

# Astronomy 230

## Section 1– MWF 1400-1450

### 106 B6 Eng Hall



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This Class (Lecture 22):

Lifetime of Civilizations

Next Class:

Class Participation

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# Outline



- We need to estimate the average lifetime of all civilizations in our galaxy.
- What factors determine the lifetime of a civilization?
  - Resource Exhaustion
  - Population growth
  - Nuclear war
  - Natural catastrophe
- Arguably, the most uncertain factor in the Drake Equation.

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**= 3.6**

## Drake Equation

Communicating Civilizations  
/century

Frank  
Drake



$$N = R_* \times f_p \times n_e \times f_l \times f_i \times f_c \times L$$

# of advanced civilizations we can contact	Rate of formation of Sun-like stars	Fraction of stars with planets	# of Earthlike planets per system	Fraction on which life arises	Fraction that evolve intelligence	Fraction that communicate	Lifetime of advanced civilizations
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**10      0.34      0.208      0.235      0.265      0.823**

Stars/year    Planetary    Livable Planets    Evolved Life    Intelligence    Comm. ET

System/star    /Planetary System    /Livable Planet    /Evolved Life    /Intelligence

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## Lifetime of Civilization



- If a civilization can communicate with other life forms, and wants to, how long can it last?
- This factor pulls a lot of weight in the Drake equation. Are we alone or are there aliens everywhere?
- Easy to envision 4 cases:
  1. Communication efforts stop. Bored with lack of success or funding issues.
  2. Civilization evolves away from interest or capability. But empires rise and fall.
  3. Technological civilization collapses: exhaustion of resources and population growth,
  4. Catastrophe! Nuclear war or various natural problems.

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## Issues



- The last 2 items:
  - Technological civilization collapses
  - Catastrophe
- Could be caused by:
  - Resource Exhaustion
  - Population growth
  - Nuclear war
  - Natural catastrophe

### Hiroshima



<http://gawain.membrane.com/hew/Japan/Hirosh.html>

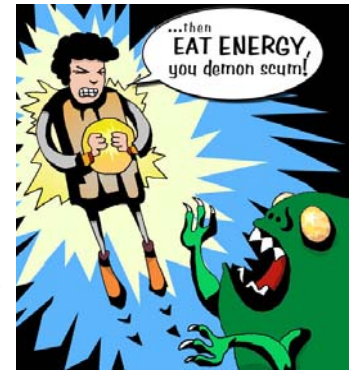
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## Depletion of Resources



- Modern life depends on metals and rare elements.
- Recycling can delay the depletion.
- Pollution of our water or air supply is still a problem.
- **But**, many of these issues can be solved with sufficient **energy**.
- Energy allows us to recycle, remove salt from the oceans, grow more crops, and generally convert material into the form we need.
- So, energy is our greatest concern.
- Remember that energy is not depleted, rather converted from useable form to less useable form (2<sup>nd</sup> law of Thermodynamics).



<http://www.timboucher.com/portfolio/eat-energy.jpg>

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## Energy



- Majority from chemical means– fossil fuels– electricity and gasoline (92% in the U.S.).
- Really are from fossils representing millions of years of life.
- And how are we spending it?
- The average US citizen uses twice that of a European, and 5 times the world average.
- Easy to obtain fossil fuels should last 50-100 yrs, coal 300-600 yrs.
- We will have to change! But US spending on renewable energy sources dropped by factor of 10 in the 1980s.
- SUVs do not help.



<http://www.orps.state.ny.us/sas/graphics/oilwells.jpg>



[http://www.dealerimpact.com/downloads/desktop\\_imgs/800x600-hummer.jpg](http://www.dealerimpact.com/downloads/desktop_imgs/800x600-hummer.jpg)

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## Nuclear Fission



- Breaking apart heavy (heavier than iron) unstable elements into lighter ones.
- Most widely used is <sup>235</sup>U– formed from supernovae– so limited amount on Earth.
- Supplies are limited and length of use controversial.
- A large reactor power plant uses 26 tons of fuel and 25 tons of waste per year.
- What do we do with the waste?
- How to prevent accidents: Three Mile Island or Chernobyl?



<http://www.ne.doe.gov/uranium/history.html>



<http://www.capefare.com/seasonone.php>

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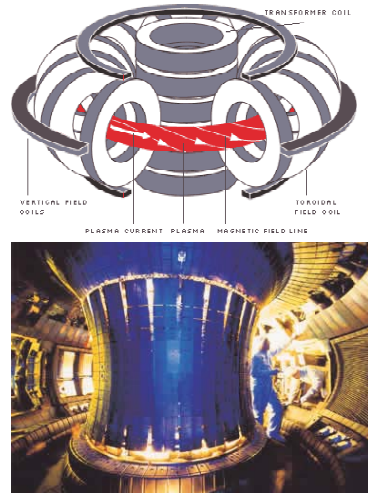
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# Nuclear Fusion



- What the Sun does for energy– H into He.
- Requires high density and temperature.
- How to contain it on Earth– Sun uses gravity.
- Magnetic confinement, but not easy.
- Research continues, but unlikely to play a large role in the next 50 yrs.
- And on Earth requires deuterium (heavy hydrogen) not as abundant as hydrogen, nonetheless very promising!

## Tokamak Fusion Reactor



<http://www.ipp.mpg.de/ippcms/eng/pr/exptypen/tokamak/magnetsputen/index.html>

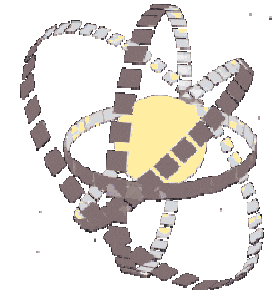
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# Long-Lived Civilizations



- Require renewable energy supplies, all Sun related.
- Hydroelectric (requires rain), windmills (winds), and solar power.
- Solar power is used today, but currently expensive because of manufacturing and tax subsidies for fossil fuels.
- Future example, could imagine a power plant that completely surrounds the Sun– e.g. Dyson sphere.



[http://www.homoexcelsior.com/omega.db/datum/megascale\\_engineering/dyson\\_sphere/237](http://www.homoexcelsior.com/omega.db/datum/megascale_engineering/dyson_sphere/237)

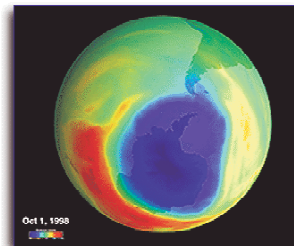
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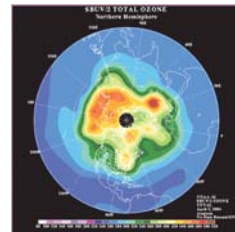
# Pollution from Civilization



- Ozone layer ( $O_3$ ) is formed when  $O_2$  is hit by ultraviolet light, which breaks up  $O_2$ .
- Ozone protects life against harmful Sun rays.
- Chlorofluorocarbons (CFCs) destroy the ozone.
- CFCs were used in A/C and refrigeration.
- Governments did not do much until a large hole appeared over Antarctica and N. America.
- Finally, being phased out, but the CFCs take about 20 yrs to reach stratosphere.
- The problem was predicted 25 years ago.



ANTARCTIC OZONE HOLE PHOTO COURTESY OF NASA



[http://www.cpc.ncep.noaa.gov/products/stratosphere/sbuv2to/gif\\_files/sbuv16\\_nh\\_latest.gif](http://www.cpc.ncep.noaa.gov/products/stratosphere/sbuv2to/gif_files/sbuv16_nh_latest.gif)

<http://www.ngdc.noaa.gov/paleo/globalwarming/images/ozone.gif>

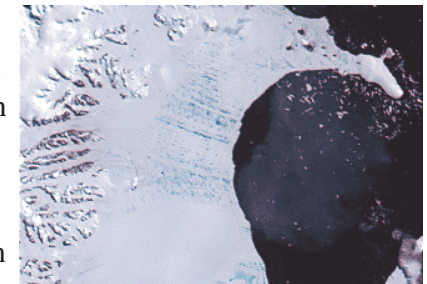
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# Global Warming



- Burning of fossil fuels releases  $CO_2$ .
- This is a greenhouse gas.
- Humans add more  $CO_2$  to the atmosphere (50-100x) than natural sources– 25 billion tons each year!
- Then why hasn't the temperature rise been more dramatic?
- The burning of coal releases sulfates form a haze that increases the albedo of Earth.
- So the effect is less than expected, but predictions suggest that  $CO_2$  content will begin to dominate in this century.
- Already, large slabs of the Antarctica ice shelf have melted.



Destruction of Larsen ice shelf 2002. 3250 km<sup>2</sup> over 35 days. That's bigger than Rhode Island! Existed for at least 400yrs maybe 12,000yrs.

<http://www-nsidc.colorado.edu/iceshelves/larsen2002/animation.html>

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# Population Growth



- Currently world population is around 5 billion ( $5 \times 10^9$ ).
- Population roughly doubles every 50 years—
  - 2050: 10 billion
  - 2100: 20 billion
  - 2150: 40 billion
  - 3000:  $2.6 \times 10^5$  times present population =  $1.3 \times 10^{15}$
- In the year 3000, each person will have 4 square feet (2' by 2') of space (including the oceans!).
- A final absurdity, in 2550 years, the weight of humans would outweigh the Earth.
- Obviously something will have to be done!



<http://w3.whosea.org/about/searo/88-97-7.htm>

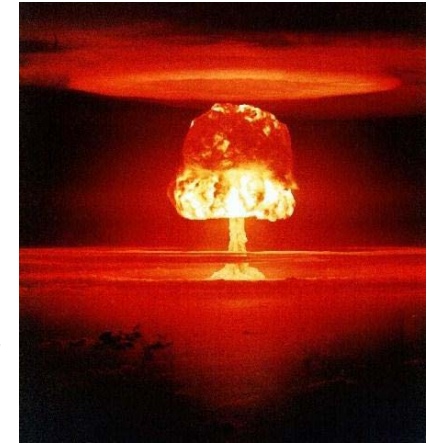
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# Nuclear War



- May be the only human activity that can catastrophically end our technological civilization.
- Effect may be seen days or years afterwards.
- Makes lots of radioactive elements with various half-lives.
- Most destructive global nuclear war could cause a nuclear winter.
- Dust and debris thrown into atmosphere around the globe would block light and lower temperatures.
- Out of control fires would add soot to the dust layer.
- Major collapse of the world's food chain.
- Possibly extinguish our species.



<http://www.dalistan.org/journal/rechist/nuclear/nuclear.html>  
<http://cosmo.pasadena.ca.us/adventures/atomic/cold-war.html>

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# Natural Catastrophes



## 1. Volcanoes

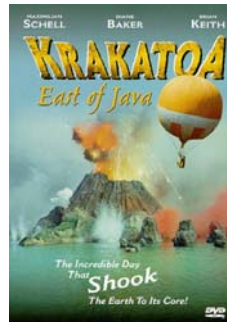
- Worldwide distribution of dust. Same idea as nuclear winter, but without radioactive fallout.
- Krakatoa eruption in 1883 near Java, blew away 75% of the island of Rakata. (Heard in Austria.)
- Prolonged low temperatures “Year with no summer”



From Simkin and Fiske, 1983



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<http://charm.hendrix.edu/astro/krakatoa.jpg>

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<http://www.vulkaner.no/v/volcan/indo/krakatau.html>

# Natural Catastrophes



## 2. Comets and Asteroids

- Many in Earth-Crossing orbits— NEOs.
- Again, creates large amounts of dust in the atmosphere leading to global cooling.
- Small objects can cause a lot of damage because the Earth's orbital velocity is 30 km/s  $\Rightarrow KE = \frac{1}{2} M V^2$
- That means that a 0.25 km radius rock releases as much energy as 7200 megatons of TNT, as much as a all-out nuclear war!
- Would make a 10 km crater a few km deep ejecting  $10^{12}$  tons of debris.

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## Be Aware



### DOCTOR FUN

11 April 96



Copyright © 1996 David Farley, d-farley@tezcat.com  
<http://sunsite.unc.edu/Dave/dr/fun.html>  
 This cartoon is made available on the Internet for personal viewing only.  
 Opinions expressed herein are solely those of the author.

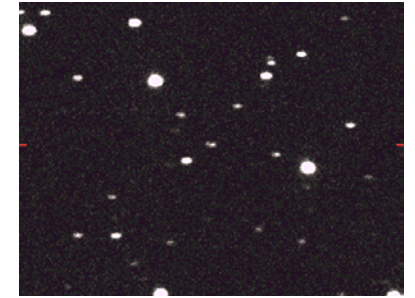
"Today's asteroid encounter was a near miss, but some scientists warn that an actual impact could have serious long-term effects on life on Earth as we now know it."

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## Killer Asteroids



- Small asteroids are often hitting the Earth's atmosphere.
- Commonly giving off around 10 kilotons of energy.
- But how often are Killer Asteroids (~0.5 km in diameter) expected?
- In 1992 congress asked NASA to find near Earth objects.
- So far over 400,000 objects.
- The most dangerous known is 1950 DA (~1km), will get close in March 2880 (0.33% chance of collision).
- We can not predict orbits more than 20 years in advance, but 1950 DA would have 100,000 Megatons of energy.



Asteroid 2004 FH. 30 meters in diameter. About 1 Megaton of TNT energy in an Earth impact! Passed within 7 Earth radii of Earth. Hiroshima was 15 kilotons.

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<http://antwrp.gsfc.nasa.gov/apod/ap040322.html>

## Killer Asteroids



- The Dino Killer was about 10 km in diameter.
- And, there are many asteroids out there that we still do not know about.
- Estimation of killer asteroids impact is about every million years or so.
- What can we do if there is an immediate threat? There may be little time.
- Diversion or destruction of object.
- With sufficient warning it doesn't take too much to miss the Earth.
- One example is to change reflectivity of surface.
- Nuclear explosions may result in many small asteroids.
- Expensive and difficult, but advanced civilizations should be able to do it.



<http://www2.ifa.hawaii.edu/newsletters/article.cfm?a=88&n=10>

<http://neat.jpl.nasa.gov/>

<http://www.ll.mit.edu/LINEAR/>

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## Natural Catastrophes



3. Stellar Evolution
  - The Sun is halfway through its lifetime on the main sequence.
  - Its luminosity will increase as it becomes a red giant.
  - In about 5 billion years, the Earth's atmosphere will probably evaporate.
  - Even earlier though, the Earth will lose its oceans in about 1-2 billion years.
  - But an advanced civilization can decrease greenhouse gases or increase dust in the atmosphere.
  - Eventually, we would have to leave the Earth and move to Mars.
  - Even shorter variations in the Sun's luminosity can result in ice ages. Again, advanced civilizations can add greenhouse gas.



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[http://www.boulder.swri.edu/~terrell/dart\\_old.htm](http://www.boulder.swri.edu/~terrell/dart_old.htm)

# Natural Catastrophes



## 4. Killer Supernovae!

- Death of a nearby massive star would be bad news.
- Explosion within 30 ly would destroy ozone layer.
- Right now, no candidates.
- Unlikely to happen in time scales of less than 2 billion years.
- As Brian Fields discussed, it is posited that a supernova event 2 Myrs ago may account for a mass extinction event.



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# Natural Catastrophes



## 5. Ultimate limit to L!

- Fate of the Universe.
- A Big Crunch:  $10^{12}$  years (a trillion years)
- But, WMAP results from the cosmic microwave background suggest that we are in a flat universe.
- Then it is an issue of energy.
- Eventually all of the stars will burn out ( $10^{12}$  years).
- Only energy source left is orbital energy.
  - Possibly extracting energy from rotating Black Holes.
- Eventually, black hole evaporate ( $10^{100}$  yrs). Remember the Universe is  $13.7 \times 10^9$  or around  $10^{10}$  years!
- But half of all protons might decay by  $10^{33}$  yrs.
- Bottom line is that the maximum age is speculative.

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# What is L?



- How long on **average** can an advanced civilization exist?
- Again, we only have a sample of 1 from which to discuss. What is our civilization's lifetime?
  - Short Term (100-1000 yrs)
    - Give up on communication due to budgets.
    - Depletion of resources.
    - Population.
    - War.
  - Long Term ( $10^5$  to  $5 \times 10^9$  yrs– age of galaxy is  $10^{10}$  yrs and we took half of that to evolve)
    - Stellar Evolution.
  - Don't forget the random volcano, asteroid, or supernova.
  - Still in many cases an advanced civilization may be prepared for many of the issues!

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