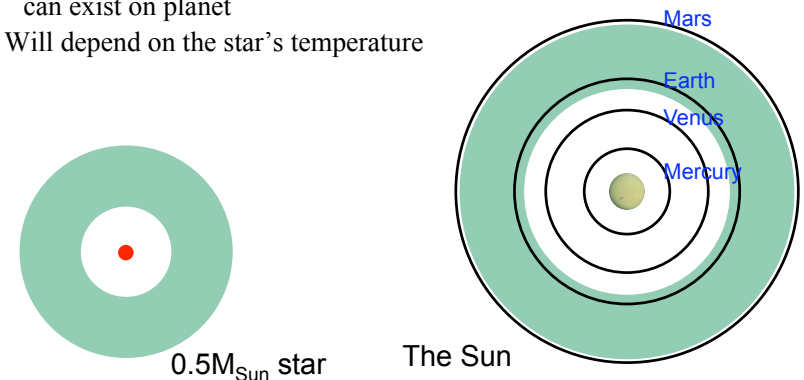


- Looking for small intensity dips in stars
- Must have planet block some of the star's light
- Playing probability game as random orbits must intersect our line of sight
  - For Earth, the chance of this happening is 0.465%.
  - If ALL stars have Earths, would see 678 Earths
- But, Kepler group won't call objects candidates until they see the dip three times.. not yet enough time for Earth's at 1 AU.

## Planets in the Habitable Zone: Would be nice for life



Distance from the star such that liquid water can exist on planet  
Will depend on the star's temperature

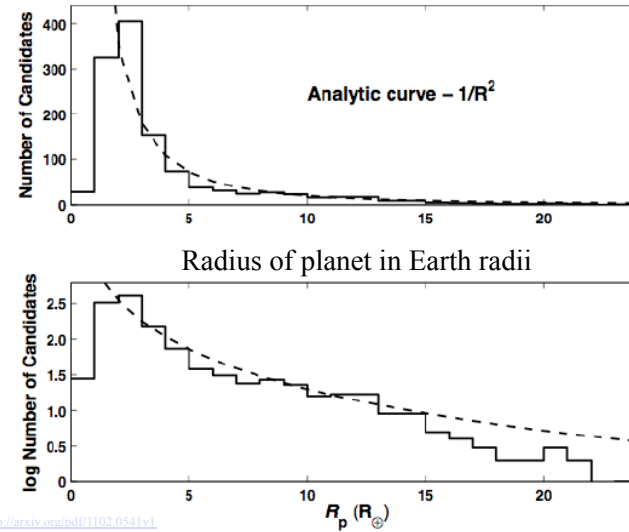
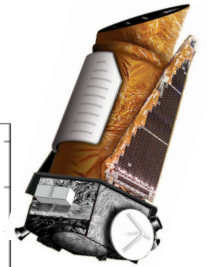


# Kepler Status



- First major release was early 2011 with another major release expected soon
- 2326 planet candidates
  - 207 Earth-sized
  - 48 planets in the habitable zone
  - Habitable zone is the region around a star where water is likely to be a liquid.
- Candidates because 90% confidence currently. Will improve with time.

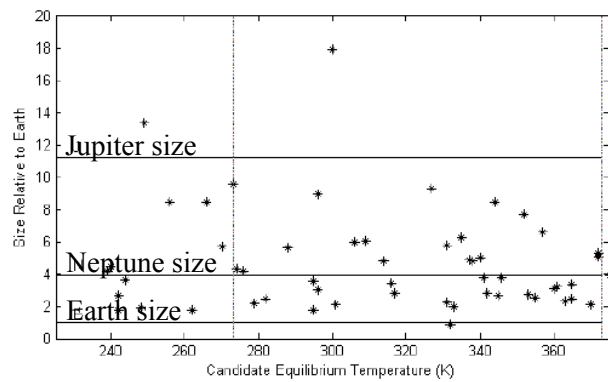
# Kepler Status



Unlike radial velocity sources, smaller planets are more common in Kepler data.

<http://arxiv.org/pdf/1102.0541v1>

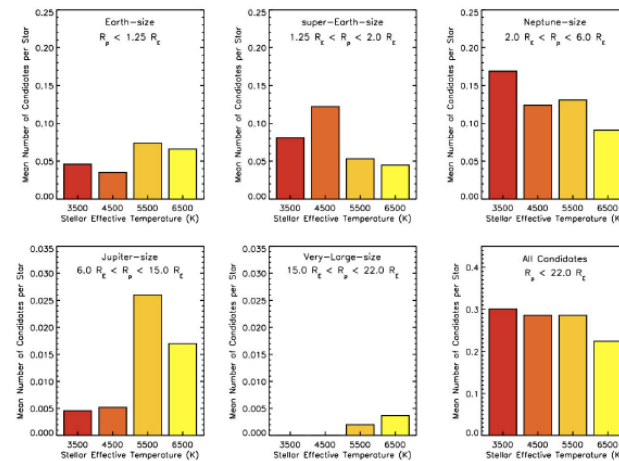
# Kepler Status



Many candidate sources in the habitable zone

<http://arxiv.org/pdf/1102.0541v1>

# Kepler Status



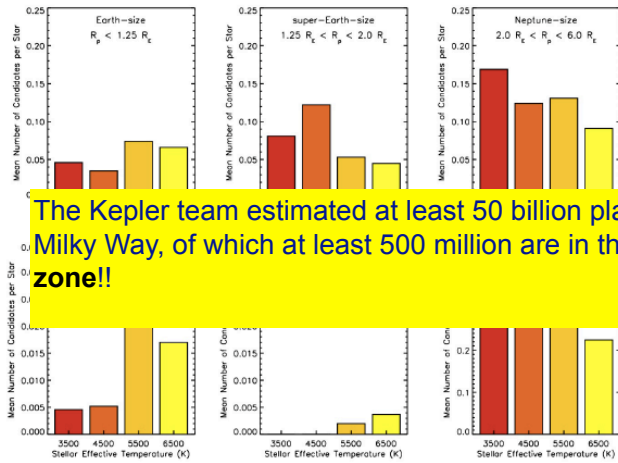
First scientific estimate of  $f_p$  ever! = 34%

Still a lower limit though.

<http://arxiv.org/pdf/1102.0541v1>



# Kepler Status



The Kepler team estimated at least 50 billion planets in the Milky Way, of which at least 500 million are in the **habitable zone!!**