### ET: Astronomy 230 **Section 1– MWF 1400-1450**

134 Astronomy Building

This Class (Lecture 8):

Planet Formation

Next Class:

Sept 12, 2005

Nature of Solar Systems

HW #2 is due on Friday

First Presentations on

19th and 23rd

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#### Outline



- From molecular clouds to baby stars—protostars.
- Star formation requires a circumstellar disk that is often seen around young stars.
- The origin of the Solar system also requires a disk of material in which dust clumped, forming planetesimals, then planets.
- Planets are different due to distance away from Sun..

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### **Presentations**



- Will be treated like a real scientific talk at a meeting.
- I will keep you to 10 minutes with 5 minutes of questions.
- Any speculative claims MUST have a scientific reference source.
  - Can't just claim that monkeys live on the Moon.



#### **Presentations**

- Can give presentation in any format you want.
- Last semester:

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- 92% powerpoint
- 5% talking with pics from webpages
- 1% dedicated webpage
- 2% overhead slides
- If presentation is electronic
  - Email me evening before
  - Or, on netfiles, email me URL location **BEFORE class**
  - Or, bring in burned CD (present to me BEFORE class)



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#### Oral Presentation



- 1. How relevant is the topic to the search for extraterrestrial life?
- 2. How interesting is the topic for the general class audience?
- 3. Rate the extent of the speakers knowledge on the topic?
- 4. Rate the quality of the overall presentation?
- 5. Does the research have a solid scientific basis?

These questions are rated 1-10 out of 10 scale by your peers!

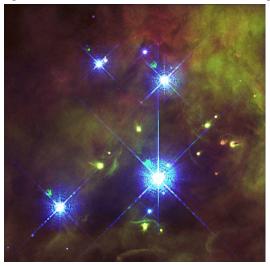
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#### The Birthplace of Stars



- Young stars often are seen in clusters
- Very young stars are also associated with clouds of gas (nebulae)



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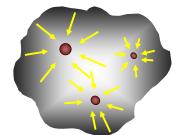
## **Gravitational Contraction**



- As we discussed for the first stars, the gravity of the gas and dust clumps push the clumps together, but there is some resistance from pressure and magnetic fields to collapse.
- Probably as the cloud core collapses, it fragments into blobs that collapse into individual stars.
- Cloud becomes denser and denser until gravity wins, and the clumps collapse under their own mass—a protostar.
- This process is slower than for the first stars where the clumps were much more massive.
- As the collapse proceeds the molecules and dust emit light, keeping the temperature of the core low.

### **Cloud Contraction**





http://www.birthingthefuture.com/AllAboutBirth/americanway.php

#### But.



- Not all mass falls in directly (radially). Why?
- All gas has a small spin that preferentially causes the formation of a flattened structure - time for an interlude.

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## Interlude: Angular Momentum



Spinning or orbiting objects in closed system have angular momentum.

Angular momentum is a single, *constant* number = *conserved*!

Keep same dist. to axis velocity same

Move closer to axis speed up!

Recall Kepler 2<sup>nd</sup> law – really due to angular momentum!

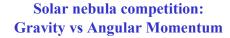


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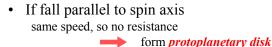
## When Doves Cry and Stars Form

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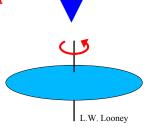








- Origin of planet's orbits!
- Organizes spins along initial spin axis

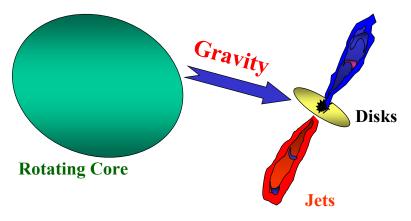


## The Protostar Stage

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## Gravity, Spin, & Magnetic Fields

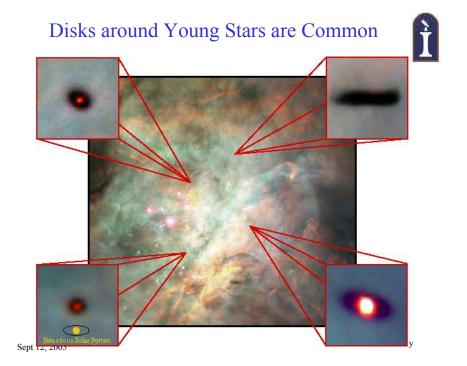


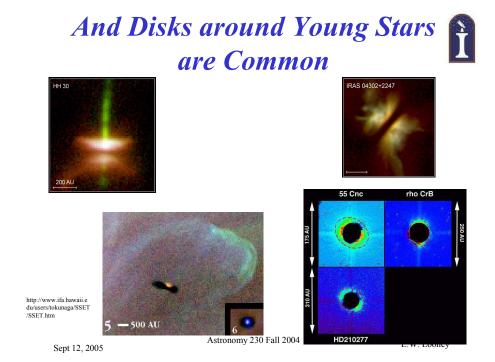
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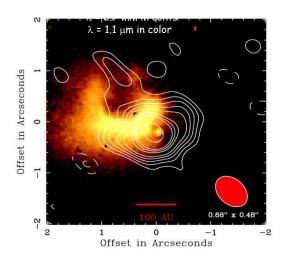
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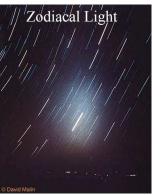
## The Circumstellar Disk of HL Tauri



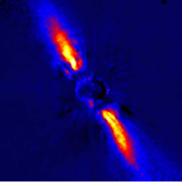
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## Do Fossil Disks Exist around other Stars?

• We see old disks around other stars (e.g. Vega and Beta Pictoris) as well as our own.



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http://www.eso.org/outreach/press-rel/pr-1997/phot-16-97.html http://antwrp.gsfc.nasa.gov/apod/ap970826.html

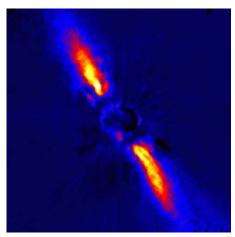
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## Disks Around Young Stars

- Many (> 50%) of newborn stars surrounded by a disk of material!
- Disks thick, blocks light
  - > Enough material to make planets
  - > Agrees with Solar Nebula theory!

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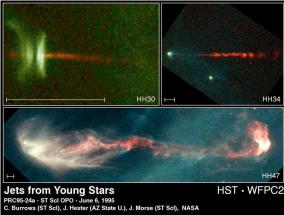


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#### **Protostellar Jets**

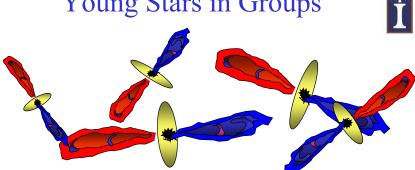




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- Most stars are in multiple systems.
- How does this effect the protostars?
- How does this effect their planet formation?
- How does this effect the possibility of life on the average star? L.W. Looney



## The Cone Nebula

A Star Forming Region



## The Protostar Archetype: T Tauri





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http://www.astrosurf.com/iw/h/ttluggneV

## On to the Main Sequence: A Star is Born!





- Density increase, temperature increases until fusion can occur.
  - Blows away most of its natal circumstellar material.
  - Becomes a star on the main sequence of the HR diagram,
  - For low mass stars, this whole process can take 10<sup>6</sup> years.
  - Expect to see a large number of embedded protostars.



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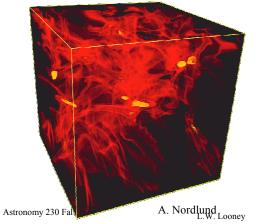
#### **Star Formation - Summary** Young stellar object with bipolar outflow Giant molecular cloud Dust-shrouded core Age $\sim 5 \times 10^{5} \text{ yr}$ Age $\sim 10^5 \text{ yr}$ Protoplanetary disk? Main-sequence star Magnetically active Age $10^7 - 10^8$ yr protostar (T Tauri star) Hydrogen fusion powered Age $\sim 5 \times 10^{6} \text{ yr}$ Creates emission or reflection nebula Gravitational collapse Inhibits / stimulates further star form. powered Sept 12, 2005

#### Some outstanding Star Formation Issues



- Why do the cores collapse, but not the entire molecular cloud?
- What sets the sizes of cores, and hence masses of stars?
- What determines how stars cluster, group together, or form

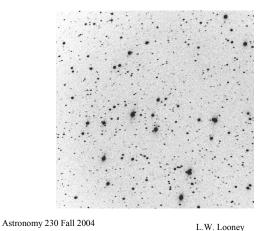
multiple systems?



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## Stars Ages and ETs





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## So, Why would Spock Care?



- If we are to suppose that ET life will be based on a planet orbiting a star, then we need to know
  - How did our solar system form?
  - How rare is it?
  - Is our solar system unusual?



mon/images/astro%05/4pocking

## What is the origin of the Solar System?



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- Explain present-day Solar System data.
- Predict results of new Solar System data.
- Should explain and predict data from other stars!

What are clues to solar system origins?

## **Some Facts of the Solar System**



- We have 8 or 9 planets.
  - Pluto, an anomaly in many ways, probably a Kuiper object or moon of Neptune. Other Kuiper objects are being found.
- So perhaps the average extrasolar system has about 10 planets (rounded off).
- The Sun has 99.9% of the mass, but the planets have 98% of the angular momentum (energy stored in orbits)

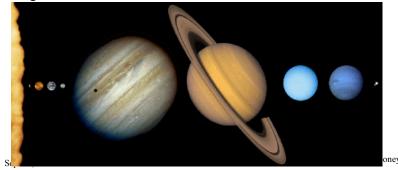


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## Some Facts of the Solar System



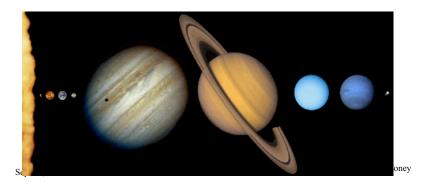
- Outer planets more massive than inner planets.
- Most of the motions in the Solar System are counter clockwise (problems with Venus, Uranus, or Pluto) in a flat system (pancake-like).
- The inner planets are rocky and the outer planets are gaseous.



## Some Facts of the Solar System



- Numerous collisions occurred in the early Solar System
  - Origin of Moon, Lunar craters, Uranus's obit, and Pluto
- Planets are not evenly spaced—factors of 1.5 to 2.
  - Sun/Saturn distance is 2x Sun/Jupiter distance
  - Sun/Mars distance is 1.5x Sun/Earth distance



Data: Planet's Dance



http://janus.astro.umd.edu/javadir/orbits/ssv.html

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# Data: The Structure of the Solar System

• What are the furthermost solar system objects from the sun and what is their distribution?

Icy objects or long period comets

Furthermost objects form the Oort cloud! So...Spherical Geometry.

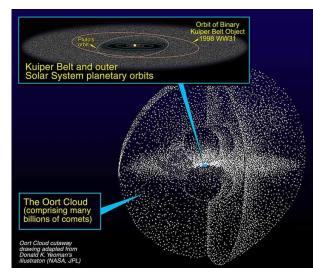
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## Data: Kuiper Belt





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## Data: What is the age of the Solar System?



- Earth: oldest rocks are 4.4 billion yrs
- Moon: oldest rocks are 4.5 billion yrs
- Mars: oldest rocks are 4.5 billion yrs
- Meteorites: oldest are 4.6 billion yrs
- Sun: models estimate an age of 4.5 billion yrs

## Age of Solar System is probably around 4.6 billion years old

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## Origin of Solar System: Solar Nebula Theory



#### **Gravitational Collapse**

• The basic idea was put forth by Immanuel Kant (the philosopher)— Solar System came from a Gas Nebula.

- 4.6 billion years ago: a slowly spinning ball of gas, dust, and ice with a composition of mostly hydrogen and helium formed the early Solar System.
- This matches nearly exactly with the idea of star formation developed last class.

"nebula" = cloud



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## The Early Solar System

- A massive cloud of gas and dust
  - Seeded with elements from
    - Big Bang (hydrogen, helium, etc.)
    - Elements from planetary nebula pushed into space by red giant.
    - Elements blown from across galaxy by supernovae.

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**Planet Formation in the Disk** 

The cloud collapsed under its gravity and formed the circumstellar disk from which our solar system formed. Most theories for solar system formation require disks with masses of 0.01 to 1 solar masses.



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#### Heavy elements clump

- 1. Dust grains collide, stick, and form planetesimals—about 10<sup>12</sup> of them, sort of like asteroids! All orbit in the same direction and in the same plane.
- 2. Gravity Effects: Big planetesimals attract the smaller planetesimals. So, fewer and fewer of large objects (100's). Collisions build-up inner planets and outer planet cores.
- 3. Collisions can also account for odd motions of Venus (backwards), Uranus (rotates on its side), and Pluto (high inclination of orbit). Proof of period of high collision evident on moon



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## Everyone Loves Disks



- As the star forms, the inner region of the disk gets much hotter than the outer regions, creating a temperature gradient.
- The inner part of the disk had a higher density than the outer regions.

• Icy mantles of dust grains (NH<sub>3</sub>, CH<sub>4</sub>, etc.) evaporated at varying distances.

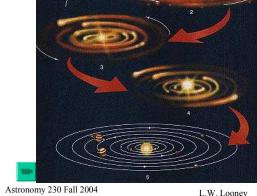


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## What it might have looked like.







## What Are We Looking For?



**General Predictions of Solar Nebula Theory** 

- Are interstellar dust clouds common? **Yes!**
- Do young stars have disks? **Yes!**
- Are the smaller planets near the star?
- Are massive planets farther away?

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Planets around other stars

Would our solar system nebula formation theory account for other solar systems around other



**Drake Equation** 





















=  $R_* \times f_p \times n_e \times f_l \times f_i \times f_c \times L$ 

advanced civilizations we can contact

Rate of star formation Fraction of stars with planets

Earthlike planets per system

Fraction Fraction on which that evolve life arises intelligence

that commun-

Lifetime of civilizations

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Test of Exoplanets



= extrasolar planets = "exoplanets"

stars?

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