

Top Ten Signs Your Astronomy Professor May Be Nuts

(based on Lee Carkner List)



- 10) The title of every lecture is: "Man, Them Stars is Hot!".
- 9) His so-called "telescopes" are really just paper towel rolls covered in aluminum foil.
- 8) To illustrate the vastness of the universe, he makes everybody walk to Springfield.
- 7) Thinks he's married to the slide remote.
- 6) Your grade is based entirely on how many ping-pong balls you can fit in your mouth.
- 5) His so-called Drake Equation video is really just an old episode of Alf.
- 4) He makes everyone wear a soup pot on their head to protect the class from "Klingon mind control lasers".
- 3) About 90% of all classes involve dressing monkeys up to look like Frank Drake.
- 2) When you go to his office hours, he is always hiding under the desk so that the "space squirrels" can't get him.
- 1) The only observing advice he ever gives is, "Keep an eye out for the mothership."

Outline



- The First Stars
- The probable fate of the Universe
- Everything depends on Dark Energy

Astronomy 330



This class (Lecture 5):

The End of the Universe

Next Class:

Molecular Clouds

**Presentation Synopsis
due Tuesday!
(grace period until Feb 3rd)**

Music: *Across The Universe* – The Beatles

Presentations



- The presentation schedule has been decided by random selection.
- It is posted in the [schedule](#) section of the webpage.
- Make sure to check those dates ASAP.

Presentations



- Will be treated like a real talk.
- I will keep you to 12 minutes with 5 minutes of questions. (Peers should deduct points if too short, or I will.)
- Any speculative claims *MUST* have a scientific reference source.
 - Can't just claim that monkeys live on the Moon.



Presentations



- Can give presentation in any format you want.
- Over last few semesters:
 - 97.9% powerpoint
 - 1% talking with pics from webpages
 - 1% dedicated webpage
 - 0.1% overhead slides
- If presentation is electronic, I want to see it 1-2 days in advance
 - Email me
 - Or, on netfiles, email me URL location
 - Or, bring in burned CD (present to me class BEFORE)
 - Or USB Flash Drive (present to me class BEFORE)



Oral Presentation



1. How relevant is the general topic to this class (e.g. search for extraterrestrial life)?
2. How interesting is the topic for the general class audience?
3. Rate the extent of the speakers knowledge on the topic?
4. Rate the quality of the overall presentation?
5. Does the research have a solid scientific basis?

These questions are rated 1-10 out of 10 scale by your peers!

Common Mistakes

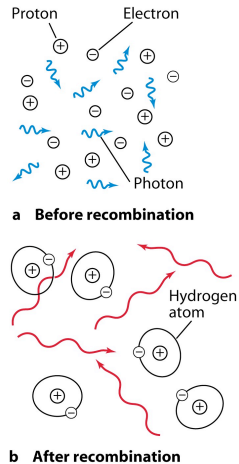


- Too much text on a slide.
- Too long (only 2% are too short).
- Background graphics or color makes text hard to read.
- Reading the slides is boring, use as points but not the whole message.
- 12 minutes is not as long as it sounds.

Era of Recombination



- In early Universe, photons were energetic, kept atoms ionized
 - protons and electrons couldn't make neutral hydrogen atoms
- After 380,000 years, photons couldn't ionize hydrogen anymore
 - Expansion of space stretched photons' wavelengths
 - Not enough energy to ionize hydrogen
 - Universe became transparent to photons
- This radiation is the source of the Cosmic Microwave Background!
- **The first H atoms proper!**



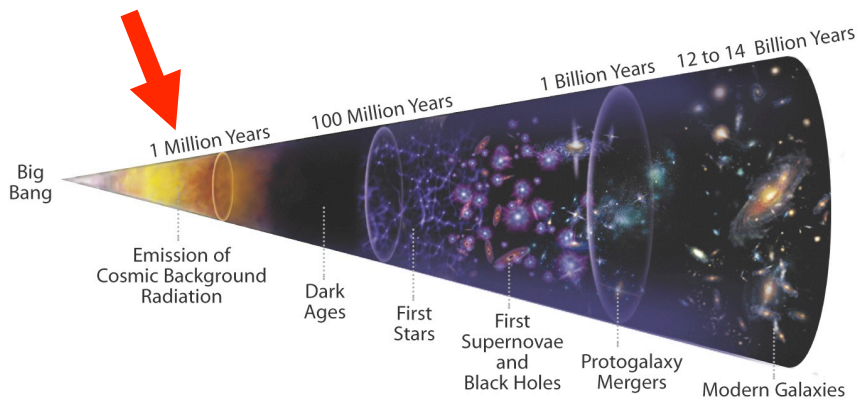
Question



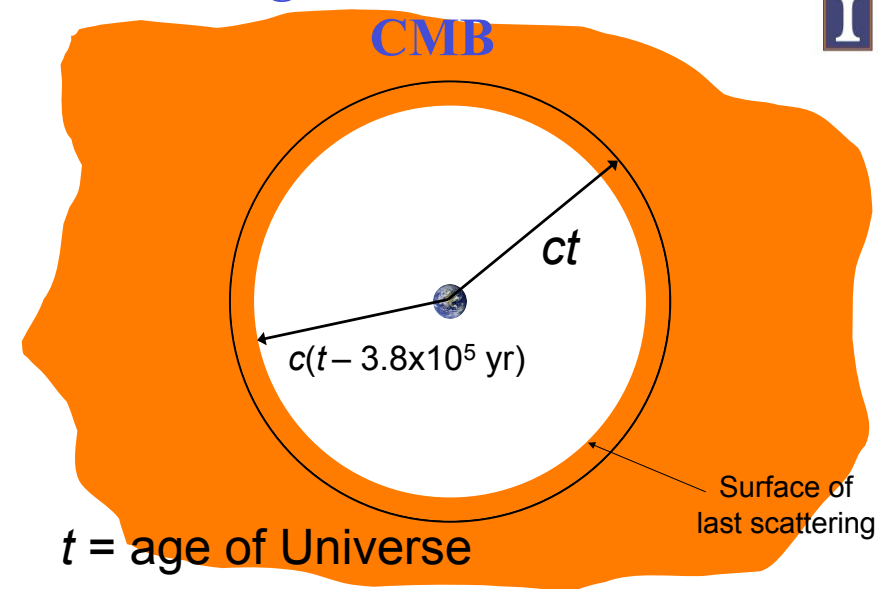
How did Hydrogen first appear in the Universe?

- When the Universe cooled and quarks combined to form the first protons, eventually gaining an electron.
- When the Universe cooled and the melted protons reformed, eventually gaining an electron.
- When the Universe cooled and the antimatter turned into matter, eventually gaining an electron.
- When the Universe cooled and the hydrogen atoms fused into helium atoms, eventually gaining an electron.
- They always existed.

Origin of the CMB



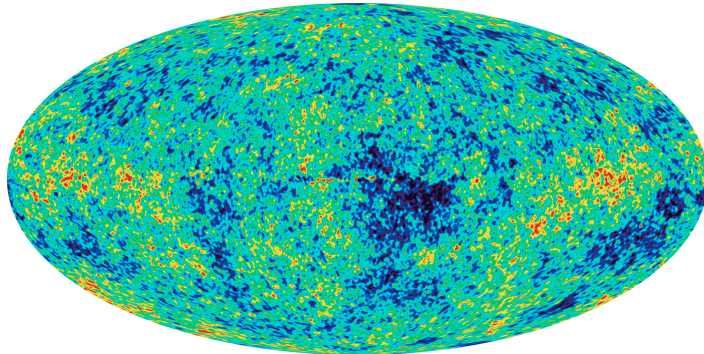
Looking Back in Time to the CMB



The Seeds of Galaxies



These small perturbations in temperature are the fluctuations (smaller than 1 in a 100,000) that caused the large scale structures we see today. This is what formed galaxies. All of this happened only 380,000 years after the Big Bang.



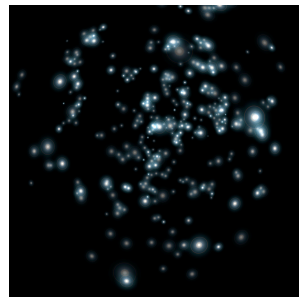
- After recombination came a period known as the Dark Ages
 - 380,000 to 200 million years
 - No light yet detected from this period
- Matter consists of warm clouds of hydrogen and helium
 - Too hot for star formation to occur
 - Gravity slowing drawing clouds together into bigger and bigger clumps
 - Proto-galaxies

<http://www.darkages.com/>

The First Stars



- From the initial seeds of the Big Bang, our local group of galaxies probably broke into clumps of hydrogen and helium.
- We'll look at star formation in detail later, but let's think of the first star to form in our proto-Milky Way
- May have formed as early as 200 million years after the Big Bang.
- Probably more massive than stars today, so lived quickly and died quickly.
- What happened? Why did this "raw" gas form anything?



<http://www.blackshoals.net/ImageBank/gallery/gallery/huge/The-first-stars-clustering.jpg>

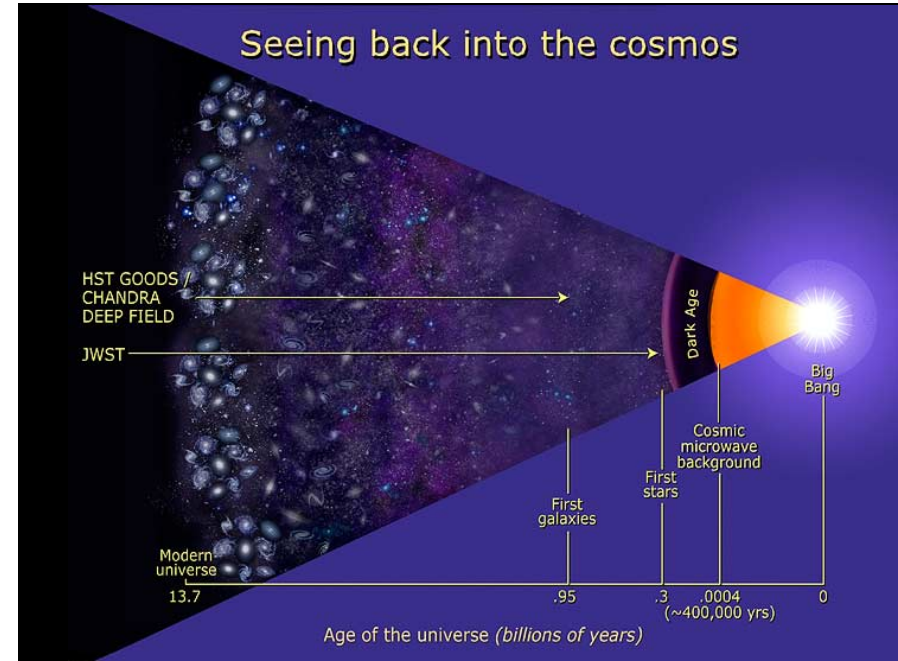
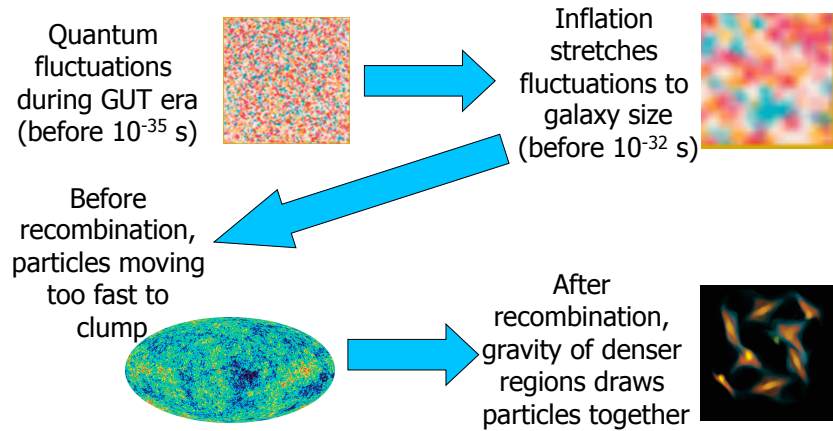
Question



A planet forms around one of the first stars in the Universe, which of the following is the most correct?

- It will be a rocky planet.
- It will be mostly made from hydrogen.
- The life that forms on this planet will be very alien.
- It will be a reddish-blue color.
- It will be made in the outer reaches of the Galaxy.

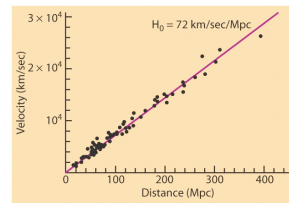
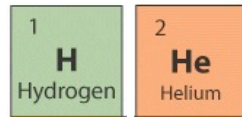
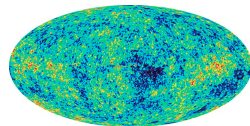
The Beginnings of Galaxies



From the Home Office in Urbana, IL Top 3 Reasons We Believe in the Big Bang



1. Cosmic Microwave Background
 - Big Bang working at about 380,000 yrs
 - Tiny fluctuations: "seeds" of galaxies
2. Big Bang Nucleosynthesis
 - H and (almost all) He come from the Big Bang
 - Big Bang working at 1 sec
3. The Hubble Law: $v=H_0d$
+ Einstein's General Relativity
= Expanding Universe with an age of 13.7 billion yrs



One of the most successful scientific theories of all time!

The Universe: Timeline



- Big Bang: 13.7 billion years ago
- GUT era: $+10^{-35}$ second, energy and quarks
- Inflation: 10^{-35} to 10^{-32} seconds, Universe expands by more than 10^{50} !
- Quark confinement: 10^{-32} to 10^{-6} seconds, protons and neutrons form
- Matter vs. antimatter: 10^{-6} seconds, matter wins
- Big Bang Nucleosynthesis: 10^{-4} seconds to 3 mins, He and some other nuclei form.
- Era of Recombination: 380,000 years. Universe becomes transparent, CMB
- Dark Ages: 380,000 to 200 million years, gravity works on stuff
- Stars: 200 million years, first stars form, protogalaxies



Fire and Ice



What is the fate of the Universe?

*Some say the world will end in fire,
 Some say in ice.
 From what I've tasted of desire
 I hold with those who favor fire.
 But if it had to perish twice,
 I think I know enough of hate
 To say that for destruction ice
 Is also great
 And would suffice.*

-- Robert Frost

What is the Universe's Fate?



Today: Universe is expanding. What do you expect to happen next?

Competition: gravity vs inertia

Compare: Pop fly and rocket!

- Quantitative question
- Launch speed vs speed to escape Earth



or



?

What is the Universe's Fate?



For Universe it is still gravity vs speed.

- Gravity acts on mass of galaxies (pulling back)
- The speed is the speed of expansion

Both are observable!



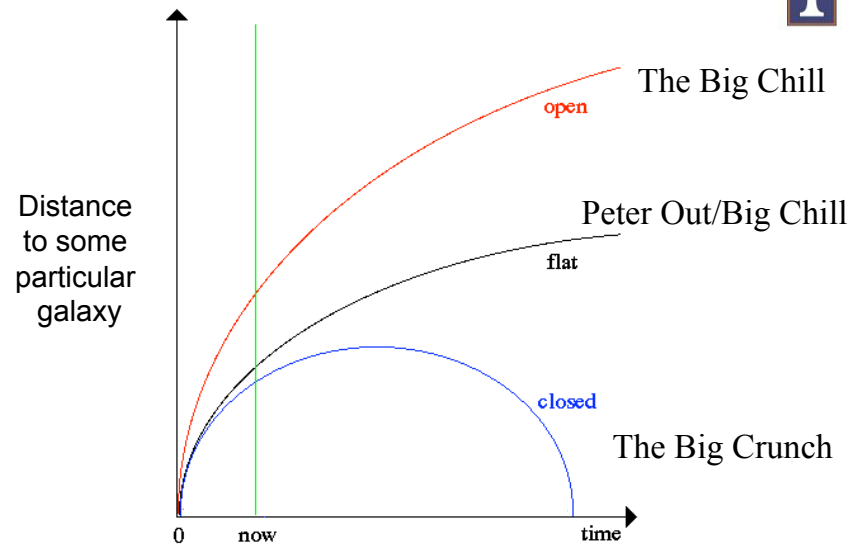
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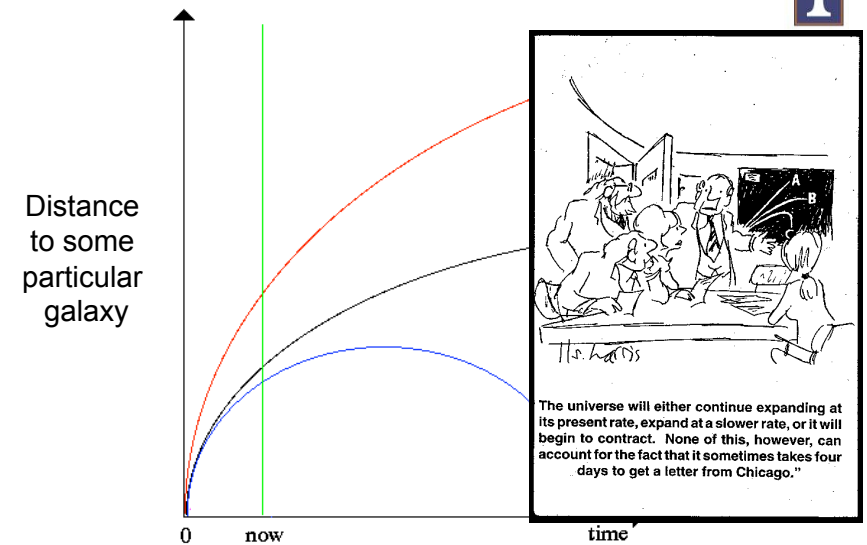
Our fate is a **quantitative** question :

- **If our mass is small enough we expand forever.**
- **If our mass is large enough expansion halts, and we collapse back.**

What kind of Universe do we live in?



What kind of Universe do we live in?



Big Chill/Big Crunch



- **Less mass:**
An open or flat Universe will end in a **Big Chill**:
 - Galaxies exhaust their gas supply
 - No more new stars
 - Old stars eventually die, leaving only dust and stellar corpses
- **More mass:**
A closed Universe will end in a **Big Crunch**:
 - Expansion will stop, and the Universe will re-collapse
 - Ends as it began, incredibly hot and dense

Question



Our Universe could be one of three types: Open, Closed, or Flat. What would happen to a closed Universe?

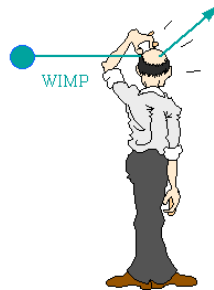
- No one else could get in.
- It would expand forever.
- It would just barely expand forever.
- It would expand for a while, then eventually begin to re-collapse on itself.
- It would expand, then slow down, then expand faster.

Dark Matter



- Outweighing regular matter is Dark Matter, probably some heavy exotic particle created during the Big Bang. (Weakly Interacting Massive Particle– WIMPs?).
- As mentioned before, most of our Milky Way is Dark Matter
 - We can't see it (only interacts via gravity)
 - We aren't sure what it is, but it is much more common than "normal matter"

How to search for WIMPs?

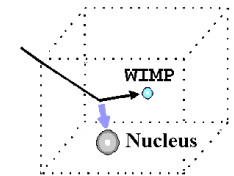
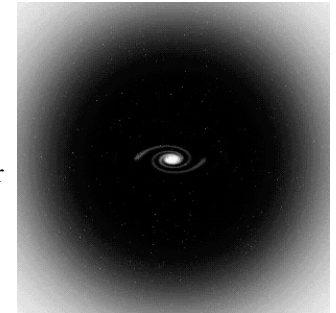


<http://www.shef.ac.uk/physics/research/pa/DM-introduction-0397.html>

Dark Matter



- The dark matter in the Galaxy is in greatly extended halo
 - Up to 90% of the Galaxy's mass is dark matter!
 - Galaxy may have over a trillion solar masses total!
- Dark matter is likely streaming through us right now!
- Recent suggestion of a detection. Stay tuned!



How Much Do We Weigh?



% of mass for closed Universe

22% Dark matter

Needed to explain:
galaxy rotation curves
clusters of galaxies

4.5% Ordinary matter

Made of protons, neutrons, and electrons

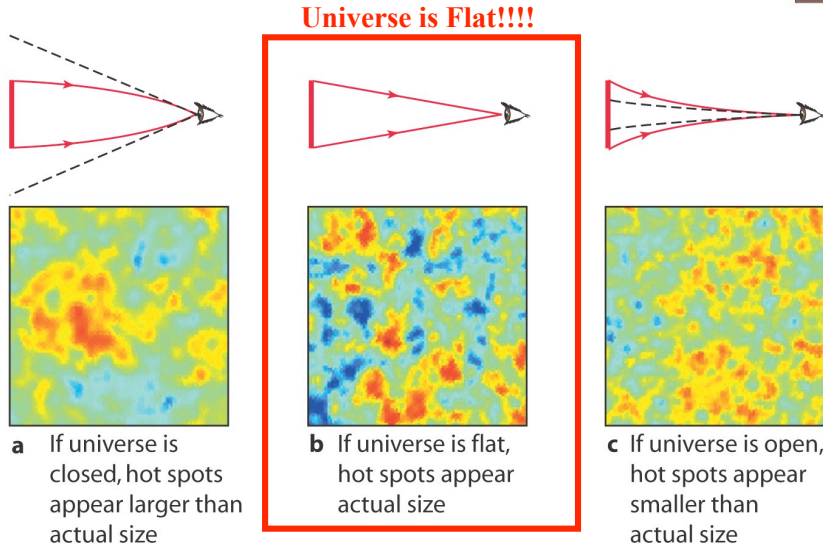
<1.5% Neutrinos

28% Total Not enough to close the Universe

So we live in an open Universe?



CMB Measurements



Peter Out/ Big Chill



- The Universe will just barely expand forever, getting cooler and cooler.
- If all of the mass, dark+regular, isn't enough, **then what's up?**
- The fate of the Universe is really dependent on the amount of matter and energy in the Universe → $E = mc^2$

We must be missing some extra mass/energy?

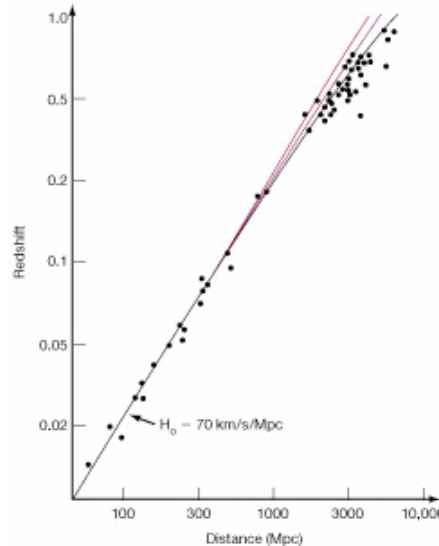


The Accelerating Universe!!!



The universe is not slowing down at all. In fact, it's speeding up!!! We live in an accelerating universe!

It's as if there's another force pushing the universe apart – a **Cosmological Constant!!!**



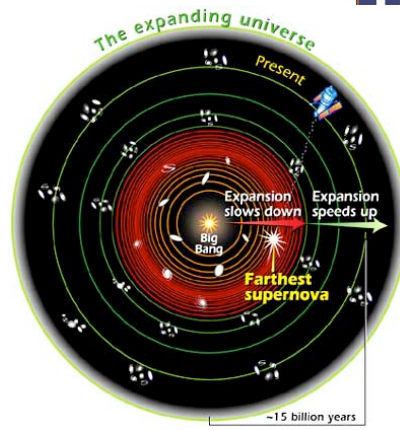
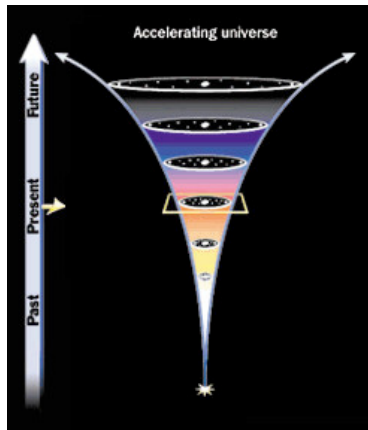
Dark Energy



- The matter census isn't enough to be flat and the expansion is accelerating!
- So, a new type of energy called **dark energy** must exist
 - Not related to dark matter
 - Acts as repulsive gravity, pushing apart.
- Dark energy is actually *accelerating* the expansion of the Universe!

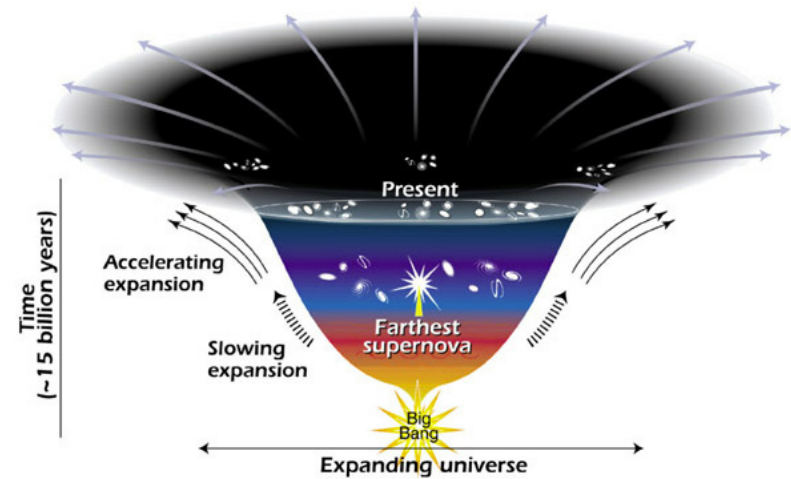


The Accelerating Universe!!!



Whatever this force is, we *think* that it is growing stronger as the universe evolves. The more empty space in the universe, the greater the acceleration – as if the vacuum of space has energy.

Effects of Dark Energy



<http://www.lbl.gov/Publications/Currents/Archive/Apr-06-2001.html>

Question



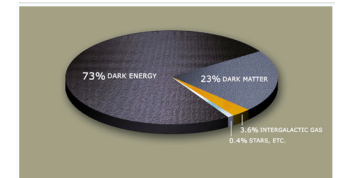
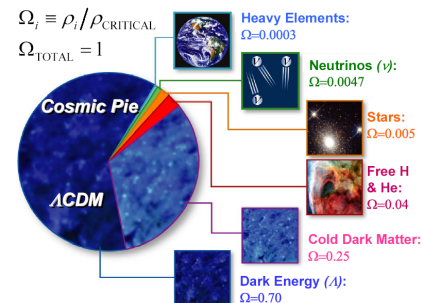
Based on measurement of the CMB, we live in a flat Universe. But there is not enough known mass to account for this. What's up?

- We must be underestimating the amount of Dark Matter.
- It would expand forever.
- It would just barely expand forever.
- We have something called Dark Energy.
- It would expand, then slow down, then expand faster.

The Accelerating Universe!!!



The universe is 13.7 billion years old, and it is now dominated by dark energy.



Dark Energy even dwarfs dark matter! Regular matter is really insignificant. We *really* don't know anything about what's going on!!

Question

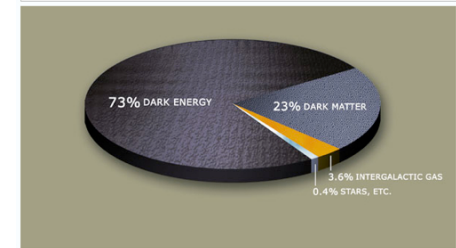
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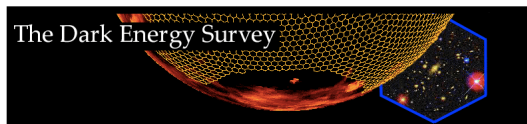
The Accelerating Universe!!!

- But, we are still in very speculative times here.
- How the Universe ends will depend on the nature of Dark Energy.
- If it really acts like a cosmological constant (go Einstein!), then we live in a flat Universe that will keep expanding forever, but if not then we don't know yet.



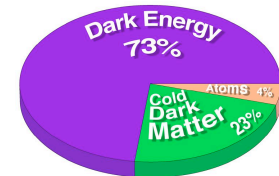
The Accelerating Universe!!!

- Understanding dark energy is one of the biggest questions for humankind today.
- There are many experiments underway to accomplish this. So we have to wait and see.
- But what are the options?



The Distant Future: The Big Rip

- Although this is not very popular, and the chances of it occurring is small, what if Dark Energy is not a cosmological constant?
- One extreme case is that it gets carried away, and rips the Universe apart.



- If repulsive force increases– Brooklyn may expand too.
- Gravity/E&M forces can not hold Galaxies rip apart
- Could rip the Milky Way apart in ~1-100 billion years
- Earth gets ripped apart soon after
- You'd get ripped apart!

<http://www.youtube.com/watch?v=oGVYG0ce1Ps>



The Early Universe?



- So, in the early Universe, the first elements formed were mostly Hydrogen (75%) and Helium (25%) by mass. What does that mean for life in the early Universe?
- Globular clusters contain the oldest stars in the Milky Way— about 10 to 13 billion years old. Should we look for life around these stars?



<http://www.shef.ac.uk/physics/research/pa/DM-introduction-0397.html>

What is the Earth made of?



- Very little hydrogen and helium. They make up less than 0.1% of the mass of the Earth.
- Life on Earth does not require any helium and only small amounts of non-H₂O hydrogen.
- All of these elements must be formed in stars. That means 2nd or 3rd or nth generation of stars are required before life can really get going. These elements were not originally formed in the Big Bang.
- **“We are star stuff!”**
- How did that come about?



What are Galaxies?



- They are really giant re-cycling plants separated by **large** distances.
- Stars are born in galaxies out of dust and gas.
- Stars turn hydrogen into helium, then into heavier elements through fusion for millions or billions of years.



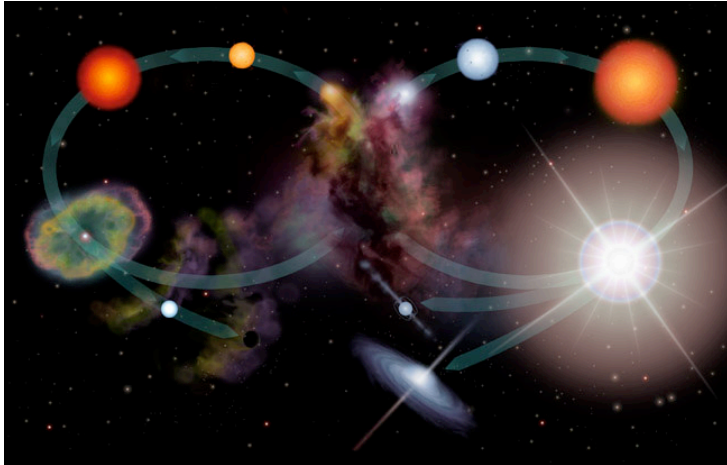
What are Galaxies?



- Stars die and eject material back into the galaxy.
- New stars are formed.
- And so on.
- Crucial to the development of life!



Stellar Evolution Re-Cycle



The Interstellar Medium (ISM)



- Stuff between the stars in a galaxy.
- Sounds sort of boring, but
 - Actually very important
 - Features complex physical processes hidden in safe dust clouds
- Every star and planet, and maybe the molecules that led to life, were formed in the dust and gas of clouds.
- Exists as either
 - Diffuse Interstellar Clouds
 - Molecular Clouds



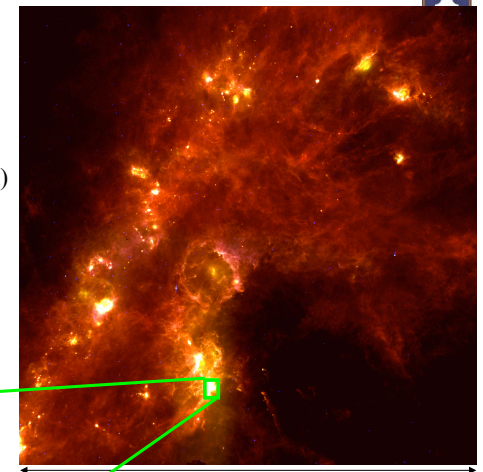
Keyhole Nebula



Giant Molecular Clouds



- Cool: < 100 K
- Dense: $10^2 - 10^5$ H_2 molecules/ cm^3
(still less dense than our best vacuum)
- Huge: 30 – 300 lyrs across,
 $10^5 - 10^6$ solar masses
- CO molecular emission & dust emission trace structure



100 degrees

Infrared image from IRAS

Orion Nebula

(near infrared)

Nearest massive star forming region with a large molecular cloud associated (distance of 1500 lys)



WFPC2
Orion Nebula • OMC-1 Region
 PRC97-13 • ST ScI OPO • May 12, 1997
 R. Thompson (Univ. Arizona), S. Stolovy (Univ. Arizona), C.R. O'Dell (Rice Univ.) and NASA

NICMOS

Hubble Space Telescope

Orion Nebula
 Subaru Telescope, National Astronomical Observatory of Japan

CISCO (J, K' & H₂ (v=1-0 S(1)))
 January 28, 1999

Trapezium cluster:

< 10⁵ yr old
 (largest star ~30 solar masses)

star density >
 10⁵ stars pc⁻³

0.07 pc

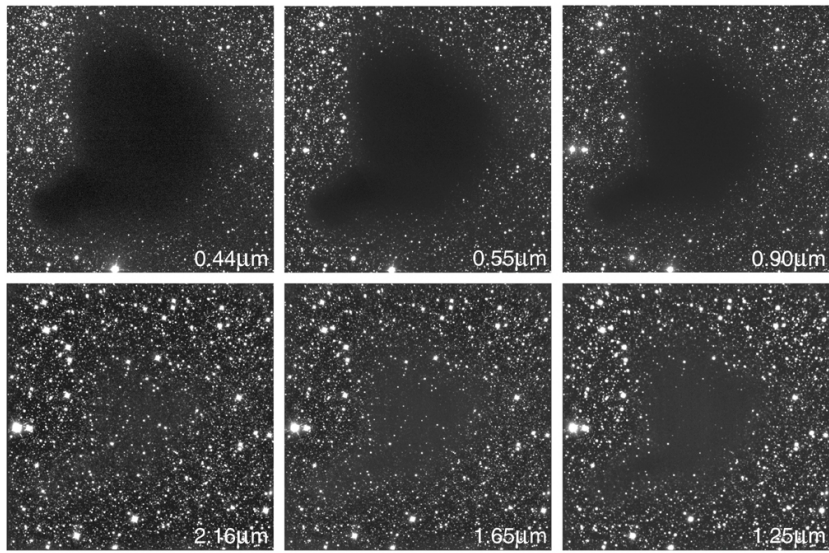
Interstellar Clouds

Bok globules

Reflection nebulae

Dust is concentrated in the galaxy's midplane

We see spiral galaxy NGC 891 nearly edge-on



The Dark Cloud B68 at Different Wavelengths (NTT + SOFI)

ESO PR Photo 29b/99 (2 July 1999)

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