Astronomy 330



Presentations



• For those presenting tomorrow, remember that the files must be emailed to me or Yiran by 1700 today!

- If not, points will be lost!

This class (Lecture 14):

Next Class:

Life in the Solar System

Origin of Life

Music: Venus as a Boy-Bjork

Outline



- Life on Europa
- Life on Titan
- Stars suitable for life.

Drake Equation





That's 2.6 planetary systems/year





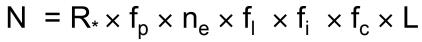












of advanced civilizations we can contact in our Galaxy today

Star formation

Fraction of stars with planets

Earthlike planets

Fraction on which that evolve life arises intelligence

Fraction

Lifetime of that advanced communcivilizations icate

stars/

yr

0.29 systems/

star

planets/ system

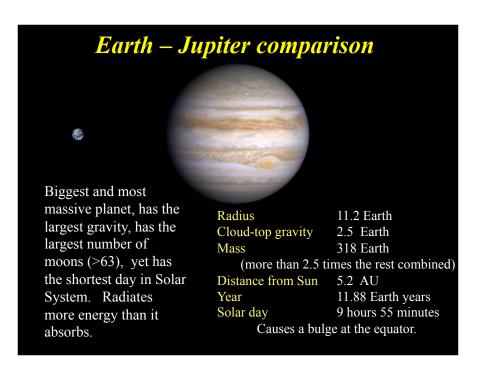
system

life/ planet

intel./ life

comm./ intel.

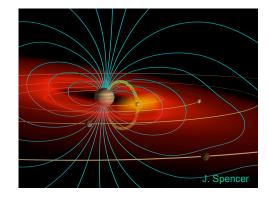
yrs/ comm.



Jupiter's Magnetosphere



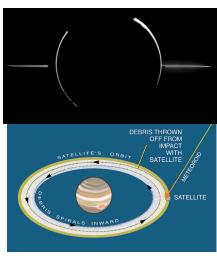
- Liquid metal hydrogen generates a magnetic field
 - 14x stronger than Earth's field
 - Over 4 million km across
- A ring of ionized particles surrounds Jupiter
 - Stripped from Jupiter's moon Io



Jupiter's Rings



- Jupiter has rings!
- Discovered by the Voyagers
- Not prominent like Saturn's
- Dusty disk of debris, probably from meteoroid impacts with small moons



The Galilean Moons



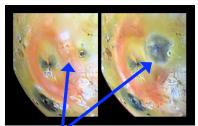
- Io is active.
- Europa is now thought to be the best option for life.
- But, Ganymede and Callisto are contenders perhaps for ancient life.

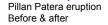


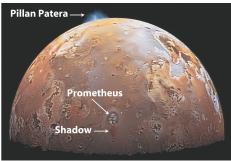
Io

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- Innermost Galilean moon the "pizza moon"
- The most volcanically active body in the solar system.
- Voyager 1 discovered presence of volcanoes
- Internal heating by Jupiter's tides
- Atmospheric gases ripped off by Jupiter's magnetic field ion torus



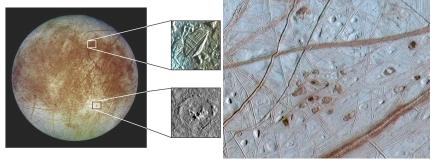




Europa

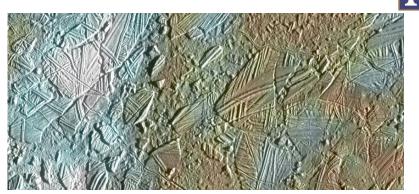


- Slightly smaller than our Moon.
- Icy crust 5 km thick. Can protect life against magnetic fields.
- Evidence for deep (50 km!) liquid water ocean beneath crust-remains liquid from tidal forces from Jupiter
- Cracks and fissures on surface upwelling?



Galileo

Europa

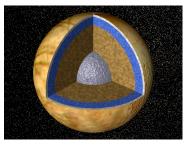


- Young surface few craters
- Tidal forces pull and push the ice
 - Like Io, it probably has strong tidal forces.

Europa

- Life would have to be below the surface, around hydrothermal vents.
- Very encouraging, as early life on Earth, might have been formed around such vents.
- We don't how thick the ice is yet future missions, will have to employ melting or smash and dive spacecraft.





Ganymede

Callisto

- Largest of the Galilean Moons
- Partly ancient surface, partly younger surface
 - Younger surfaces about the age of the Moon's maria
- Compared to our Moon:
 - 50% larger
 - 100% more massive
 - 40% less dense
- Interior more differentiated than Callisto, probably has an iron core
- May have a water ocean under surface.



Moons from Jupiter • Ancient surface,

- covered with craters
- Compared to our Moon:

• Furthest of the Galilean

- 40% larger
- 50% more massive
- 45% less dense
- Surface is made of "dirty ice"
- Interior is rocky, mixed with ice

Finding JIMO

- Jupiter Icy Moon Orbiter
 - To launch in 2017 or later
- Study Callisto, Ganymede, and Europa
 - Investigate makeup
 - Histories
 - Potential for sustaining life



- ESA-working title: Laplace
- Early planning stages of NASA/ESA/JAXA mission.
- Two or three orbiters
 - Launch date around 2020





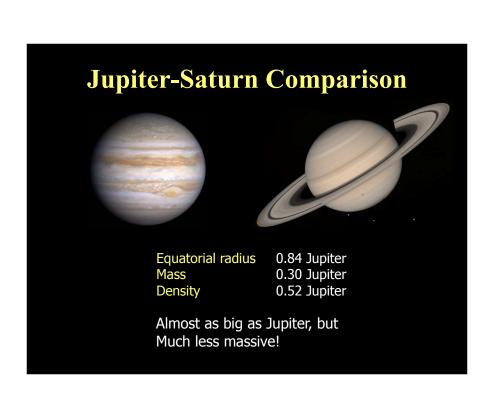


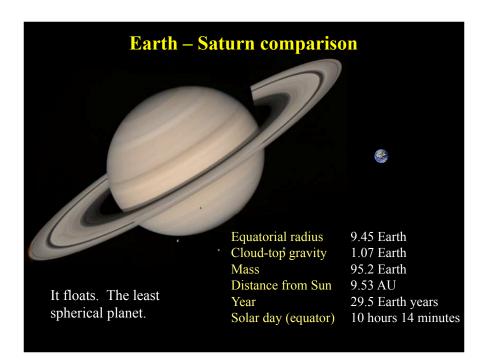
Question



The best place to look for life in the Jupiter system is

- a) in the frozen oceans of Callisto.
- b) in the frozen oceans of Ganymede.
- c) in the upper atmospheres of Jupiter, floating life.
- d) deep in the atmosphere of Jupiter, diamond bodied life to withstand the pressures.
- e) under the ice on Europa.







- Named for the father of the Roman gods
- Saturn is very similar to Jupiter
 - Large planet
 - Mostly liquid hydrogen
 - Has a mini-solar system
 - At least 60 moons
 - Most are small



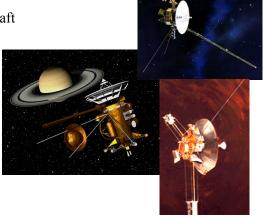
http://www.solarviews.com/cap/ sat/saturn.htm

http://saturn.jpl.nasa.gov/cgibin/ gs2.cgi?path=./multimedia/ images/saturn/images/ PIA05380.jpg&type=image

Missions to Saturn

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- There have been 4 unmanned spacecraft missions to Saturn
- Pioneer 11
 - Flyby 1979
- Voyager 1Flyby 1980
- Voyager 2
 - Flyby 1981
- Cassini-Huygens
 - Arrived 2004



The Cassini Mission



- Launched on October 15th, 1997
- Arrived at Saturn on July 1st, 2004
- Orbiting Saturn, making flybys of the planet, its rings, and some of its moons
- Contains 12 scientific instruments
- Also carries the Huygens probe, which was dropped onto Titan, Saturn's largest moon on Jan 2005. Remember?

Saturn's Atmosphere

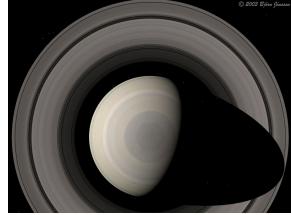


- Composition similar to Jupiter
 - Mostly hydrogen and helium
- Atmosphere more "spread out"
 - Less gravity
 - Contrast of cloud bands reduced
- Wind speeds fastest at the equator
 - 1000 km per hour!

Driving Saturn's Weather

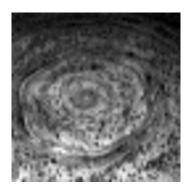


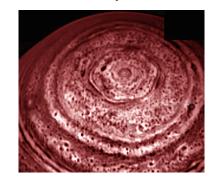
- As on Jupiter, Saturn's internal heat drives weather
 - Saturn radiates 80% more heat than it receives from the Sun
 - Like Jupiter,Saturn is still contracting!
 - As is contracts, heat is produced



Driving Saturn's Weather

- As on Jupiter, storms are produced between cloud bands
 - No long lasting storm like the Great Red Spot, but hexagon cloud at pole has been stable for 20+ years.

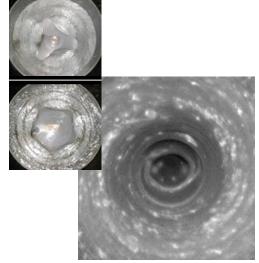




Driving Saturn's Weather



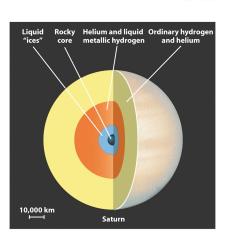
- Spinning water bucket experiments show similar features.
- Pseudoscience posit sound wave reflections.
- Saturn's South Pole also has an unusual structure.



Saturn's Interior

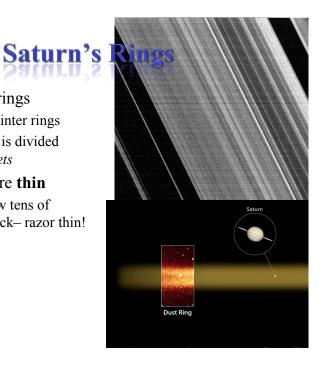


- Similar structure to Jupiter's
 - But Saturn is less massive
 - The interior is less compressed
- Liquid metallic hydrogen creates a magnetic field
 - 30% weaker than Earth's



• Two main rings

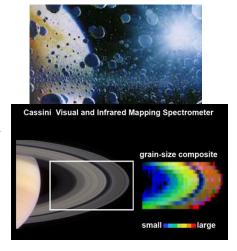
- Several fainter rings
 - Each ring is divided into ringlets
- The rings are thin
 - Only a few tens of meters thick-razor thin!



Makeup of the Rings

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- The rings of Saturn are **not** solid rings
 - Made of icy rocks
 - 1cm to 10m across
- New Cassini data shows ring particle size varies with distance from Saturn
 - Note the gap is filled with small particles



Saturn's Moons



- Saturn has a large number of moons
 - At least 60
- Only Titan is comparable to Jupiter's Galilean moons
- Smaller moons are mostly ice, some rock

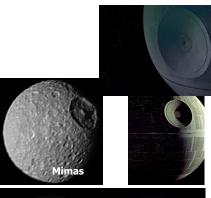


Saturn's Odd Moons

- **Mimas** Crater two-thirds its own radius
- Enceladus Fresh ice surface, water volcanoes?
- **Hyperion** Irregularly shaped
- **Iapetus** Half its surface is 10x darker than the other half
- Phoebe Orbits Saturn backwards



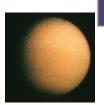




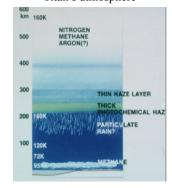


Titan

- Saturn's largest moon– bigger than Mercury.
- 2nd largest moon in the solar system after Ganymede.
- Discovered in 1655 by Christiaan Huygens
- Only moon to have a dense atmosphere
 - Dense nitrogen atmosphere
 - Small greenhouse effect
 - 98% nitrogen
 - -Only Earth is comparable
 - -Methane (something producing it)
 - Much like ancient Earth!

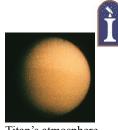


Titan's atmosphere

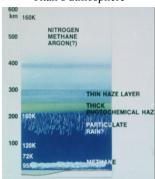


Titan

- Atmospheric pressure is 1.5 times Earth's
- Organic compounds life?
 - Probably not too cold: 95 K
 - May be a "deep freeze" of the chemical composition of ancient Earth



Titan's atmosphere



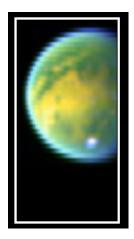
Piercing the Smog



- Cassini has special infrared cameras to see through Titan's smog
- Green areas are water ice
- Yellow-orange areas are hydrocarbon ice

Jan 14, 2005

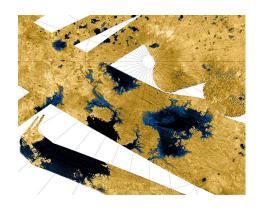
• White area is a methane cloud over the south pole



Surface Liquid



- · Now confirmed to have liquid on surface.
- · Only body besides the Earth.
- Too cold for water, so most likely filled with liquid ethane, methane, and dissolved nitrogen



http://photojournal.jpl.nasa.gov/catalog/?IDNumber=PIA10008

Cassini-Huygens Arrival at Saturn July 1, 2004 Huygens Probe descent to Titan

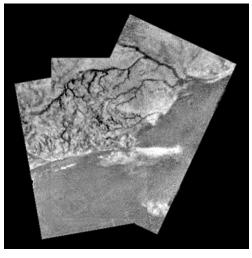
A Possible Landing

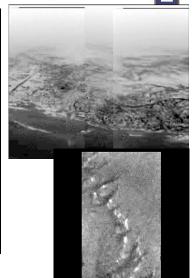


- The probe floating in the methane/ethane sea of Titan.
- Mountains in the distance.

http://saturn.jpl.nasa.gov/cgibin/gs2.cgi?path=../multimedia/images/artwork/images/

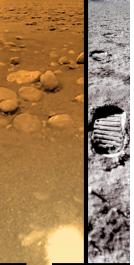
Mapping Titan



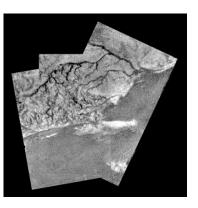


Mapping Titan





Mapping Titan

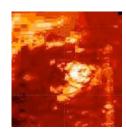


http://esamultimedia.esa.int/multimedia/esc/esaspacecast001.mp4



Cryovolcanoes

- Methane may come from volcanoes.
- Volcanoes heat up rock hard ice, spewing "lava" made up of water and ammonia
- Two hot spots found in atmosphere, suggesting eruptions.
- Mountains found, suggesting some sort of plate tectonics.



Conclusion

- No conclusive evidence exists for life in our solar system besides on Earth
- But, possibilities exist for life
 - Venus's clouds may have migrated life.
 - Mars may have some microbial history linked to water, and perhaps some subsurface life.
 - Jupiter's reducing atmosphere may harbor sinkers.
 - Europa's sub-crustal oceans may harbor life, even fish-like life.
 - Titan is still very interesting
 - · Thick atmosphere
 - · Reducing chemistry



Life on Titan



- Conditions much like the early Earth.
- Can organic chemistry work well in this environment?
- If found, would revolutionize our understanding of life.
- Some researchers suggest that panspermia from Earth is likely, so might find our cousins.
- Future missions will need to have biological component.

Question



Why is Titan an interesting place to look for life?

- a) It will revolutionize how we think about ET life.
- b) It will create new life hybrids.
- c) There is no chance of life there.
- d) The life is in early state if at all.
- e) Black beans.



Optimism?

- Carl Sagan argues for $n_p > 3$.
 - If Venus had less clouds (less greenhouse) it could have been cool enough for life.
 - If Mars had a thicker atmosphere it could have been warm enough for life.
 - If solvents other than water were used, maybe the moons of the outer planets?
 - Giant Jupiter-like planets close in?
 - Non-Earth life?



http://www.uranos.eu.org/biogr/sagane.html http://spider.ipac.caltech.edu/staft/jarrett/sagan/sagan.html

n_p: number of life planets per planetary system (average)



- Can range from 0.01 to >3.
 - Is seismic activity necessary to recycle bioelements?
 - How important is the first atmosphere? Ozone?
 - Is a moon needed? A large Jupiter-like planet?
 - Is liquid water a requirement? Other solvents okay?
 - Not too hot, not too cold; not too much pressure, not too little– Goldilocks requirement?
 - Habitable Zone around the star.
 - Galactic Habitable Zone
 - Does atmosphere need feedback mechanism?
 - But in our solar system, maybe 5 nearly possible life planets.



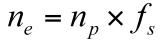
Pessimism?



- We only considered temperature. What about:
 - Gravity?
 - Atmospheric pressure?
 - Size of the moon or planet?
 - Does life need a Moon-like moon? Does life need the tides? Does the Moon protect the Earth's rotation? Is a Jupiter needed?
- If we impose Earth chauvinism, we can easily reduce to $n_p \sim 0.1$



http://sagiru.tripod.com/Travel/Lost_in_the_Sahara/lost_in_the_sahara



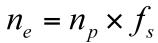


n_p: number of planets suitable for life per planetary system

Class number is?



http://nike.cecs.csulb.edu/~kjlivio/Wallpapers/Planets%2001.jpg





n_p: number of planets suitable for life per planetary system

f_s: fraction of stars whose properties are suitable for life to develop on one of its planets

We can list 5 situations that will have an effect on f_s .



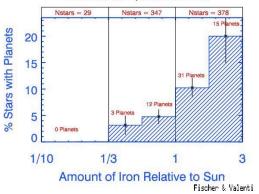
http://nike.cecs.csulb.edu/~kjlivio/Wallpapers/Planets%2001.jpg

Differences of Stars to Life



Metal rich stars. Stars with heavy elements, probably more likely to have planets. Suggested in the current planet searches. About 90% of all stars have metals.

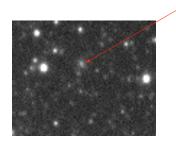
Planet Occurrence Depends on Iron in Stars



Differences of Stars to Life



2. <u>Main sequence stars</u>. Need the brightness to stay as constant as possible. Otherwise the temperature changes dramatically on the planets. This is 99% of all stars.

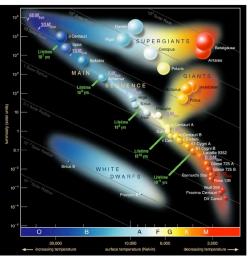


Differences of Stars to Life



3. Length of time on the main sequence.
We needed temperature stability for 5 billion years to get intelligence on

Earth. This rules out stars more massive than 1.25 solar masses! 90% of all stars are less massive than that.

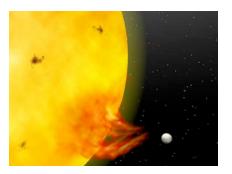


http://mjbs.org/hr.jpg

Differences of Stars to Life



4. Minimum mass of star. If ice exists close to the star. that would imply the formation of Jupiter-like planets not Earth-like planets. And, any life bearing planet would have to be closer to the star- and closer to stellar effects (e.g. tidal locking and more flares from low mass stars). That limits us to a minimum of 0.5 solar masses. 25% of all stars are more massive than that.

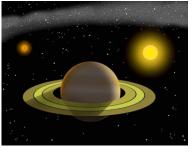


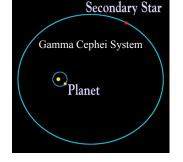
http://spaceflightnow.com/news/n0401/19planet/planet.jpg

Differences of Stars to Life



5. Binarity. Planets may form. But they may have odd orbits unless the 2 stars are far enough apart or the planet orbits the pair. Only 30% of all stars are single stars. 50% of all stars are single stars or wide binary stars.





http://spaceflightnow.com/news/n0210/11planet/

Question



The best type of life sustaining stars are

- a) Low mass stars (less than 0.5 solar masses), as life can exist nearer the star where more terrestrial planets are probably located.
- b) Binary stars, as they double the chances of life.
- Stars off the main sequence, as they have lived the longest, they are the best chance for finding intelligent life.
- Middle mass stars (less than 1.25 and more than 0.5 solar masses), as they live longer and don't require the planets to be too close.
- e) Massive stars (more than 2 solar masses), as they have more mass from which to form planets.

Drake Equation





That's 2.6 planetary systems/year

















 $N = R_* \times f_D \times n_e \times f_I \times f_i \times f_c \times L$

of advanced civilizations we can contact in our Galaxy today

Star formation rate

Fraction of stars with planets

Earthlike planets system

Fraction on which

Fraction that evolve life arises intelligence

that communicate

Lifetime of advanced civilizations

stars/

vr

0.29 systems/

star

planets/ system

life/ planet

intel / life

comm./ intel.

yrs/ comm.

So Far, We have Studied



- The Universe
 - Big Bang
 - Creation of hydrogen, helium...
 - Galaxy formation
 - Swirls of elements embedded in self-gravitating cloud of dark matter
 - Star birth
 - Energy generation and element production in selfgravitating mass of gas
 - Planets
 - Ice, rock, gas surrounding stars form planetesimals, then planets

Life on Earth



- In our scientific approach, we look at life as a result of chemical evolution of complexity.
- We will view the formation of "life" on planets as we did star formation
 - A natural consequence of natural laws
 - More specifically, as a consequence of the complex chemistry that is sometimes achieved.



http://www.toothpastefordinner.com/052802/science-only-happens.git

Life on Earth



- Time to examine terrestrial evolution.
- Need to understand what is needed for life to arise.
- Again, some Earth chauvinism.
- Relies on chemical evolution
- Eventually life began?



http://www.accessexcellence.org/bioforum/bf02/awramik/bf02a1.html

Cosmic Imperative?

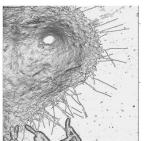


- But is life a cosmic imperative?
- Just like gas forms galaxies, and in galaxies stars and planets form, do chemicals on some planets form molecules that lead to life?















Element Basis of Life



• About 95% of the mass of all terrestrial organisms is composed of only 4 out of 90 elements

(61% in humans) Hydrogen Oxygen (26% in humans) (2.4% in humans) Nitrogen (10.5% in humans) Carbon

• HONC is essential to life, and it's common in space.

Question



Life on Earth is varied in how its made on the molecular level, i.e. elephants are made out of different stuff than bacteria.

- a) True
- b) False

Trace Elements



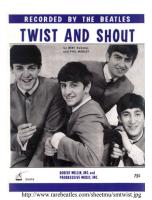
In addition to HONC, there are some other elements that are essential for life but in *smaller* amounts:

- Sulfur, magnesium, chlorine, potassium, sodium
 - These other elements make up about 1% of the mass of living organisms
 - Exist in roughly the same concentration in organisms as in ocean water
 - Highly suggestive that life began in oceans
 - Furthermore suggests that the evolutionary processes occurred on Earth. Panspermia problems?



Good News

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- H,O,N,C is very common in universe; everywhere as far as we can tell
 - If life were based totally on rare elements, we might expect its occurrence to be extremely rare...
- So, we expect ET life to be based primarily on HONC.
 - The four primary chemical elements of life with some other simple components can produce staggering complexity.
- But, each planet will feature its own environment of trace elements giving each planet's life a unique twist to the standard HONC chemistry



We Are Special Stuff?



- Why is Earth life based on H,O,N,C instead of the more abundant elements found on Earth?
 - Suggests that the formation of life is not able to be formed just out of anything lying around.
 - The selection of H,O,N,C seems to be a <u>necessity</u> of the chemistry of life.
 - In general, Earth life is a carbon based life. Carbon is the main backbone of the chemistry.
- Is this good news?

Nature's Complexity



- The workings of biological molecules are an absolute marvel
 - How did this complexity develop?
 - How did it evolve?
- As complex and mysterious as life on Earth may be, we can begin to understand it
- Start with the basics:
 - Why are H,O,N,C the basis for living organisms?
 - How do the molecules formed by these (and other elements) work to make DNA, proteins, life?

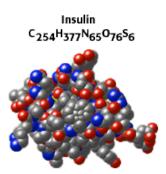


http://europa.eu.int/comm/environment/life/toolbox/logo life high resolution 2.jpg

Why Carbon Based Life?



- Carbon's electronic structure allows it to form long chains
 - Chains of atoms and chains of molecules complexity
 - Life needs bonds to be stable but breakable
- Good for us, at temperatures at which water is liquid, carbon bonds are stable but breakable
- Organic chemistry is the special branch devoted to carbon chemistry.



http://www.biology.arizona.edu/biochemistry/tutorials/chemistry/page 2.html