

Astronomy 122



This Class (Lecture 15):

Exoplanets

Next Class:

Midterm

Midterm on Thursday!

Music: *For Science* – They Might Be Giants

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Midterm



- 50 min exam in this classroom
- Will cover material up to and including star formation
- 40 multiple choice questions.
- Exam will have 105 points graded out of 100 (i.e. extra credit)
- You may bring normal-sized sheet with notes on each side.
- Study guide was posted on class webpage.
- Discussion will focus on your questions and perhaps a Jeopardy game.
- May want to bring a calculator, but don't really NEED one.

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Night Observing



- Make-up dates added to list!
 - March 4-5th: Tuesday-Wednesday
- Bad semester!
- Observing sessions are from 7:30pm-9:30pm (allow 45 mins to complete)



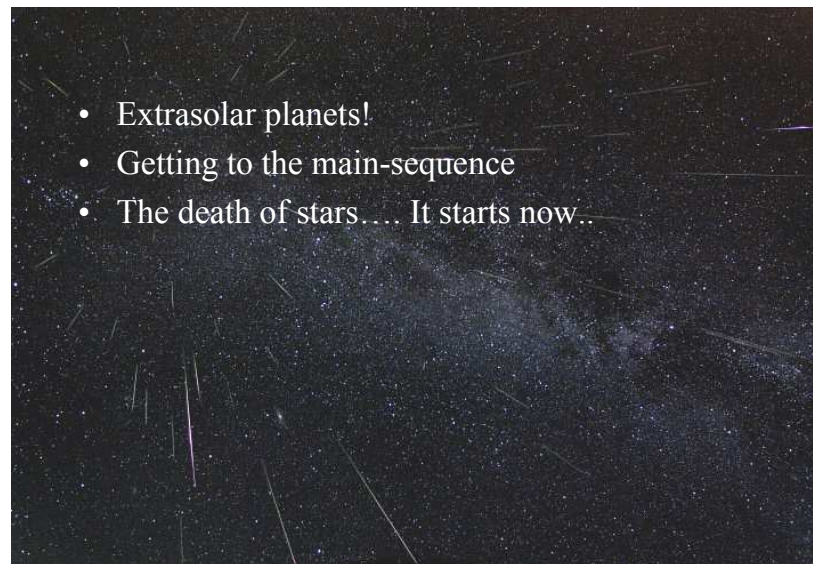
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Outline



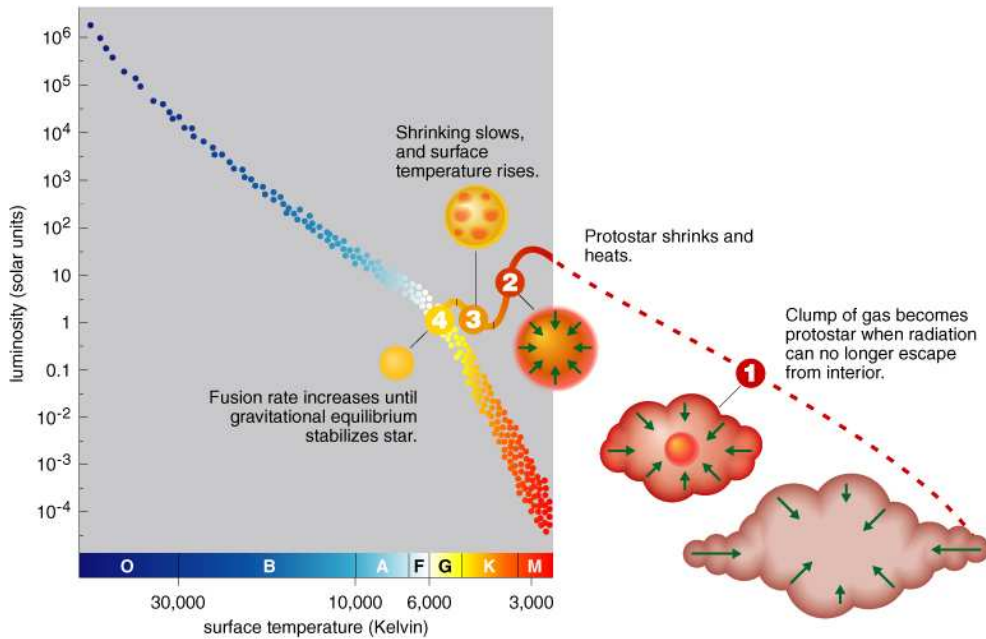
- Extrasolar planets!
- Getting to the main-sequence
- The death of stars.... It starts now..



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A Star is Born!

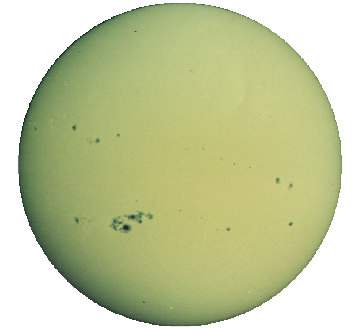


© Addison-Wesley Longman

Becoming a Star



- When the temperature in the pre-main-sequence star's core reaches about 10 million K, hydrogen fusion starts
- Gravitational collapse stops, pressure from fusion in the core halts contraction
- Now a main sequence star!



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Question



A star was born and is now on the main sequence.
Which of the following is **incorrect**?

- It is now burning hydrogen into helium
- It contracted from a molecular cloud
- It is now in hydrostatic equilibrium
- It will spend most of its life on the main sequence
- It will live for as long as nuclear fission can occur in the core.

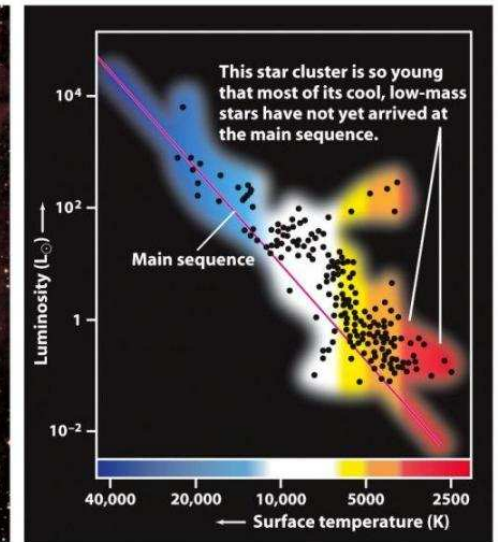
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Very Young Cluster



(a) The star cluster NGC 2264



(b) An H-R diagram of the stars in NGC 2264

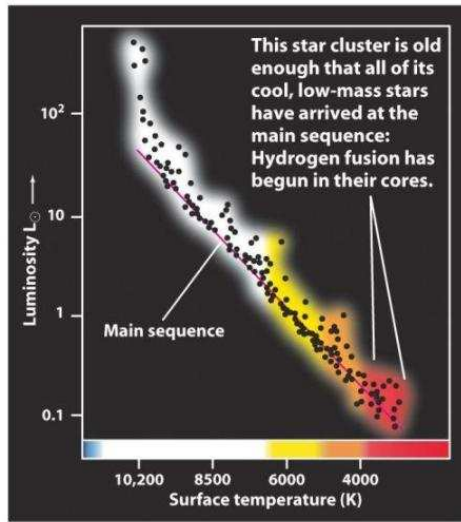
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Older Cluster



(a) The Pleiades star cluster



(b) An H-R diagram of the stars in the Pleiades

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What Are We Looking For? General Predictions of Solar Nebula Theory



- ☺ Are interstellar dust clouds common? **Yes!**
- ☺ Do young stars have disks? **Yes!**
- ? Are the smaller planets near the star?
- ? Are massive planets farther away?

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Test Of Exoplanets



Planets around other stars

= extrasolar planets = **“exoplanets”**

Would our solar system nebula formation theory account for other solar systems around other stars?

Hard to find!

Reflected light from the Earth is 1 billion times fainter than the Sun!!!!

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Finding Planets



1. **Radial Velocity:** **Stars will wobble.**
2. Astrometry: See the stars move.
3. Transit Method: Occultation.
4. Optical Detection: Direct.

Arguable 2 extrasolar planets have been detected directly in the IR. Remember that planets in our Solar System seem bright because they reflect light from the Sun in the visible.

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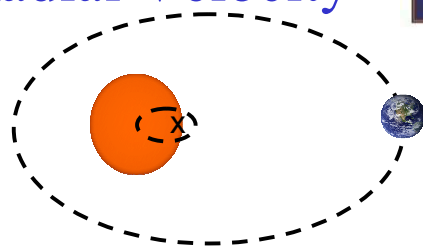
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Star Wobble: Radial Velocity

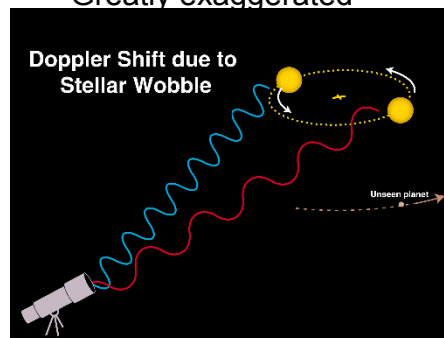


Newton's 3rd Law:

- Both planet and star move
- Both orbits fixed around the "center of gravity"
- Star's period? Place your bets...
 - Same as planet
- Star movement too small to see
 - Moves in small, tight circle
 - But "wobble" in star speed detected!



Greatly exaggerated



<http://www.howstuffworks.com/planet-hunting2.htm>

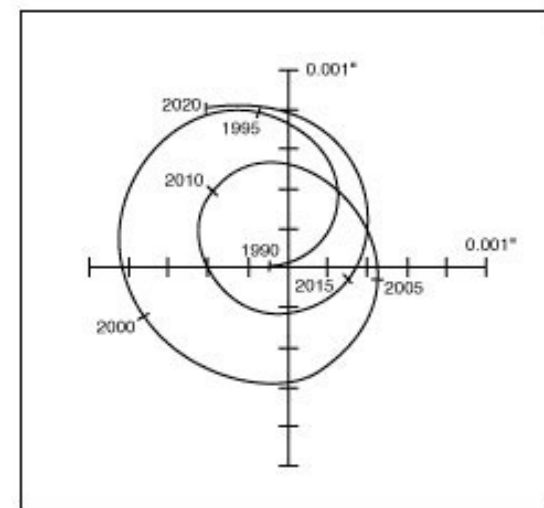
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The Sun's Wobble



Astrometric displacement of the Sun due to Jupiter (and other planets) as at it would be observed from 10 parsecs, or about 33 light-years.



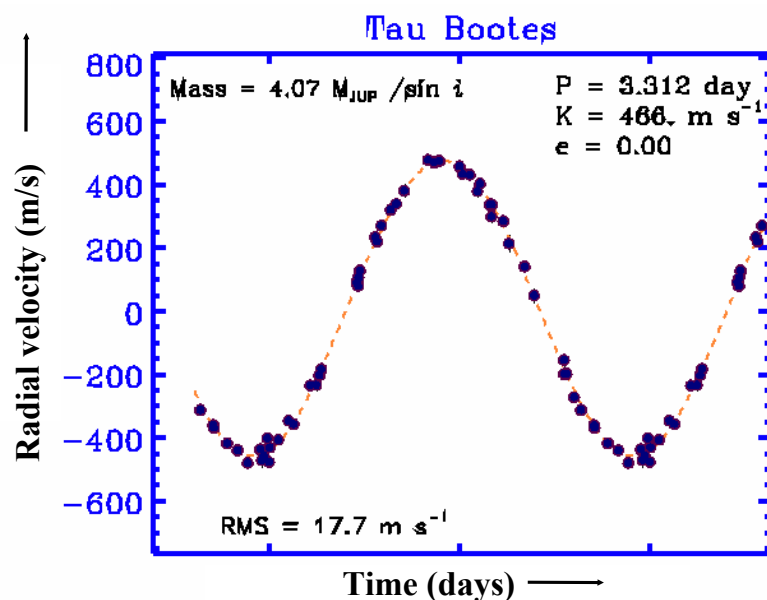
http://planetquest.jpl.nasa.gov/Keck/astro_tech.html

If we could observe this, we could derive the planetary systems— also called astrometry.

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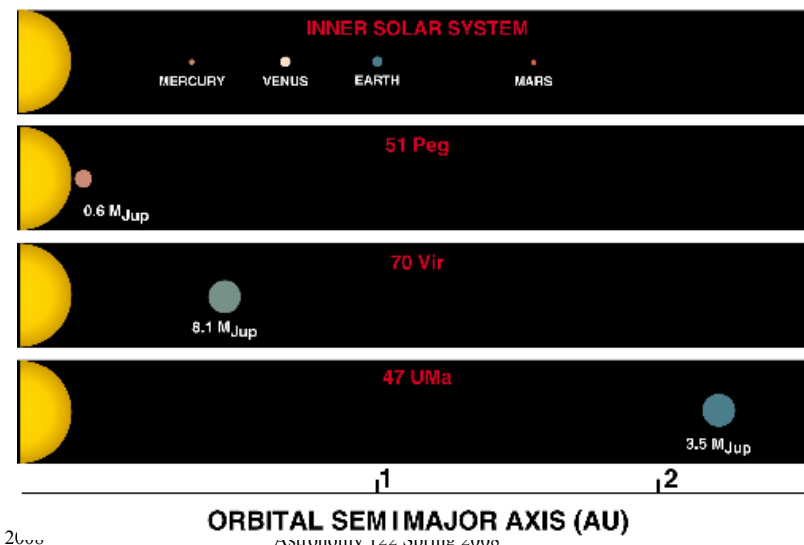
Radial Velocity Shifts: Planets around other Stars?



Early Discovery-- 1996



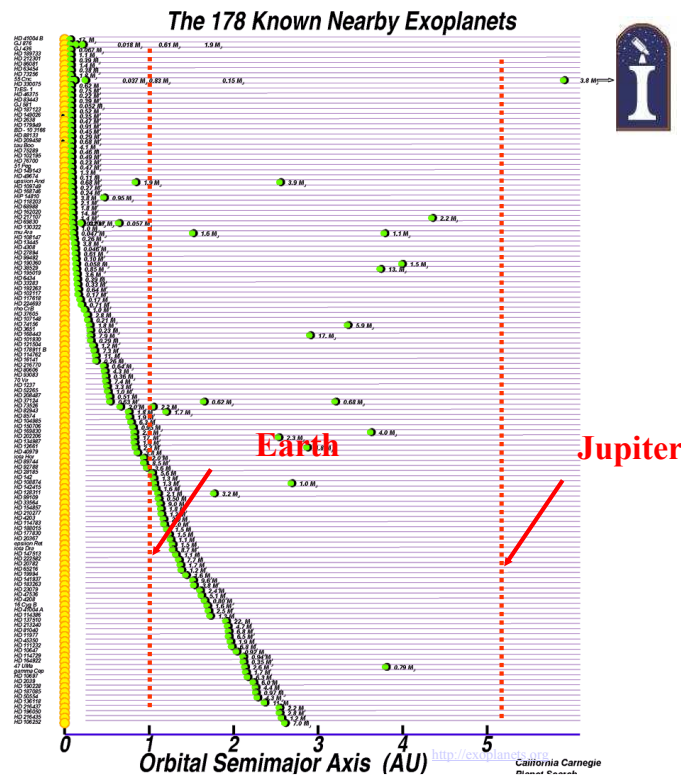
PLANETS AROUND NORMAL STARS



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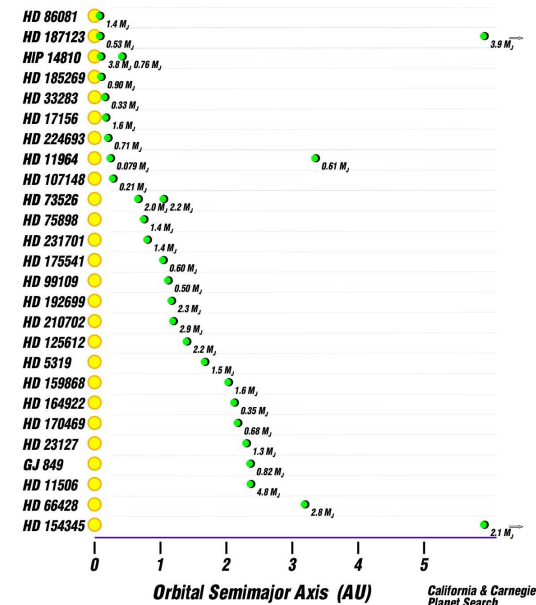
As of today,
there are
228 planets
known
around
nearby
stars.



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Recent Ones

28 New Exoplanets



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Question

Most exoplanetary systems found so far are

- Just like our own solar system.
- Very different than our solar system.
- It is hard to determine since we have only found a handful so far.
- None of the above.

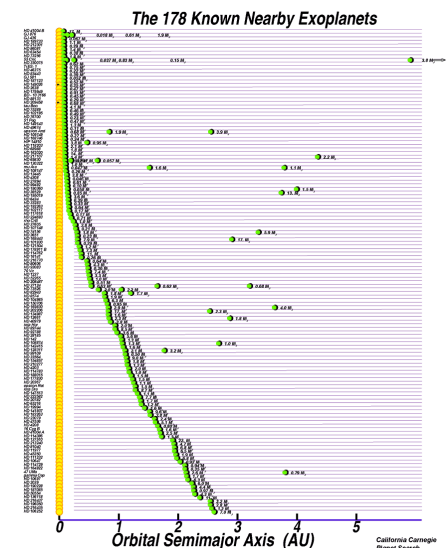
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Exoplanets: Results to Date

Over 228 planets detected so far

- More than 25 times the number in our Solar System!
- By measuring the wobble variation:
 - With time, gives the planet distance: Kepler's 3rd law
 - The orbital speed of the star gives masses: the bigger the wobble amplitude, the heavier the planet



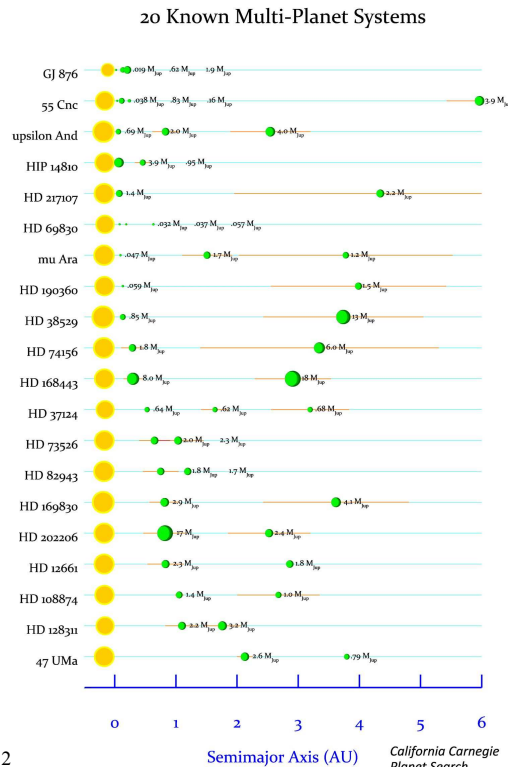
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Exoplanets: Results to Date

At least 20 multi-planet
systems!

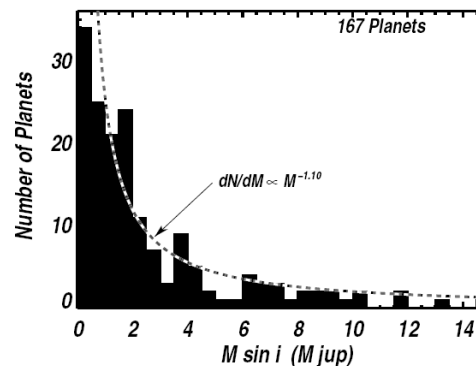
Note: Jupiter is 318 time the
mass of Earth or
 $M_E = 0.003 M_J$
 $M_J = 0.001 M_{Sun}$
 $M_{red\ dwarf} = 80 M_J$
 $M_{brown\ dwarf} = 18 M_J$
 $Period_J = 12\ years$



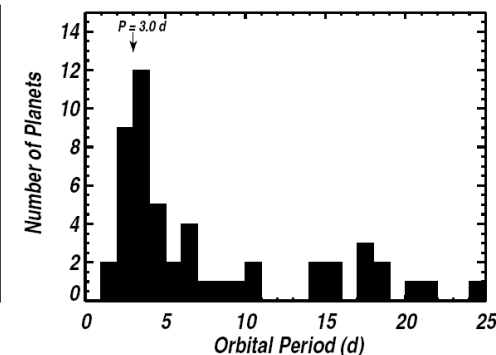
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Masses/Periods of Extrasolar Planets



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<http://exoplanets.org>

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List

<http://exoplanets.org/>

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Other Planets, Other Stars

47 Ursae Majoris
System– 51 light
years away (near
the Big Dipper).
13 years of data
has shown 2
planets– 1 Jupiter
like and 1 Saturn
like.



Wow! Among the most similar to our own system

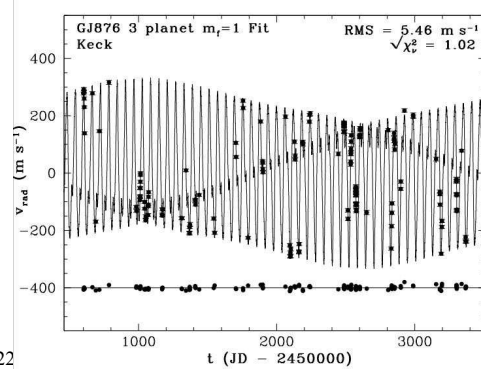
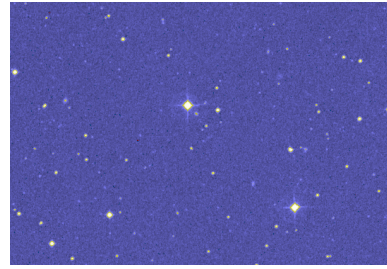
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The Lowest Mass to Date



GJ 876 – a Red Dwarf that is 15 light years away (in Aquarius). Has three planets! 2 Jupiter-like and one that is 6-8 Earth masses! But all are inside 1 AU!



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The Lowest Mass to Date



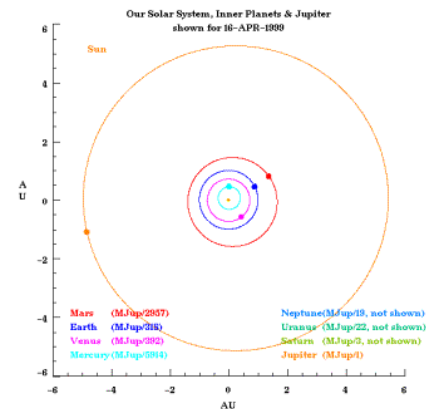
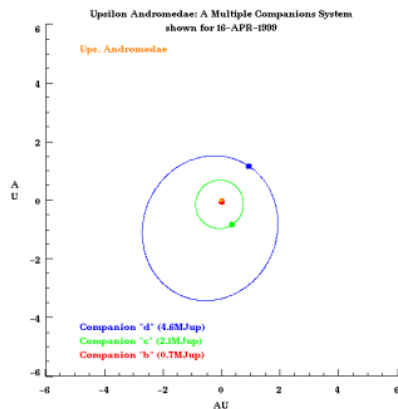
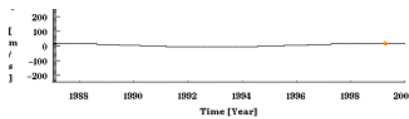
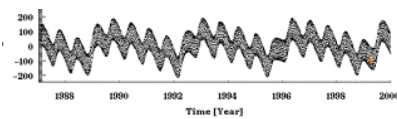
GJ 876 – a Red Dwarf that is 15 light years away (in Aquarius). Has three planets! 2 Jupiter-like and one that is 6-8 Earth masses! But all are inside 1 AU!



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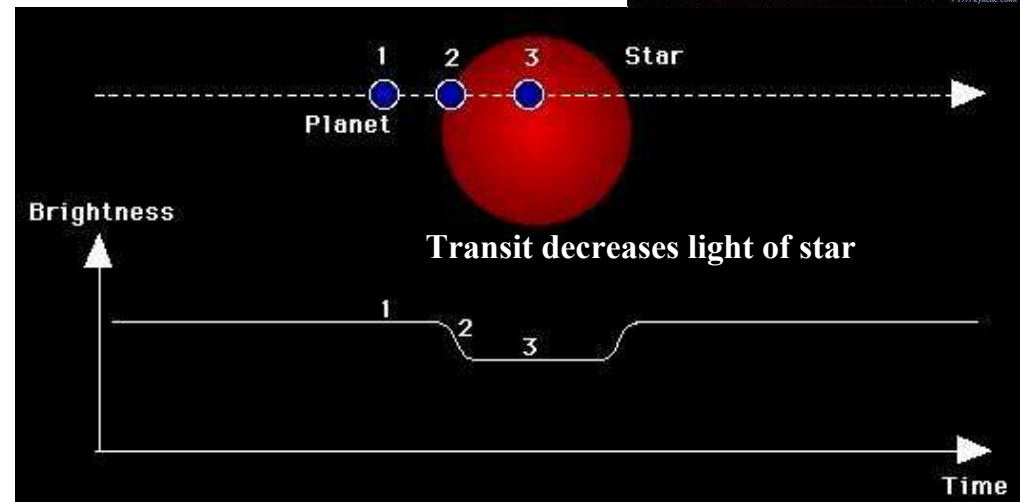
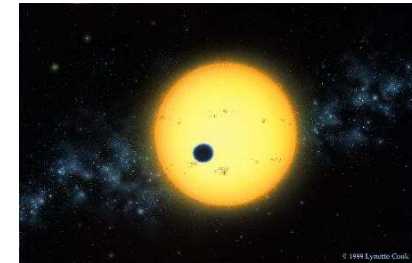
Detecting the Solar System



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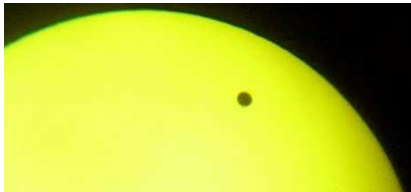
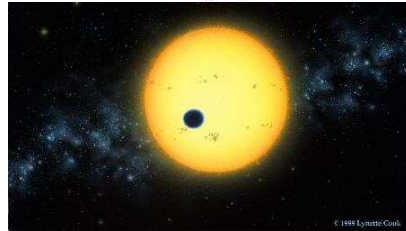
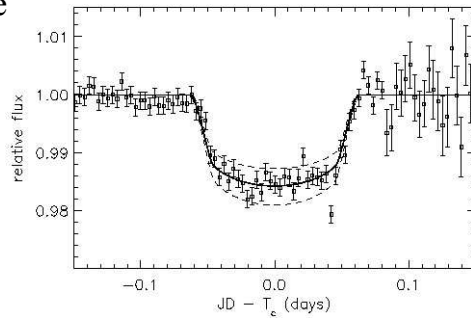
Planet Transits



Transits



- The planet passes in front of the star— like Venus 2004.
- Can find planet radius
- Best chance of finding Earth-like planets
- Requires the extrasolar planet's orbital plane to be pointed at Earth
- <http://www.howstuffworks.com/planet-hunting2.htm>

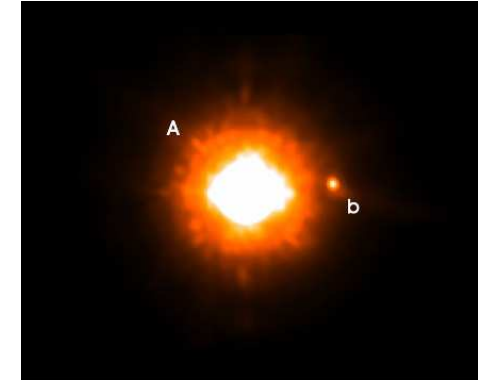


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Direct Detection?



- The race is on to directly image a planet in the IR, it is still difficult to determine the stellar mass.
- Best example so far is an adaptive optic image from April, so planet or brown dwarf?



The Sub-Stellar Companion to GQ Lupi (NACO/VLT)
190 PR Photo 10a/05 (7 April 2005)
© European Southern Observatory

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Results to Date



No surprise

- ✓ New planets are massive
- ✓ Why? Big planets make a big wobble
- ✓ If not massive, we could not have found them
- ✓ About 3-5% of all stars have some type of planet.



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Results to Date



Big surprises

- ? Most periods of only *a few days*!
- ? Most planets are very near their stars!
- ? τ Bootes' planet is 3.6 times Jupiter's mass, but it's orbit smaller than Mercury's!
- ? If a Jupiter-like planet formed close in, perhaps that prevents terrestrial planets from forming.



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What Are We Looking For?

General Predictions of Solar Nebula Theory



- ☺ Are interstellar dust clouds common? **Yes!**
- ☺ Do young stars have disks? **Yes!**
- ? Are the smaller planets near the star?
Not the ones found so far! Haven't found smaller planets yet!
- ? Are massive planets farther away?
Not most of the ones found so far!

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Important Caveat



- Our current observations of extrasolar planets do **not** exclude planetary systems like our solar system
- Current instruments are most sensitive to large planets close to their stars
 - Big planet - big wobble
 - Close planet - fast wobble
- We only have a little over 10 years of data – 1 orbit's worth for Jupiter
- To find solar-type systems, we need more sensitive equipment

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Exoplanets: Implications



Solar Nebula **Theory**:

- Giant planets born far from star

Exoplanet Data:

- Giant planets found very close

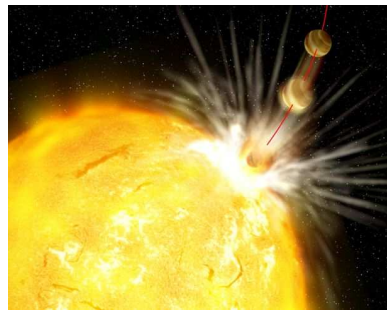
Theory is incomplete/wrong!

New questions:

- ? Who is normal: Them or us?
- ? Are giant planets born close in?
- ? Are some giant planets born far out, move in?
“planet swallowing”!?!

Anyway: Planets are common!

- ✓ Good news in search for life elsewhere...maybe



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Exo-Question



What do the exoplanets suggest about our idea of how planets form?

- a) To date, the planets we have found, although in the wrong orbits, are similar to our Solar System.
- b) Something is wrong with our picture of star formation.
- c) Something may be wrong with our picture, we need more data.
- d) Sometimes some planetary systems are formed that are different, but we don't yet know who is weird– us or them.
- e) None of the above.

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Future Projects



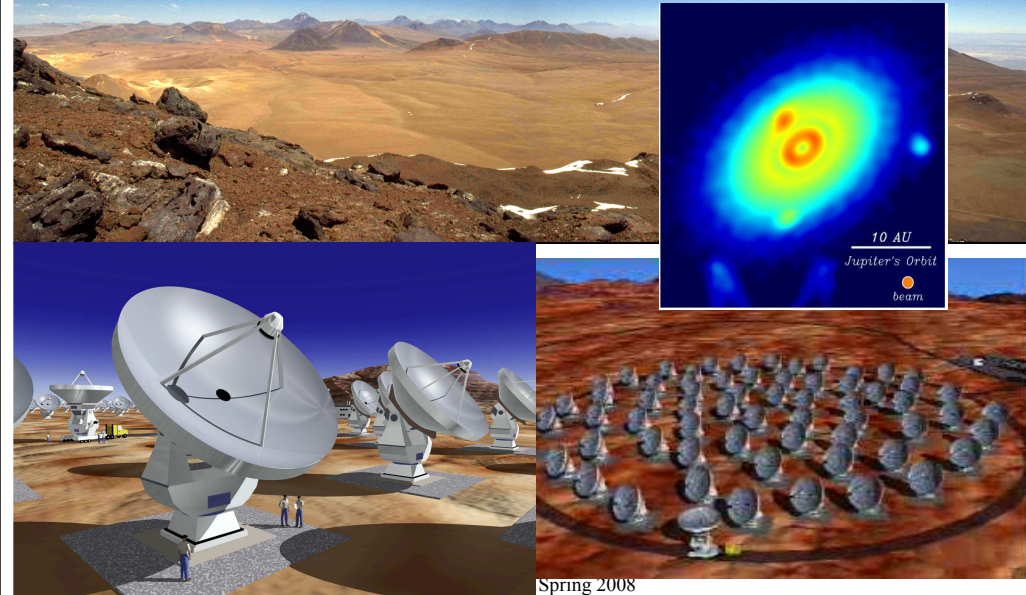
- Atacama Large Millimeter Array (ALMA): 2010
 - mm interferometer:
 - direct detection of young gas giants
- Kepler: Feb 2009
 - Planet Transits
- Next Generation Space Telescope
 - James Webb Space Telescope (JWST): 2013
 - Direct imaging of forming gas giants?
- Space Interferometry Mission (SIM): 2016?
 - Astrometry
- Terrestrial Planet Finder (TPF): Mission 1: deferred
 - Coronagraph
 - IR interferometer
- Terrestrial Planet Finder (TPF): Mission 2: deferred
 - A large-baseline infrared interferometer. Imaging extrasolar Earths!!!!

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50?
~~64~~ × 12 m @ 16,400 ft Chajnantor Chile

ALMA -- 2010

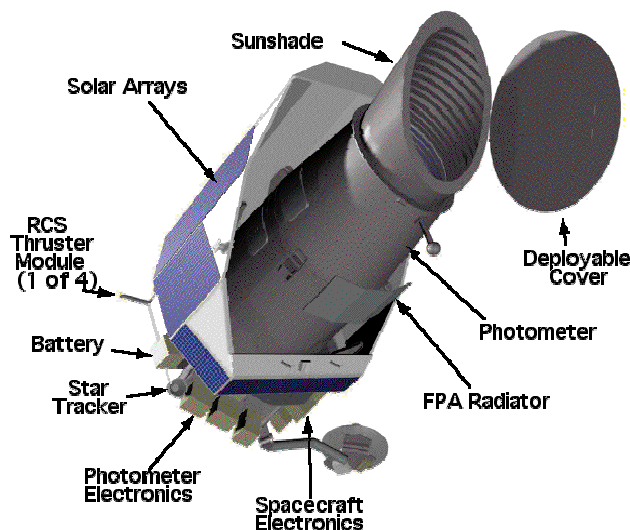


Kepler



1.4 meter mirror,
measuring accurate
brightness of stars.

A terrestrial-sized
Earth-like planet
would dim the star's
light by 1/10,000th –
comparable to
watching a gnat fly
across the beam of a
searchlight.



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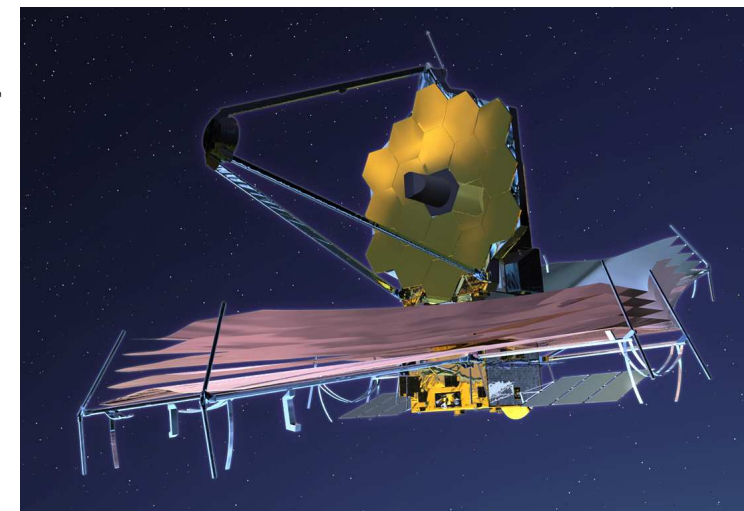
JWST



James Webb
Space Telescope:
Successor to HST

6.5 meter
observatory

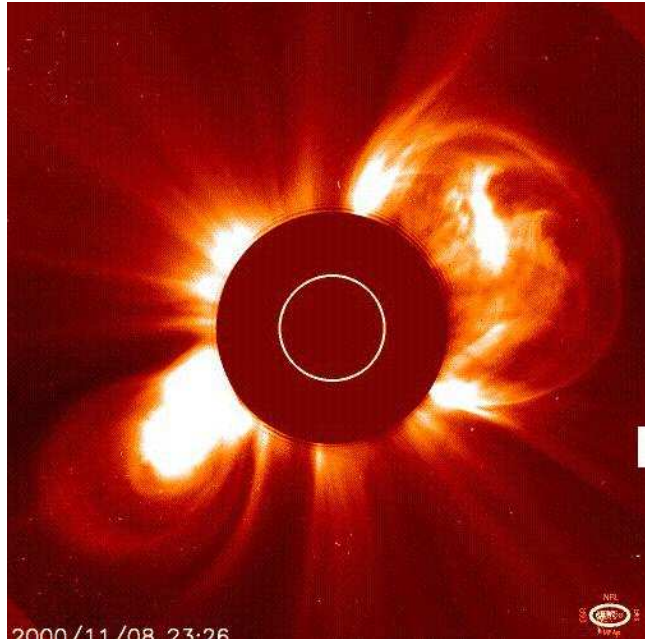
Working in the
infrared with a
coronagraph.



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The Coronagraph Advantage



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Space Interferometry Mission



Accurately measure location of stars to micro-arcseconds.

Need to know relative location of components to 50 pm.

Funding in question.



http://planetquest.jpl.nasa.gov/SIM/sim_index.html

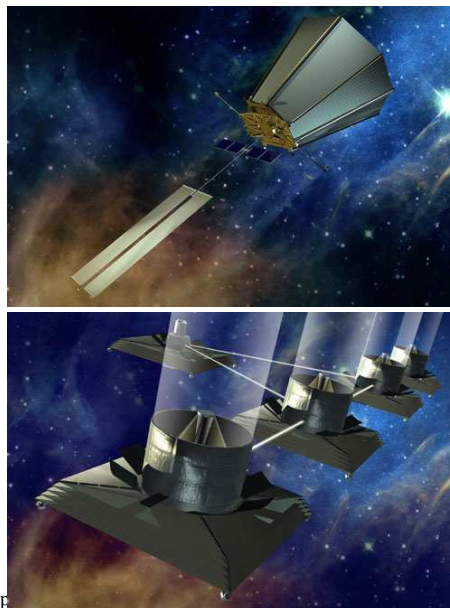
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Terrestrial Planet Finder Mission: Two Telescopes



- Survey nearby stars looking for terrestrial-size planets in the "habitable zone"
- Follow up brightest candidates looking for atmospheric signatures, habitability, or life itself
- Then, the ultimate: image the little blue dots



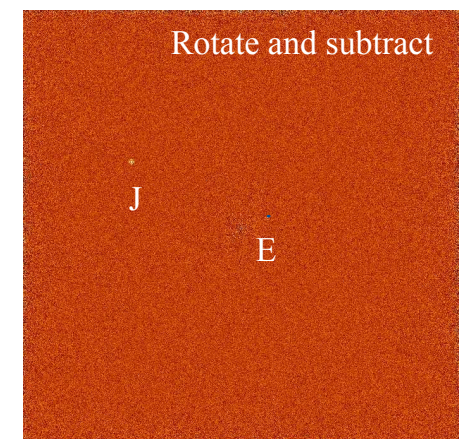
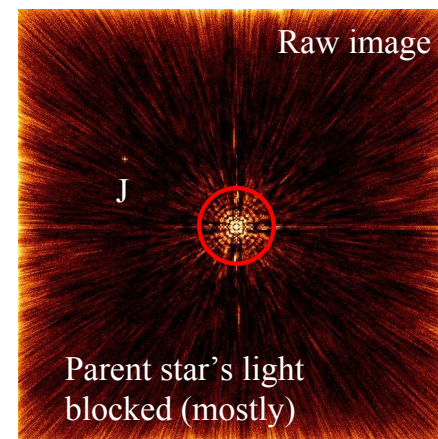
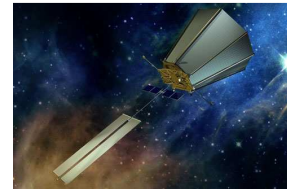
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TPF

Visual wavelength 'coronagraph'

- Find Earth-like planets
- Characterize their atmospheres, surfaces
- Search for bio-signatures of life (O_2 , H_2O , etc)



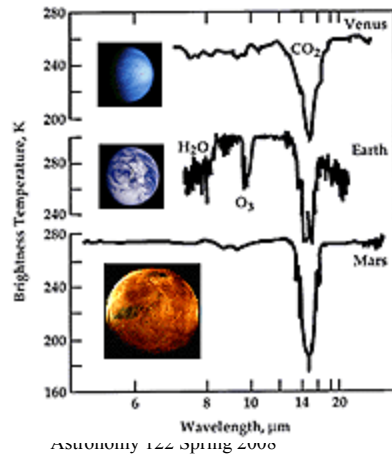
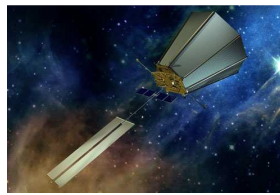
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TPF

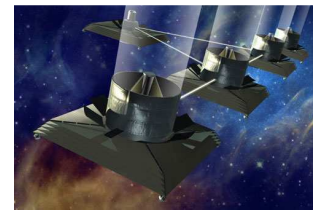
Visual wavelength 'coronagraph'

- Find Earth-like planets
- Characterize their atmospheres, surfaces
- Search for bio-signatures of life (O_2 , H_2O , etc)



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TPF: Step 2



The goal of imaging an Earth-like planet.

5 platforms of 4 eight meter interferometer in space.



spider.ipac.caltech.edu/staff/jarrett
frank.LIU/origins/openhouse30.html

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TPI -- Scales



Pixel / Diameter	Pixel size @ planet (km)	Image	Interferometer Requirements	
			Collecting Area	Baseline
400	32		IR Visible 144 km ² 1,296 km ²	100,000 km 5,000 km
100	128		IR Visible 0.64 km ² 5.76 km ²	24,000 km 1,200 km
			Collecting Area	Baseline
25	510		IR Visible 1,024 m ² 9,216 m ²	6,000 km 303 km
10	1276		IR Visible 64 m ² 576 m ²	2,4km 120 km

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