### Astronomy 230

y	230
	This class (Lecture 13):
	Life in the Solar System Heath Murra & Ryan Peterson
	Next Class:

Life in the Solar System

#### -HW 5 is due Thursday

Oct 17: Amanda Schultz & Joel Bonasera

#### Music: Venus as a Boy – Björk

Alien boogers

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## HW #3

- Ken Sampson: <u>http://sprott.physics.wisc.edu/PICKOVER/aliensc</u> p.html
- Amanda Schultz: <u>http://www.ufoevidence.org/</u>
- Bryan White:

http://www.ufoevidence.org/documents/doc220.ht m

- Ian Gentile: <u>http://www.iwasabducted.com/</u>
- Joe Coletta: <u>http://www.iwasabducted.com/</u>

### Midterm

- 1 hour exam in this classroom.
- It will cover material up to, but not including, "Life in the Solar System"
- Will consist of 15 multiple choice/ true-false questions (worth 40 points) and 2 essay questions (25, and 40 points each).
- A total of 105 points, so 5 extra credit points.
- You can bring a normal-sized sheet of paper with notes on both sides.

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## **Presentations**

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- Heath Murra : <u>Lunar Colonization</u>
- Ryan Peterson : Creating a Planet

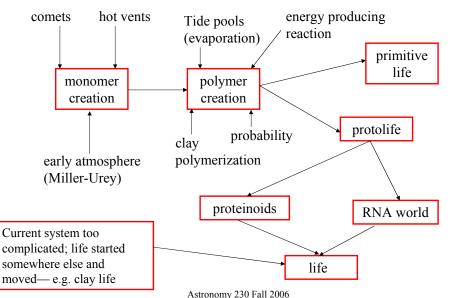
### Outline

- $f_1$ ٠
- Life in our Solar System?
  - Venus
  - Mars

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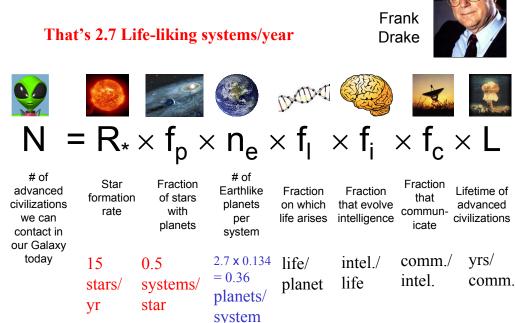
## **Pathways on Earth**





#### **Drake Equation**





### **Summing Up**



- Existence of organic molecules in space implies that amino acid complexity is common.
- Fact: On Earth polymers arose and evolved to life.
- Life it seems evolves naturally through a number of intermediate steps if conditions are right and  $f_1 = 1$
- But how often are the conditions right?
- Nonetheless, even with only a vague notion of how life on Earth evolved, it seems that there are possible pathways that take the mysterious polymerization to transition to life steps.
- Still a number of questions:

#### Summing for f<sub>l</sub>



- Is life a natural occurring consequence of the laws of nature?
- Will each planet from n<sub>e</sub> outgas and produce water?
- Will it have a reducing atmosphere?
- Will it have the right energy sources to produce life's monomers?
- Monomers from space?
- Will polymerization occur?
- Are tides necessary to wash polymers back into liquid water?
- Will basic life occur? Protolife or life?
- Alternative life?
- Maybe the conditions that produced life on Earth are unusual or maybe common.
- That means  $f_1$  can range from small numbers 0.0001 to 1.

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#### Earth – Venus comparison



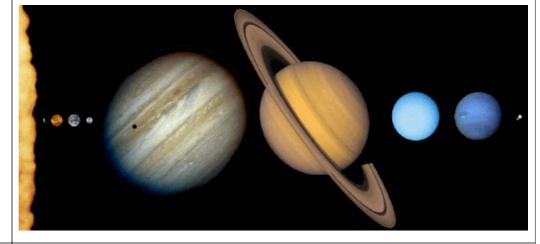
Venus is the hottest planet, the closest in size to Earth, the closest in distance to Earth, and the planet with the longest day. Radius Surface gravity Mass Distance from Sun Average Temp Year Length of Day Atmosphere

0.95 Earth 0.91 Earth 0.81 Earth 0.72 AU 475 C 224.7 Earth days 116.8 Earth days 96%  $CO_2$ 

## Life in the Solar System?



- We want to examine in more detail the backyard of humans.
- What we find may change our estimates of  $n_e$  or even  $f_l$ .



## What We Used to Think



Venus must be hotter, as it is closer the Sun, but the cloud cover must reflect back a large amount of the heat.

In 1918, a Swedish chemist and Nobel laureate concluded:

- Everything on Venus is dripping wet.
- Most of the surface is no doubt covered with swamps.
- The constantly uniform climatic conditions result in an entire absence of adaptation to changing exterior conditions.

• Only low forms of life are therefore represented, mostly no doubt, belonging to the vegetable kingdom; and the organisms are nearly of the same kind all over the planet.

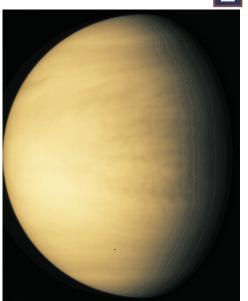
## **Turns Out that Venus is Hell**

- The surface is hot enough to melt lead
- There is a runaway greenhouse effect •
- There is almost no water
- There is sulfuric acid rain
- Not a place to visit for Spring Break.



#### **Our "Twin"**

- Always covered in thick clouds of  $CO_2$ , which make it the hottest planet in the Solar System.
- Pressure on surface is 90 times that on Earth–like 1 km under the sea
- Often called the morning star or the evening star. 3<sup>rd</sup> brightest object in the sky. Often mistaken for UFO.
- Retrograde rotation Sun rises in west
- No moons, no magnetic field



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http://antwrp.gsfc.nasa.gov/apod/ap960923.html

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## **Soviet Satellites on Venus**

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Color as seen on the surface of Venus

Venera 13

Color with atmospheric effects removed



USSR Academy of Sciences / Brown University

Mostly Basalts-like rocks, indicative of volcanoes Astronomy 230 Fall 2006

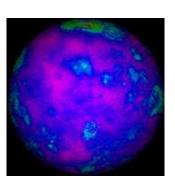
## **The Venusian Surface Revealed**

- We can't see Venus' surface in visible light, clouds block the view
- Magellan's Radar showed the surface
- Most of surface is smooth lava flows
- Many large volcanoes ٠
- Probable ongoing volcanism



### **Surface of Venus: Radar**





http://www.solarviews.com/raw/venus

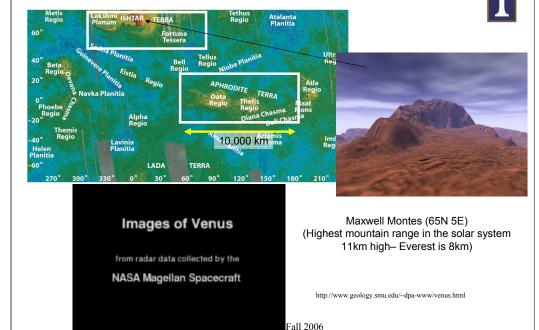
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## **Impacts on Venus**

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- Venus has about 1,000 craters, often clustered
- No trace of heavy bombardment
- Cratering rate indicates Venus' surface about 500 million yrs old
- Why?
  - Possibility: Extreme temperatures soften rock, making the surface subject to catastrophic volcanic upheaval

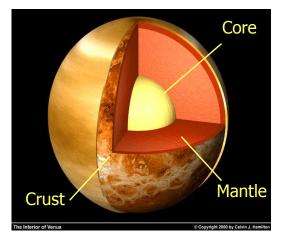




#### **Venus' Interior**



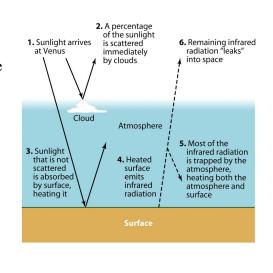
- Venus' size and density are roughly equal to Earth's
  - Indicates iron core of similar size
- No magnetic field
  - Very slow rotation 243 Earth days



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### **Runaway Greenhouse**

- On Earth, greenhouse gasses insulate us
  - Keep Earth 35 K warmer than it would be otherwise
- On Venus, massive amounts of CO<sub>2</sub> keep it incredibly hot
  - Almost 300 K warmer!
  - The hottest planet in the Solar System



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## Why So Different?

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- Earth's carbon is locked up
  - Dissolved in the oceans
  - Locked into rocks and life



- Venus' carbon is in its atmosphere
  - Too close to the Sun for liquid water
  - No oceans to trap the carbon dioxide
  - No life to process the carbon into sedimentary rocks



## What Happened to Venus?

- It really should have been more like Earth, but the atmosphere is much different.
- Earth's atmosphere is mostly  $O_2$  from life, but early Earth was N.
- Earth and Venus have similar amounts of carbon &

nitrogen, but Venus's carbon is loose in the atmosphere.



http://www.digitalart.ab.ca/art/ren/images/birth-of-venus.jpg

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What Happened to Venus?



- Apparently Venus lost its H<sub>2</sub>O– no oceans and no sediments.
- Probably the atmospheric temperature was hot enough for water to travel high enough to be broken apart by UV radiation, the H was lost and the O reacted with something else.
- Irreversible procedure!
- Which is why greenhouse effect is worrisome here too!
- The Earth traps water vapor in the cool tropopause at 14km.

http://photos1.blogger.com/blogger/4103/1148/1600/Venus%20Wimbeldon05.jpg



## Life on Venus?

- Surface is far too hot
  - If lead is liquid, think of what heat would do to complex organic polymers
  - No cooler polar regions exist
    - Heat is uniform!
    - But, high in the clouds it should be cooler?! Temperature (°C)

80

Altitude (km) 05 09

- Maybe life can still exist in the clouds?
- At 50 km up, the temperature is not too hot and the pressure is 1 atmosphere.



# Life on Venus?

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0.0001

0.001

0.01

Pressure (atm)

Haze lave

Upper cloud laye

Middle cloud lave

ower cloud lave

400 500 600 700

Temperature (K)

- One possibility is that microbes living in the clouds could be combining sulfur dioxide with carbon monoxide and possibly hydrogen sulphide or carbonyl sulphide in a metabolism similar to that of some early terrestrial microorganisms.
- Given that the temperature on Venus was once much cooler, there may once have been oceans on the planet. Life could have started there and retreated to stable niches once the runaway greenhouse effect began.

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• Maybe a mission to scoop up some atmosphere?



## **Chemical Disequilibrium**

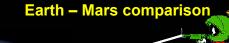
- High clouds in the atmosphere contain chemicals that <u>hint</u> at the presence of some kind of biological activity.
- Hydrogen sulfide and sulfur dioxide two gases that react with each other- exists in the clouds.
- Something is probably producing them.
- Hardly any carbon monoxide. So something is perhaps removing the gas.



http://www.manson-valley.de/fotogalerie/manson/images/acss/acss\_32.jpg

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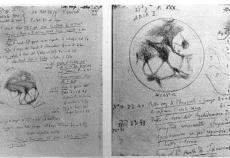
Mars has the Solar System's largest Volcano, Olympus Mons – 27 km tall. Radius Surface gravity Mass Distance from Sun Average Temp Max Temp Year Length of Day Atmosphere 0.53 Earth 0.38 Earth 0.11 Earth 1.5 AU -63 C 20 C 687 Earth days 24 hours 39 minutes  $CO_2$  95%



### What we used to think.

- Was thought to be similar to the Earth in many ways.
- Life was argued to exist on Mars by many astronomers.
- The astronomer Schiaparelli announced that he saw regular linear markings on the surface, which he named canali.
- Technically, in Italian means channels, but it was mistranslated to canals.





Pages from Schiaparelli's observing notebook, 1879

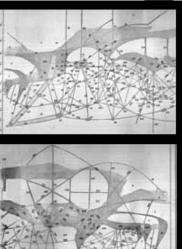
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## **Percival Lowell's Canals**

- Evidence for intelligent life?
- Mapped the civilization.
- Influenced culture.







Martian "canals" as mapped by

Percival Lowell in the late 1800s.

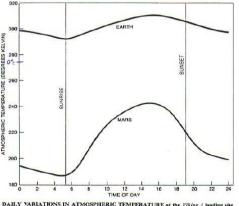
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## The Martian Atmosphere

- 95% carbon dioxide
- Atmospheric pressure 0.6% of Earth's – like 40 km altitude on Earth
- But too thin for significant greenhouse effect.
- Pressure is too low for liquid water.
- Not protected by a global magnetosphere like Earth's



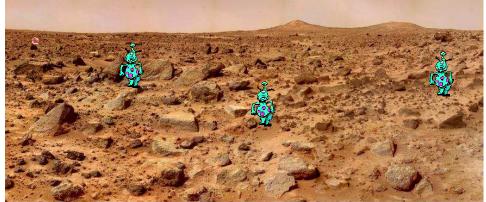
• Large daily and seasonal swings in surface temperature



are qualitatively similar to those at China Lake, Calif., a desert site (*bl*) the temperature touches a minimum around sunrise and reaches a peak ab ter. The daily range, however, is about three times greater on Mars than it is on the earth. At te range is 55 degrees, from about 187 to 242 degrees Kelvin (-86 to China Lake mage is 18 degrees, from 292 to 310 degrees K. (19

#### **The Surface of Mars**

- Mars is a desert!
- Iron oxide in soil gives reddish cast.



View of "Twin Peaks" from Mars Pathfinder

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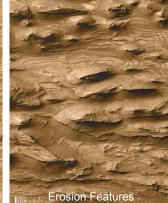
Astronomy 230 Fall 2006 http://www.grc.nasa.gov/WWW/PAO/html/marspath.htm

#### Liquid water on Mars?



- Water erosion features visible from space
- Atmospheric pressure too low for liquid water to exist
- Perhaps at some point in the past?





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"Islands"



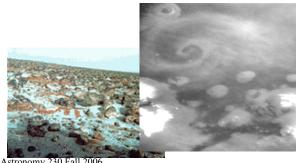
Flood erosion

### Water on Mars

- There is water on Mars
  - North and south polar caps (mostly CO<sub>2</sub>)
  - Some water vapor in the air
  - Frost on rocks
  - Clouds (ice crystals)
- No *liquid* water now





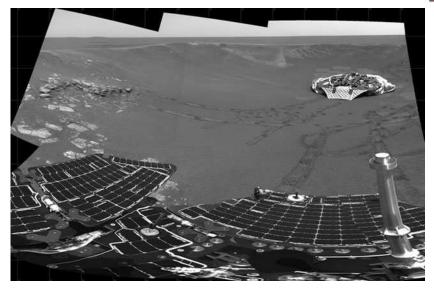


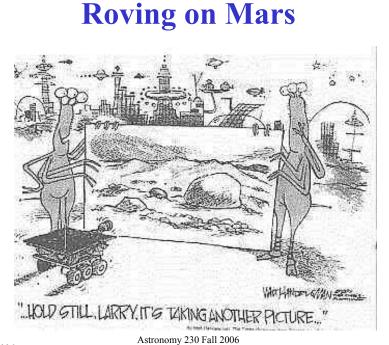
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## The Surface of Mars: Opportunity





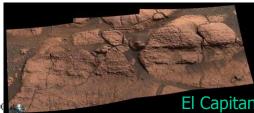


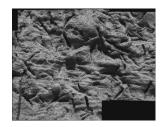
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## **Standing Water on Mars**



- The new data from the rovers are highly suggestive of ancient standing water on the Meridiani Planum.
- 3 pieces of evidence:
  - Physical appearance of rocks
  - Rocks with niches where crystals appear to have grown
  - Rocks with sulfates left after the water evaporated
- Is it a former sea floor or just an area that had ground-water?







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Roving on Mars:

Spirit and Opportunity find evidence of ancient

liquid water

http://antwrp.gsfc.nasa.gov/apod/image/0403/emptynest\_

opportunity\_big.jpg

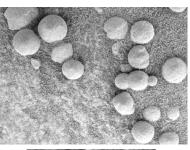


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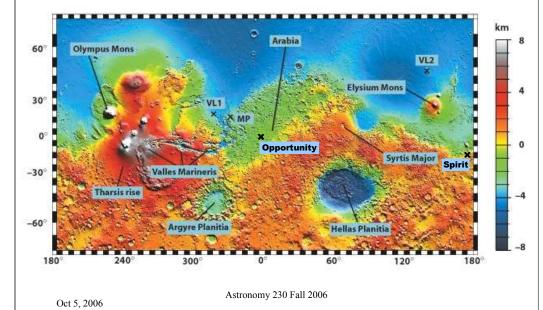
## What Happened to the Water?

- That is the big question
  - Both Opportunity and Spirit have found evidence of water
  - Did the water escape to space with the air?
  - Is it frozen beneath the surface?
- The rovers are continuing their exploration
- More missions are planned





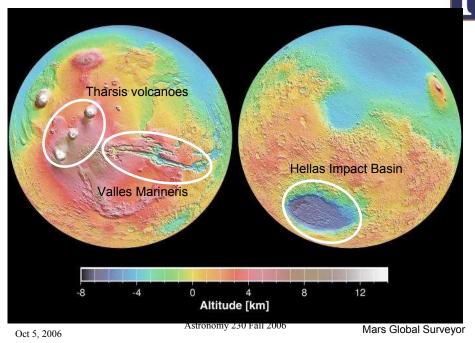






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#### The Surface of Mars



### **Olympus Mons**

- The largest mountain in the Solar System rising 26 km high
- A shield volcano, like Hawaii on Earth
- Its caldera is 90 km across

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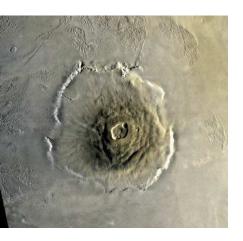




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## **Olympus Mons**

- Its base is more than 500 km in diameter
- As long as the entire Hawaiian island chain
- Rimmed by a 6 km high cliff
- Last erupted 25 million years ago
- Probably so big, due to lack of plate tectonics

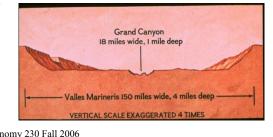


#### **Valles Marineris**

- A series of fault canyons
- 5000 km long
  - A big as the U.S.!
- A giant crack in the crust of Mars
  - Formed as the planet cooled
  - Expanded by water flow





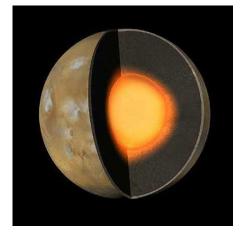


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## **Mars' Interior**

- Like Earth, Mars has an iron core
  - About half of the planet's radius in size
  - Heavily contaminated with sulfur
  - Weak magnetic field suggests a thin layer of liquid iron, mostly solid



### Mars' Past

- Early in Mars' history it was likely more Earth-like
  - Geologically active
  - Volcanic eruptions created a thick carbon dioxide, nitrogen atmosphere
  - Greenhouse effect made it warm enough for liquid water
  - Oceans? Rivers? Glaciers by the poles?
  - Life?



## What Happened?

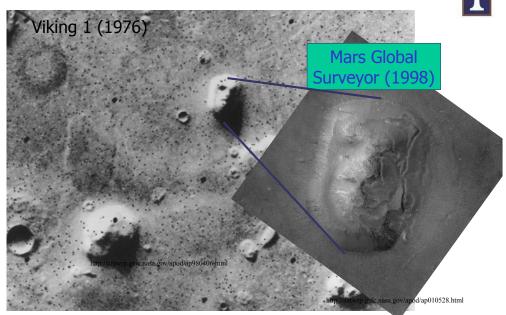
- Mars was too small
  Not enough internal heat
- Plate tectonics stopped
  - Volcanoes sat over "hotspots" grew to immense sizes
- Volcanic activity slowed as the interior cooled
- The atmosphere escaped
- The planet froze

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• Did life move underground?



#### The "Face" of Mars?



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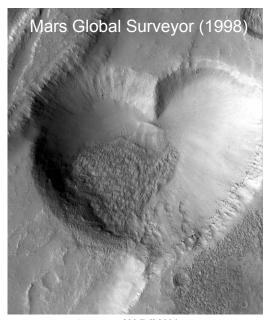
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#### **Other Faces**

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#### **Other Places**





## **The Search for Mars Life**

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- Viking 1 and 2 carried several experiments to detect life
- The results were ambiguous. The soil reacted vigorously with the Viking nutrients, then tapered off in activity.
- The conclusion of most scientists is that the reactions were due to inorganic chemical reactions.



#### **Martians?**

- In August 1996, evidence for microbial life was found in a Martian meteorite.
  - ALH84001 (3Gyrs): Found in Antarctica, composition suggests it was knocked from Mars
  - About 14 such Mars rocks have been found on Earth
- David McKay *et al.* suggested that there was fossil evidence for bacteria in the meteorite.



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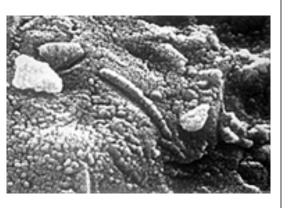


# **Martian Microbe Fossils?**

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- Microscopic shapes that resemble living and fossil bacteria on Earth– nanobacteria, but much smaller than on Earth.
- Microscopic mineral grains like some produced by living and fossil bacteria on Earth
- Organic chemical compounds that resemble the decay products of bacteria on Earth.
- In the end, not impelling enough. Non-biological processes can probably produce the observed features



## Phobos & Deimos

- Mars' moons
- Likely captured asteroids
- Very small
  - About 15-25 km in size
  - Shaped like potatoes







## **Mars Missions**

- Mars Reconnaissance Orbiter (now)
  - Will study the geology and climate of Mars
  - Look for ancient sea shores
  - Survey potential landing sites
- Phoenix (2007)
  - Will analyze water ice at Mars' north pole



## **Manned Mars Exploration**

- NASA's plans to send a manned expedition to Mars
- Timetable:
  - Complete Space Station by 2010
  - Return to Moon by 2020
  - Then, on to Mars (no date)
- No cost estimates
  - Some funds from to-be-retired shuttle fleet



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