

Astronomy 122



This Class (Lecture 7):

Telescopes and Light

Next Class:

How to do a jump shot

Homework #3 due Sun!

Music: Kelly Watch the Stars – Air

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Night Observing



- Night observing starts next week!
 - Feb 11-14th: Monday-Thursday
 - Feb 20th: Wednesday (special Lunar Eclipse!)
 - Feb 25-28th: Monday-Thursday
- Don't wait until last minute (never know about Illinois weather)!
- Observing sessions are from 7:30pm-9:30pm (allow 45 mins to complete)



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Night Observing



- Sign up for sessions at <http://www.astro.uiuc.edu/classes/nightobs/>
- Check weather status before you go at <http://www.astro.uiuc.edu/classes/nightobs/status.php>
- Download Astro 122 worksheet at <http://eeyore.astro.uiuc.edu/~lwl/classes/astro122/spring08/hw.html>

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Outline



- Telescopes
 - Refractive
 - Reflective
- Light



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Astronomy as a Hobby



- Did you know you can see a galaxy 2½ million light-years away with your unaided eyes?
- Or that you can see craters on the Moon with binoculars?



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Your First Steps...



- **Read**
 - The night sky is beautiful to behold, but astronomy is a *learning* hobby
 - You can find good guides to the night sky at your local library or bookstore
 - Get a copy of *Sky & Telescope* from the library
 - Offers a big evening-sky map for beginners
 - Practical observing tips



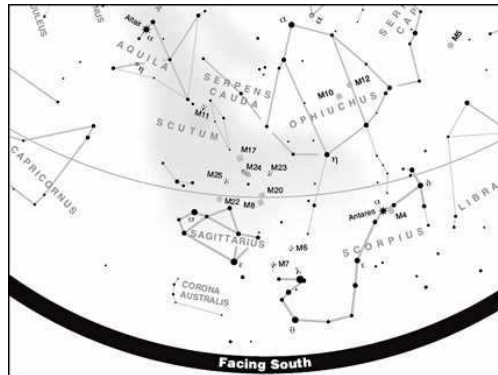
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Learn The Sky



- **Learn the sky with the naked eye**
 - Download star charts from *Sky & Telescope*
 - Use your planisphere from the book
 - Generate sky charts with the Starry Night software that came with your textbook



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Start With Binoculars



- Binoculars are an ideal first telescope
 - Wide field of view, making it easy to find your way around
 - Relatively inexpensive
 - Widely available
 - Easy to carry and store
 - Allow you to easily see lunar craters, Jupiter's moons, and the brighter star clusters, galaxies, and nebulae
 - The larger the front lenses are the better



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Seek Out Others



- There are two amateur astronomy clubs here
 - University of Illinois Astronomical Society
 - Champaign-Urbana Astronomical Society
- Attend star parties where you can meet members and discuss astronomy
 - Try out different types of telescopes
 - Get advice



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Your Own Telescope



- When you're ready, it's time for your own telescope
- Don't skimp on quality, you'll regret it later
- What do you want?
 - Solid, steady, smoothly working mount
 - High quality optics
 - Large aperture – but not too large, you have to carry it!
 - The best telescope for you *is the one you'll use most!*



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Relax and Have Fun!



- This is the most important step!
- Take pleasure in whatever your eyes, binoculars, or telescope can show you
- The more you look, the more you will see, and the more you will become at home in the night sky
- Set your own pace, and revel in the beauty and mystery of our amazing universe!



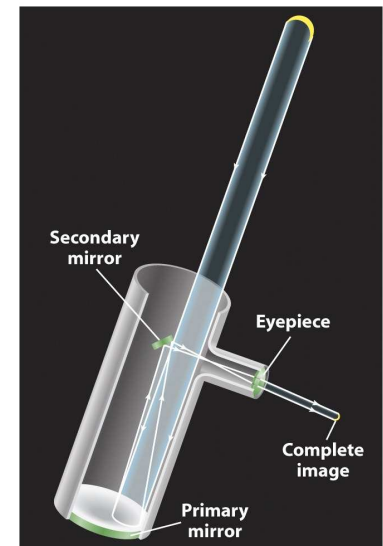
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Functions of a Telescope



- Telescope functions
 - Collect light over a large area
 - Resolve image onto an eyepiece or a scientific instrument
- Can do this with either lenses (refracting) or mirrors (reflecting)
- Three priorities (in order)
 1. **Gathering light**
 2. **Angular resolution**
 3. **Magnification**



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First, Light Gathering



- Top priority since most celestial objects are dim
- Telescope = “light bucket”
- Key: collecting area
- Human eye – ~5 mm,
- Subaru telescope mirror – 8.3 m
 - 3 million times the area of your eye!



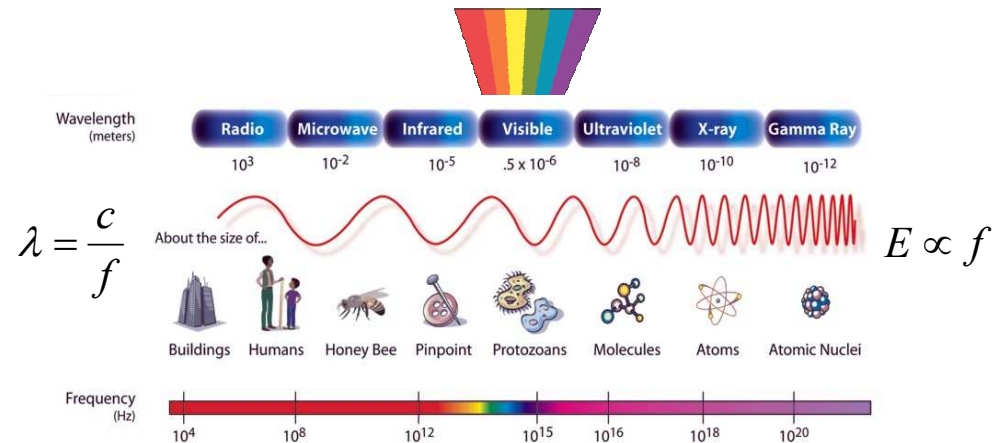
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The electromagnetic spectrum



- Visible light is only a tiny portion of the full electromagnetic spectrum
- Light comes in many colors that you can not see! The color x-ray or color radio or color microwave.
- Divisions between regions are really only from biology or technologies.



Question



If I want to contact an alien civilization near Orion (1500 light years away), which of the following will give the fastest communication?

- a) Gamma-rays
- b) Red light
- c) X-rays
- d) Radio waves
- e) All of the above travels at the speed of light.

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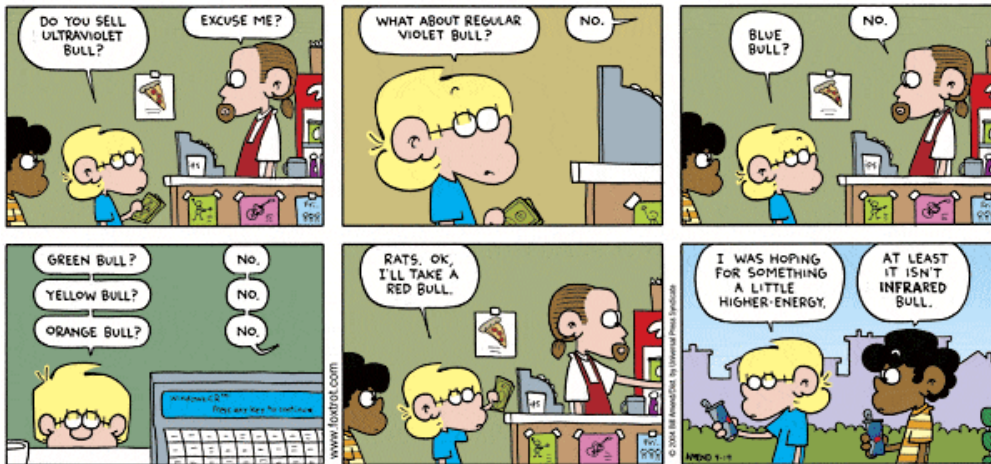
Where Do Colors Come From?



- The color of light is determined by its wavelength
- Visible light has extremely small wavelengths
 - Wavelengths range from 400 nm (violet light) to 700 nm (red light)
- Colors, from longest wavelength to shortest: red, orange, yellow, green, blue, violet
- But just as valid, x-ray is a color, just higher energy.

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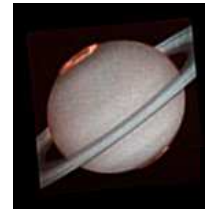
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Saturn at Multiple Wavelengths



Ultraviolet



Visible



Infrared



Radio

<http://www.ipac.caltech.edu/Outreach/Multiwave/gallery.html>

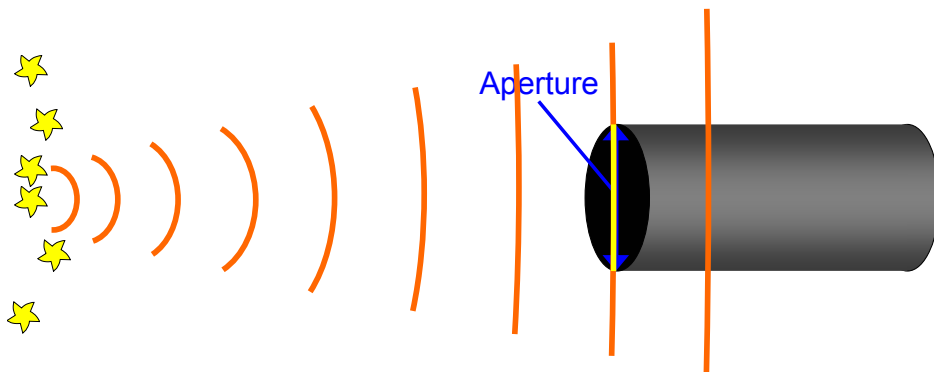
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Telescopes



- A **telescope** collects light
- The larger the **aperture**, the more light can be collected in a given amount of time



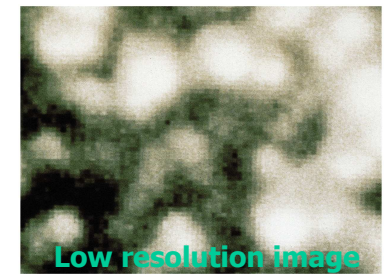
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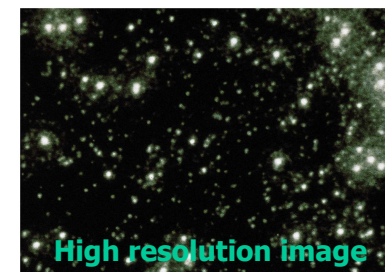
Secondly, Angular Resolution



- Reveal details of objects
- Angular resolution:
 - Measures finest detail that is not smeared out
 - Smallest angle for which two stars aren't smeared together to one
 - e.g., human eye resolution = $1/60^{\text{th}}$ of a degree
 - Hubble Space Telescope resolution $< 1/36,000^{\text{th}}$ of a degree



Low resolution image

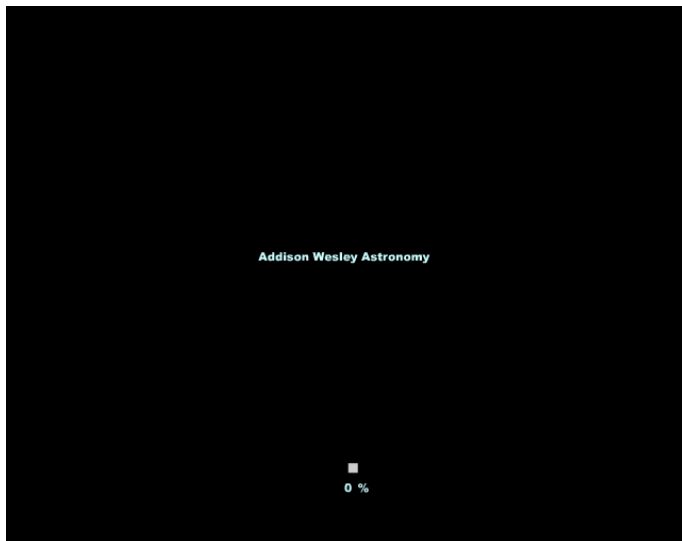


High resolution image

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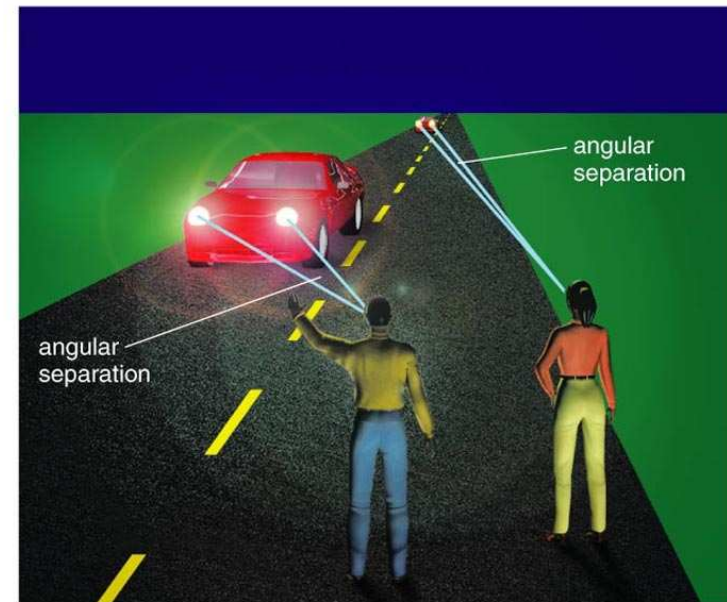
Example of Angular Resolution



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Example of Angular Resolution



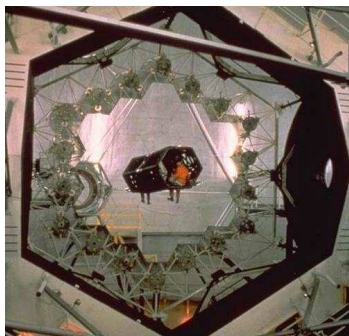
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Resolve This



- What is the limitation on how well a telescope can resolve objects?
 - The size of the telescope, silly



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Resolve This



- The best resolution of a telescope is

$$\theta_{\text{diff}} = 2.5 \times 10^5 \lambda / D$$

λ and D in meters, then θ in arcseconds

 - We want the angle to be small as possible
 - So, again we want a large telescope!
 - The Keck 10 meter has a $\theta_{\text{diff}} = 0.0125$ arcsec with $\lambda = 500 \text{ nm}$ (5×10^{-7} meters)
- The shorter the wavelength of light the better!

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Resolve This



- The best resolution of a telescope is

$$\theta_{\text{diff}} = 2.5 \times 10^5 \lambda / D$$

λ and D in meters, then θ in arcseconds

- The Keck 10 meter has a $\theta_{\text{diff}} = 0.0125$ arcsec in optical
- A 10 meter radio telescope ($\lambda = 1$ cm), then $\theta_{\text{diff}} = 250$ arcseconds!

– But, there is another limitation!

- The atmosphere

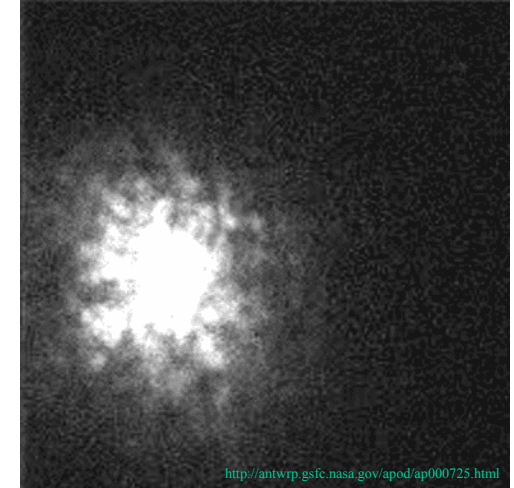
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Twinkle, Twinkle Little Star



- Turbulence in the atmosphere “jiggles” image
- We see it as stars “twinkling”
 - A good atmosphere will allow $\theta \sim 1$ arcsec.
 - A great atmosphere will allow $\theta \sim 0.3$ arcsec.
- So for modern telescopes, we are limited by the atmosphere.



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Lastly, Magnification



- Makes the object appear larger
- Useful for studying detail
- Least important issue
 - If you don't have the other two, this is not at all relevant
 - No good to magnify a blurry image
- Magnification is ratio of focal length of telescope and focal length of eyepiece

$$f_{\text{tel}} / f_{\text{eye}} = M$$



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Question



Which of the following is the most important aspect of a modern research telescope.

- Magnification
- Collecting area
- Resolution
- Expense
- None of the above.

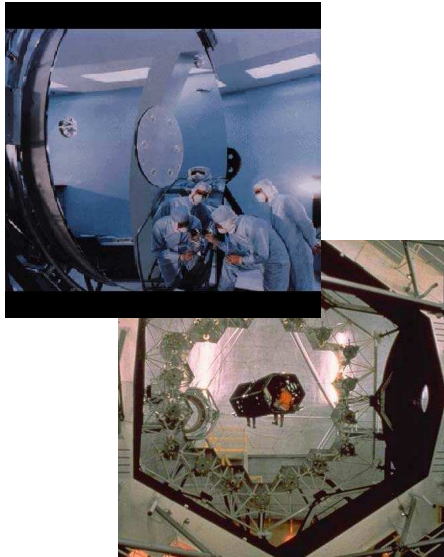
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In The End Size Does Matters



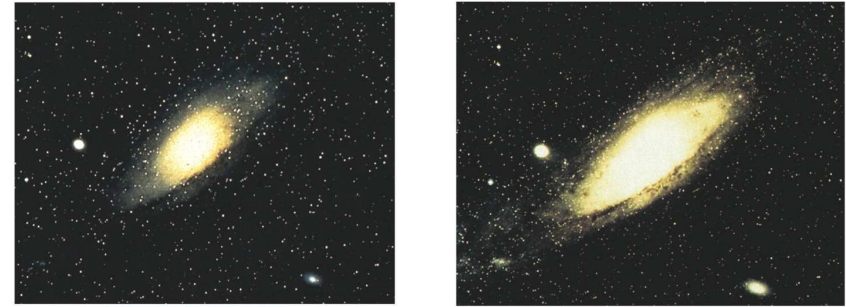
- Both light collecting and resolution improve as the diameter of the scope – its lens or mirror – increases
- Bigger is better!**



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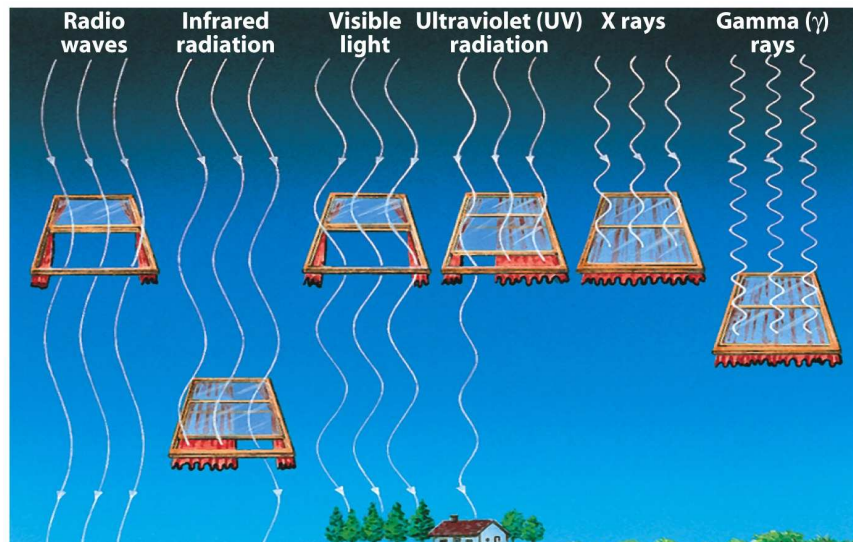
Bigger Is Better!



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The atmosphere absorbs some wavelengths and not others



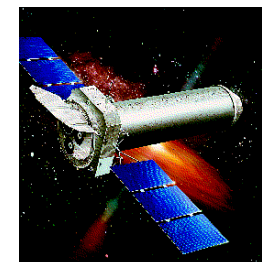
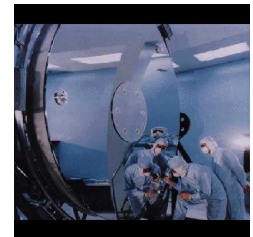
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Telescope Types



- Optical (visible light)
 - Refracting
 - Reflecting
- Radio, infrared, ultraviolet
 - Reflecting
- X-ray
 - Reflecting (grazing incidence)

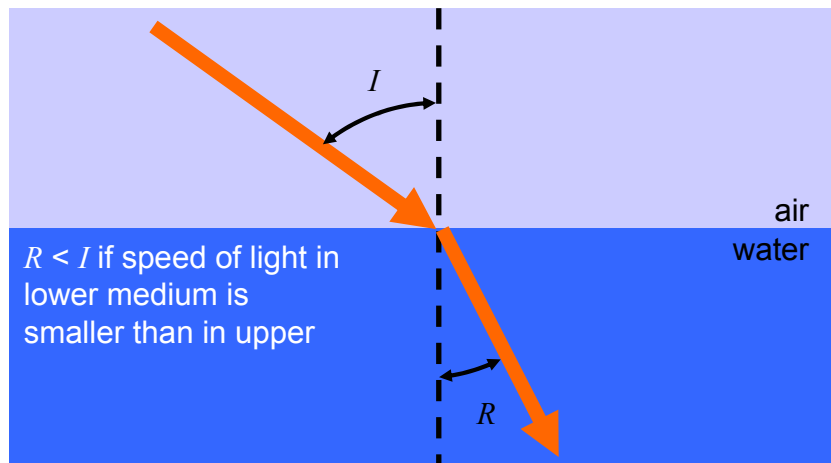


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Refraction



- Light travels more slowly in transparent materials than it does in vacuum
- When passing from one medium to another (e.g. air to water), light is bent (**refracted**)



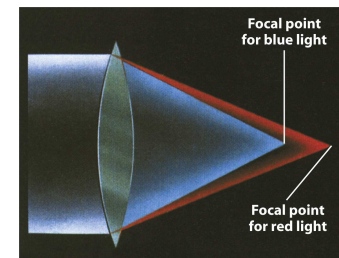
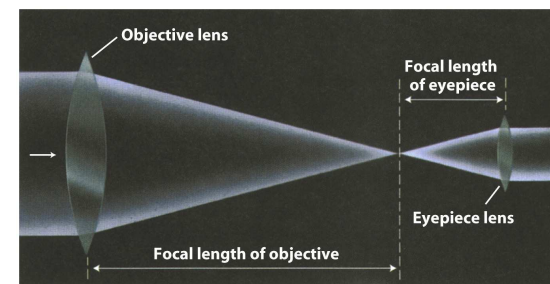
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Refracting Telescopes: Lenses



- Use Lenses
 - Curved glass
 - Light bent to focus
- Problems:
 - Lenses focus colors differently
 - Sag of lens from gravity
 - Large lens distorted as it hangs
 - Limits lens size
 - Limited wavelengths



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The Largest Refractor



- At Yerkes Observatory near Chicago
- 40 inch diameter lens, 63½ feet long!

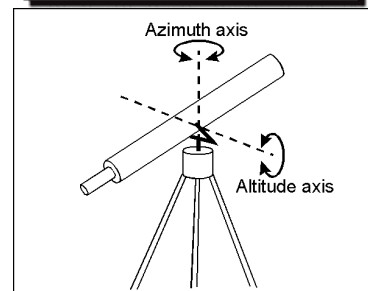
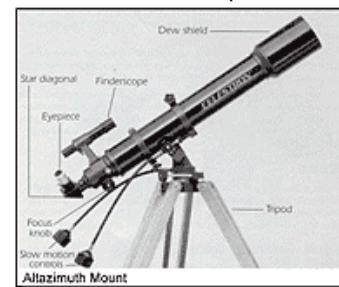


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Telescope mounts



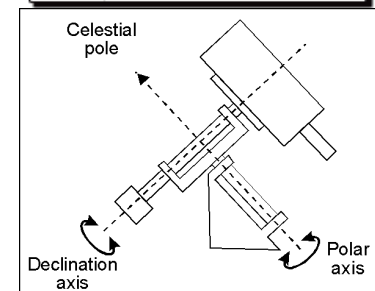
Altitude-azimuth (alt-azimuth)



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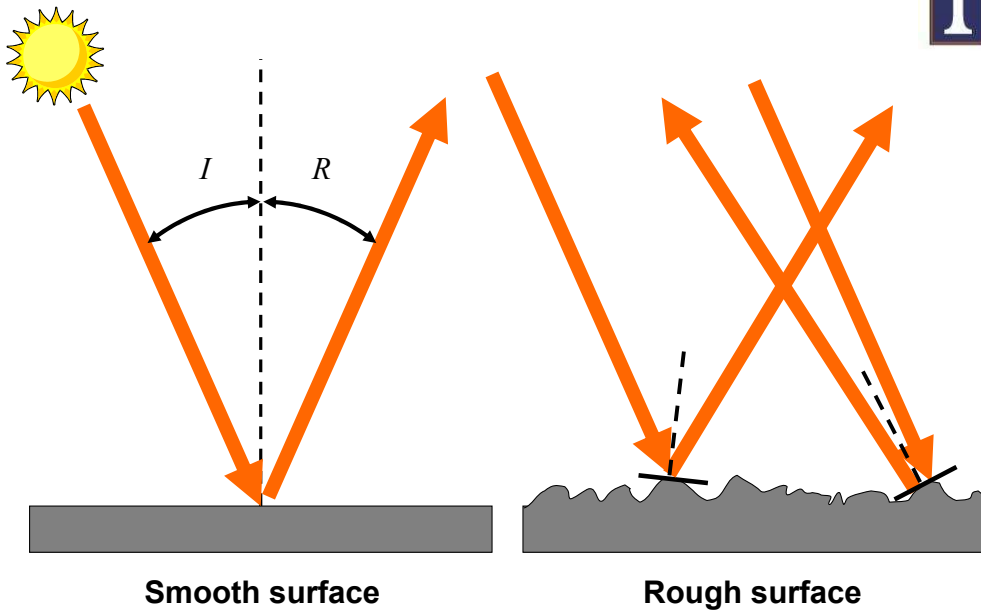
Morehead Planetarium

Equatorial



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Reflection



Smooth surface
Angle of incidence I = Angle of reflection R

Rough surface

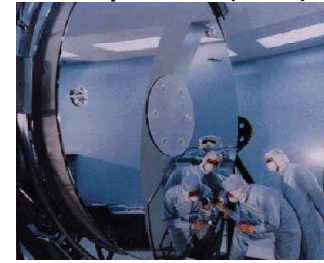
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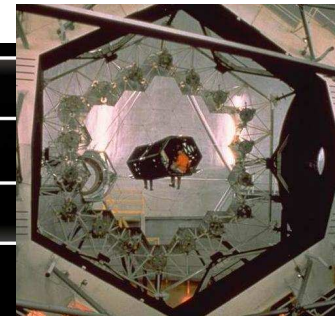
Reflecting Telescopes



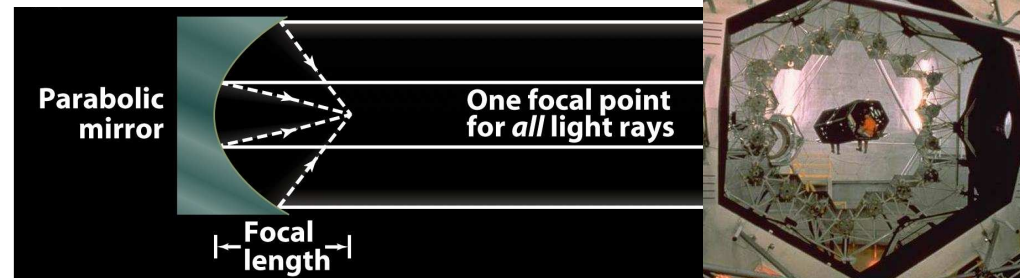
Hubble Telescope Mirror (2.4 m)



Keck Telescope Mirror (10m)



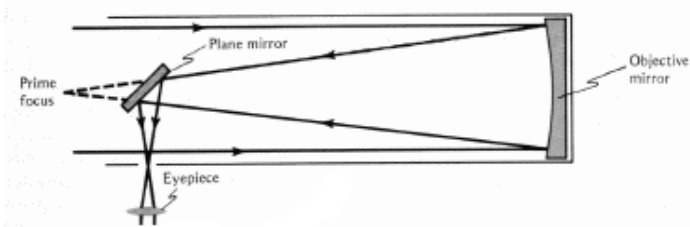
- Use Mirrors
 - Sag not a problem – support the mirror from below
 - Parabolic mirror gives good focus
- Today, all large professional telescopes are reflectors



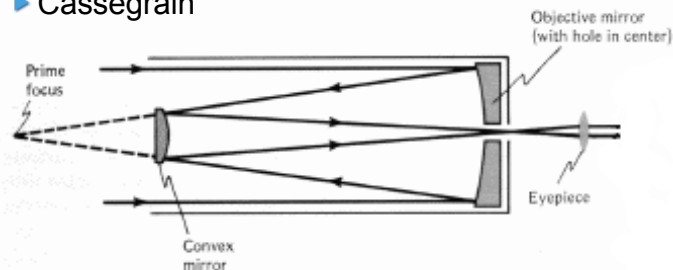
Reflecting telescope designs



► Newtonian



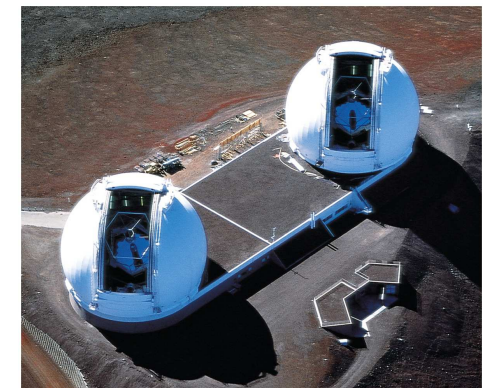
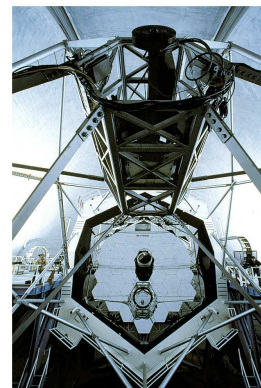
► Cassegrain



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The Largest Reflector



- Keck Observatory on Mauna Kea in Hawai'i
- Twin 10-meter reflecting telescopes
- Completed 1993 & 1996

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Gemini Telescopes



- Twin telescopes
- One in Hawaii, one in Chile
- 8-meter mirrors



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Light Pollution



- Another problem for astronomers is light pollution
- City lights raise the “background light” level
- Makes it more difficult to collect light from stars



Tucson, 1959

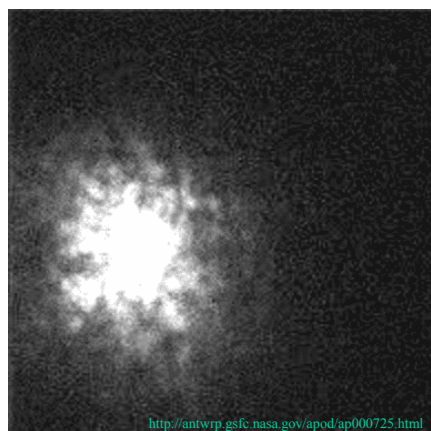


Tucson, 1989

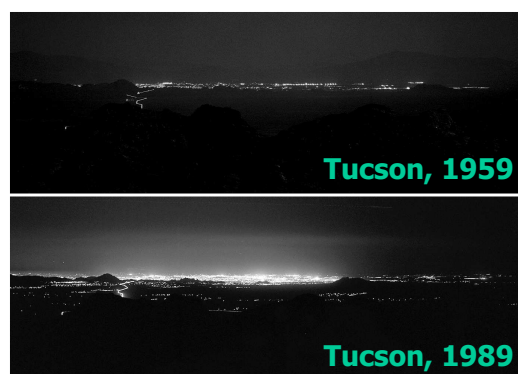
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Twinkling & Light Pollution



<http://antwrp.gsfc.nasa.gov/apod/ap000725.html>



Tucson, 1959

Tucson, 1989

How do we combat these problems?

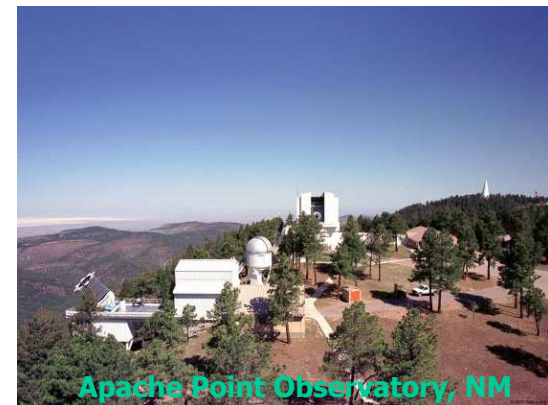
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Remote Mountains



- One solution: Build telescopes at sites high, dry, and away from civilization
- While this solves the scientific problems, it introduces its own complications
 - Providing facilities
 - Environmental impact
 - Cultural conflicts



Apache Point Observatory, NM

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Mauna Kea, Hawai'i



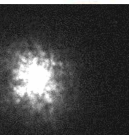
- Mauna Kea is the best place on Earth for astronomical telescopes
 - High elevation
 - Far from urban lights
 - Reasonably easy access
 - Generally good weather
- Mauna Kea is also a sacred place
- Also an environmentally sensitive area



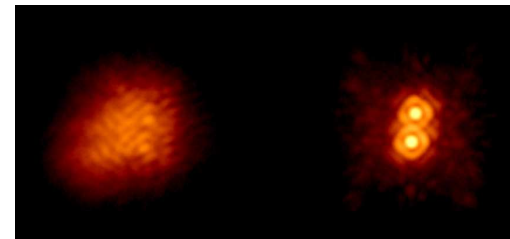
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Adaptive Optics



- If you can observe a star quickly, and change the mirror to correct for the atmosphere you can observe closer to the diffraction limit of the telescope.
- Is working on many telescopes now.
- If no star nearby, can make one with a laser.



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Light Detection



Once light collected and focused need detector

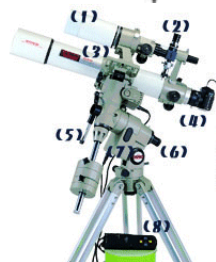
Human eye— just look

- Least sensitive (1% of photons)
- No permanent record
- Only optical wavelengths



Photographic film

- Telescope as camera
- Accumulates light: see dimmer objects
- Provides a permanent record
- Small efficiency (a few % of photons)
- Non-linear response



<http://www.pandia.com/graphics/hemera/womanandchildlooking.jpg>

<http://www.sundu.co.kr/telescope/camera/main.gif>

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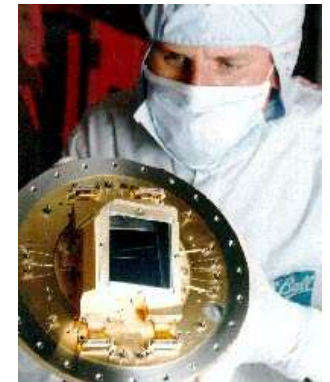
Light Detection



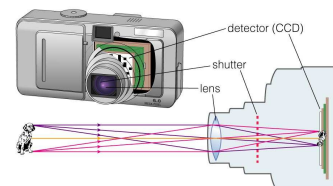
Once light collected and focused need detector

Electronic “film” (CCD)

- Charged Coupled Device
- Like Digital camera/camcorder
- Photons hit silicon chip and electrons kicked-out
- One measures the electrons created in a pixel.



Hubble Telescope CCD



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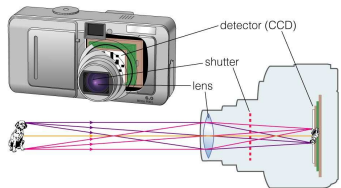
Light Detection



Once light collected and focused need detector

Electronic “film” (CCD)

- About 80% photons detected
- Much more sensitive
- Detector of choice!
- All modern professional astronomy done this way— costly to make large CCDs
- Bonus: digital data great for computers!



Hubble Telescope CCD

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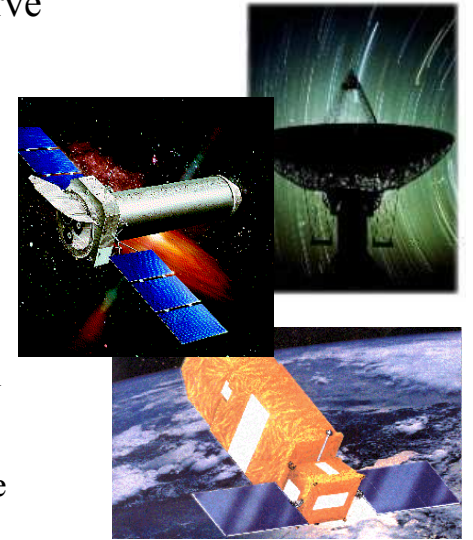
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Invisible Astronomy



- Astronomers want to observe all types of light

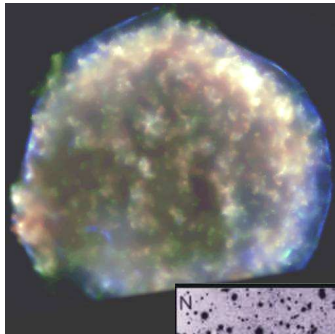
- To see into the dust-enshrouded regions of newly-forming stars
- To peer into the heart of the Milky Way itself
- To study the remains of solar-type stars
- To detect the emission from gases heated to millions of degrees by the powerful explosions of dying massive stars



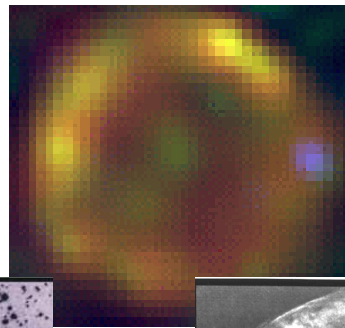
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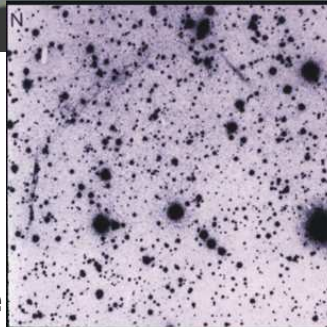
Tycho's Supernova



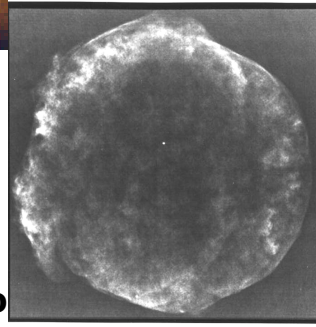
X-rays



Infrared



Visible



Radio

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Radio telescopes



First detection of cosmic radio sources by Karl Jansky at Bell Labs (1932)

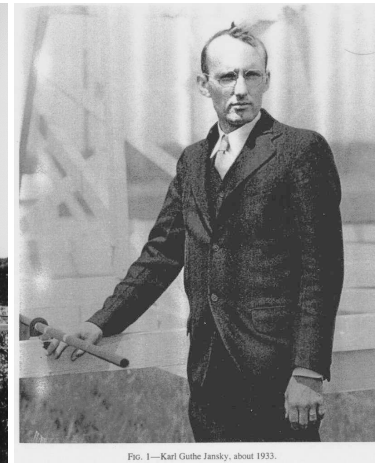
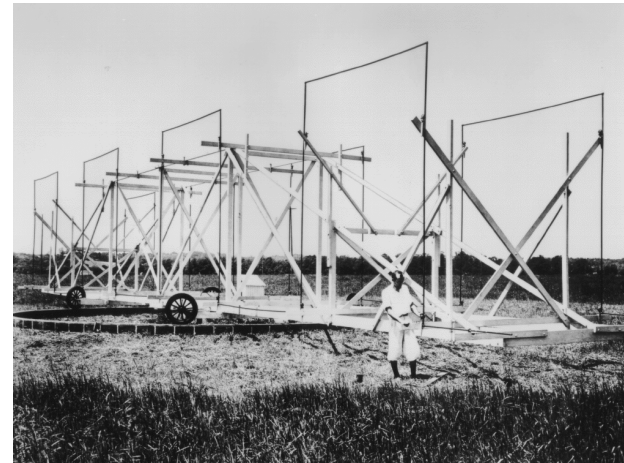


FIG. 1.—Karl Guthe Jansky, about 1933.

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Radio telescopes

Pioneering work by Grote Reber in back yard, Wheaton, Illinois.
(He died in 2002)



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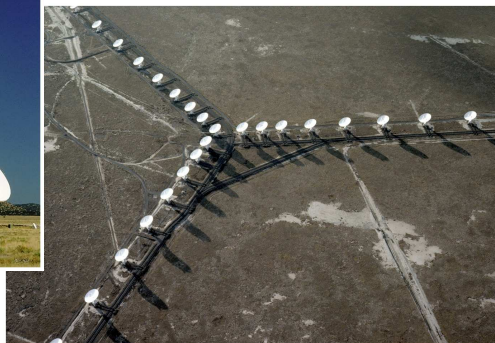
Arecibo Observatory, Puerto Rico



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Very Large Array (VLA), NM



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CARMA



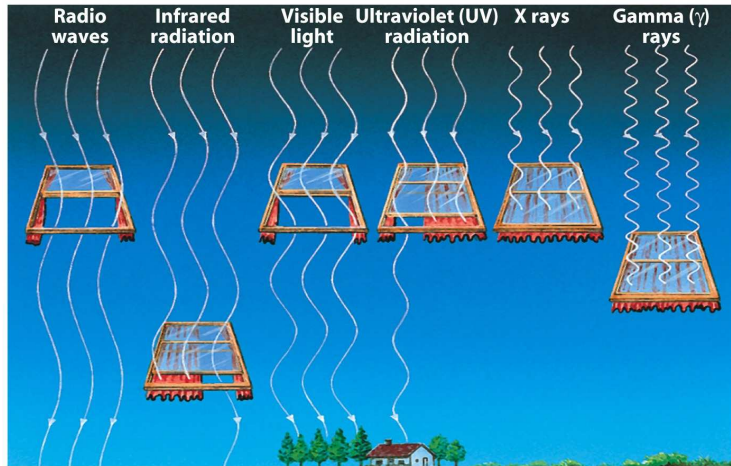
A millimeter array of 15 telescopes (9 six meter and 6 ten meter) owned and operated by CalTech, UC Berkeley, **UIUC**, and UMd in White Mountains, California. Wavelength of 1.4 millimeters – frequency of 220 GHz. Works night and day. Why?



Opaque Atmosphere



- The atmosphere blocks some wavelengths
- Must observe some wavelengths from space!



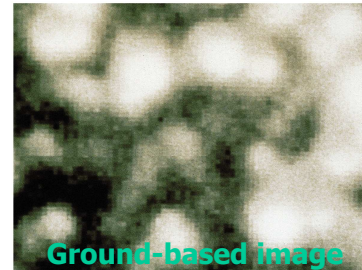
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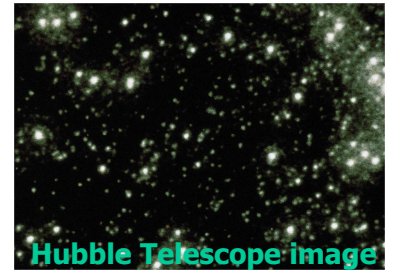
Hubble Space Telescope



- 2.5 meter reflecting telescope in space
- Above the atmosphere
 - No “twinkling” effects
 - No light pollution



Ground-based image



Hubble Telescope image

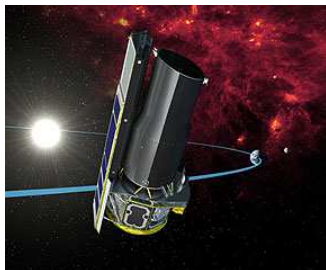
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Spitzer Space Telescope



- 0.85 meter infrared telescope
- Launched August 2003
- Cooled to near absolute zero so that its own heat doesn't confuse the results



Reflection Nebula NGC 7129 Spitzer Space Telescope • IRAC
NASA / JPL-Caltech / S.T. Megeath (Harvard-Smithsonian CfA)

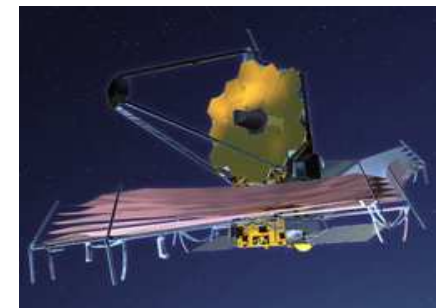
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James Webb Space Telescope



- The next space telescope – 2011
- Observe in the near and mid-infrared
- Will be the biggest telescope in space – 6 meters! (Must fold up for launch)
- Will take 3 months to reach position – no service missions



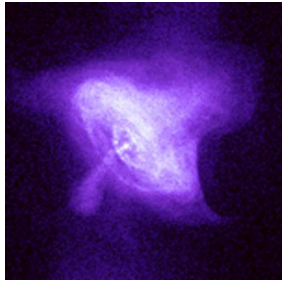
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Chandra X-ray Observatory



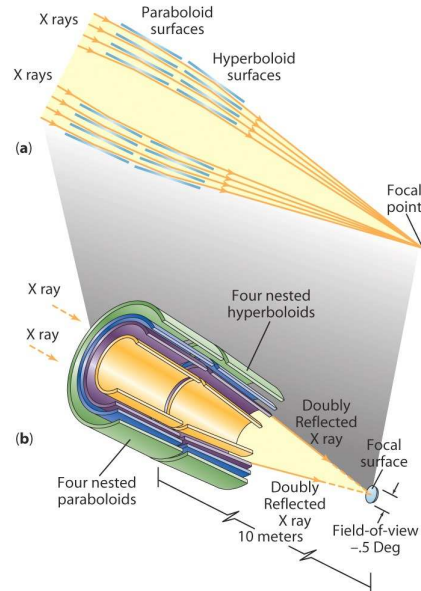
Launched 1999



Crab Nebula in X-rays

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Question



Which of the following is **not** a reason to place telescopes in space?

- a) Can put the biggest telescopes in space.
- b) There is no light pollution.
- c) Can observe at wavelengths that are blocked by the Earth's atmosphere.
- d) There is no Atmospheric "twinkling".

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SOFIA



- Stratospheric Observatory For Infrared Astronomy (SOFIA)
- Modified Boeing 747
- Operation height: 39000 to 45000 ft (11.8 to 13.7 km)
- 2.7m telescope
- Stratosphere is high enough to observe infrared
- Currently in air testing



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