Astronomy 330

This class (Lecture 14):

<u>Next Class:</u>

Life on Earth Alesia Prakapenka Anthony Salis Origin of Life *MidTerm Due!* Sonja Bromann Michelle Boehm

Music: Bring Me to Life-Evanescence

HW 2

- Carolyn Buesing http://www.ufos-aliens.co.uk/cosmicabduct.htm
- Maura Walsh
 http://aliensandchildren.org



- Emailed to everyone after class last Thursday.
 - 50%: 4 short (few paragraphs) essays
 - 50%: 1 large (~1 page) essay (with definition terms)
- Must be typed, not handwritten.
- Will cover material up to and including last Thursday.
- It is a closed notes exam (honor system!).
- You can make 1 page of notes that you use during the exam.

Presentations

- Alesia Prakapenka
 <u>Aliens in Religion</u>
- Anthony Salis <u>Alien Math</u>

Outline

- What type of stars are good for life?
- Monomers and polymers
- Proteins and nucleic acid?
- Where did the monomers of life come from?

 $n_e = n_p \times f_s$



Class number is 1.3



http://nike.cecs.csulb.edu/~kjlivio/Wallpapers/Planets%2001.jpg

 $n_e = n_p \times f_s$

- n_p: number of planets suitable for life per planetary system
- f_s : fraction of stars whose properties are suitable for life to develop on one of its planets

We can list 5 situations that will have an effect on f_s .



http://nike.cecs.csulb.edu/~kjlivio/Wallpapers/Planets%2001.jpg

Differences of Stars to Life

1. <u>Metal rich stars</u>. Stars with heavy elements, probably more likely to have planets. Suggested in the current planet searches. About 90% of all stars have metals.



Differences of Stars to Life

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- 2. <u>Main sequence stars</u>. Need the brightness to stay as constant as possible. Otherwise the temperature changes dramatically on the planets. This is 99% of all stars.



Differences of Stars to Life

3. <u>Length of time on</u> <u>the main sequence</u>.

We needed temperature stability for 5 billion years to get intelligence on Earth. This rules out stars more massive than 1.25 solar masses! 90% of all stars are less massive than that.



http://mjbs.org/hr.jpg

Differences of Stars to Life

4. Minimum mass of star

For low-mass stars, any life bearing planetd would have to be closer to the star- and closer to stellar effects (e.g. tidal locking and more flares from low mass stars). That limits us to a minimum of 0.5 solar masses. 25% of all stars are more massive than that.







5. <u>Binarity</u>. Planets may form. But they may have odd orbits unless the 2 stars are far enough apart or the planet orbits the pair. Only 30% of all stars are single stars. 50% of all stars are single stars or wide binary stars.





http://spaceflightnow.com/news/n0210/11planet/

Adding it all up

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Stellar Requirement	Mass Limit	Fraction OK	Cumulative Fraction
Heavy Elements		0.9	0.9
Main Sequence		0.99	0.891
Main Sequence Lifetime	M < 1.25 M _{sun}	0.90	
Synchronous Rotation/ Flares	$M > 0.5 \ M_{Sun}$	0.25	
Not a Binary		0.30	0.267
Wide Binary Separation		0.50	

Adding it all up

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	Stellar Requirement	Mass Limit	Fraction OK	Cumulative Fraction
	Heavy Elements		0.9	
	Main Sequence		0.99	
<	Main Sequence Lifetime	M < 1.25 M _{sun}	0.90	
<	Synchronous Rotation/ Flares	M > 0.5 M _{Sun}	0.25	
	Not a Binary		0.30	
•	Wide Binary Separation		0.50	

Adding it all up

	Stellar Requirement	Mass Limit	Fraction OK	Cumulative Fraction
✓	Heavy Elements		0.9	
✓	Main Sequence		0.99	
✓	Main Sequence Lifetime	M < 1.25 M _{sun}	0.90	
✓	Synchronous Rotation/ Flares	$M > 0.5 \ M_{Sun}$	0.25	
	Not a Binary		0.30	
	Wide Binary Separation		0.50	

f_s: fraction of stars that life can exist around

Stellar Requirement	Mass Limit	Fraction OK	Cumulative Fraction
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Wide Binary Separation		0.50	

Value can range from ~ 0.06 to ?

Drake Equation Frank That's 1.27 Life-liking systems/year Drake Dalpal probably located. = $R_* \times f_p \times n_e \times f_1 \times f_i \times f_c \times L$ Ν c) # of # of Star Fraction Fraction Earthlike Lifetime of advanced Fraction Fraction life. formation of stars that civilizations planets on which that evolve advanced rate with communlife arises intelligence civilizations we can per icate planets system contact in our Galaxy today vrs/ 15 0.65 1.3 x 0.1 life/ intel./ comm./ to be too close. = 0.13 life intel. comm. planet systems/ stars/ planets/ vr star system

So Far, We have Studied

- The Universe
 - Big Bang
 - Creation of hydrogen, helium...
 - Galaxy formation
 - · Swirls of elements embedded in self-gravitating cloud of dark matter
 - Star birth
 - Energy generation and element production in selfgravitating mass of gas
 - Planets
 - · Ice, rock, gas surrounding stars form planetesimals, then planets

Ouestion

The best type of life sustaining stars are

- a) Low mass stars (less than 0.5 solar masses), as life can exist nearer the star where more terrestrial planets are
- b) Binary stars, as they double the chances of life.
- Stars off the main sequence, as they have lived the longest, they are the best chance for finding intelligent
- d) Middle mass stars (less than 1.25 and more than 0.5 solar masses), as they live longer and don't require the planets
- e) Massive stars (more than 2 solar masses), as they have more mass from which to form planets.

Life on Earth

- Time to examine terrestrial evolution.
- Need to understand what is needed for life to arise.
- Again, some Earth chauvinism.
- Relies on chemical evolution
- Eventually life began?



http://www.accessexcellence.org/bioforum/bf02/awramik/bf02a1.html



Life on Earth

- In our scientific approach, we look at life as a result of chemical evolution of complexity.
- We will view the formation of "life" on planets as we did star formation
 - A natural consequence of natural laws
 - More specifically, as a consequence of the complex chemistry that is sometimes achieved.

Science only happens when you are not watching

http://www.toothpastefordinner.com/052802/science-only-happens.gif

Cosmic Imperative?



- But is life a cosmic imperative?
- Just like gas forms galaxies, and in galaxies stars and planets form, do chemicals on some planets form molecules that lead to life?