Astronomy 122 Section 1– TR 1300-1350



1320 Digital Computer Laboratory

Leslie Looney This Class (Lecture 4):

Phone: 244-3615
Email: lwl @ uiuc . edu

The Earth-Moon System

Office: Astro Building #218

Office Hours:

T 10:30-11:30 a.m. or by appointment

Next Class:

Gravity and the Planets

Homework #1 due Fri!

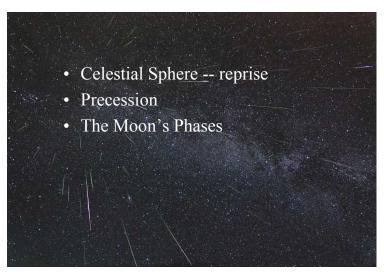
http://eeyore.astro.uiuc.edu/~lwl/classes/astro122/spring06/

Music: Walking on the Moon – The Police

Jan 26, 2005 Astronomy 122 Spring 2006

Outline





Astronomy 122 Spring 2006

Jan 26, 2005

Celestial Sphere



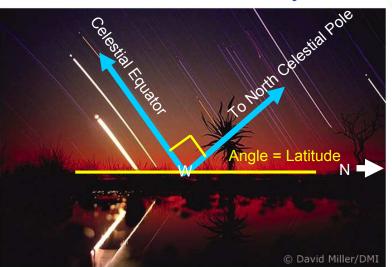
• ..\..\..\animation\celsphere1.avi



 $http://brahms.phy.vanderbilt.edu/{\sim}rknop/astromovies/celsphere1.html\\$

Motions in the Sky



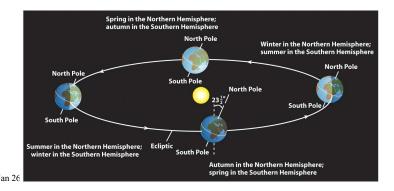


What Causes the Seasons?



So what does cause the seasons?

- It's the tilt of the Earth's spin axis
 - · Affects the length of day and intensity of sunlight



"Thinking Cap" Question



June 21st is the summer solstice and December 21st is the winter solstice. However, they are not the hottest and coldest days of the year (those occur in July and January, on average). Why is this?

Jan 26, 2005

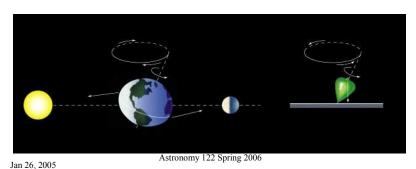
Astronomy 122 Spring 2006

Precession



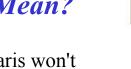
- As the Earth spins it also wobbles slowly, like a top
 - This wobble takes 26,000 years
 - Geek-speak: Called **precession**





Astronomy 122 Spring 2006

What Does This Mean?



- CYGNUS

 Deneb

 AD 8000

 CASSIOPEIA

 LYRA

 Vega

 AD 15,000

 DRACO

 Path of north celestial pole

 CORONA
 BOREALI9

 Deneb

 AD 1000

 DRACO

 Path of north celestial pole

 URSA
 MINOR

 Thuban

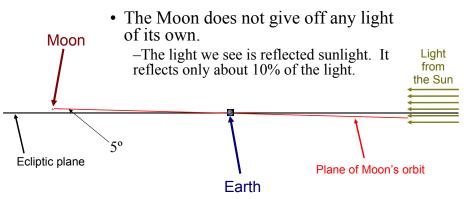
 5000 BC

 URSA MAJOR
- Polaris won't always be the North Star!
- In the time of the Egyptian pharaohs, it was Thuban
- In about 13,000 AD it will be Vega (sort of)

Jan 26, 2005

The Lunar Orbit

- Ì
- The Moon is Earth's nearest neighbour in space
 - About 30 Earth diameters away
- Orbits the Earth once in a little under a month
 - Like the Earth orbits the Sun, the Moon orbits the Earth counter-clockwise



Phases of the Moon



- Over the orbit, the Moon's appearance changes radically
- The apparent **Phases** of the Moon depend on how much of the sun-lit side of the Moon we can see.
- This is caused by the relative positions of the Earth, Moon, and Sun.
- Not caused by the shadow of the Earth



Jan 26, 2005

Astronomy 122 Spring 2006

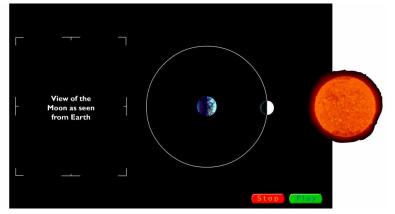


Waxing crescent

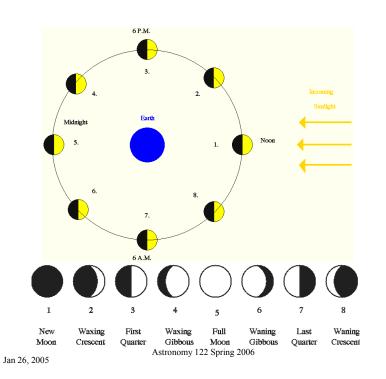
The Cycle of Phases



As the Moon orbits the Earth, we see it go through a cycle of phases



Jan 26, 2005



At Home Phases Demo



- Hold a softball (or equivalent) toward the sun (or a lightbulb)
- Spin around, and watch the ball experience phases like the Moon!





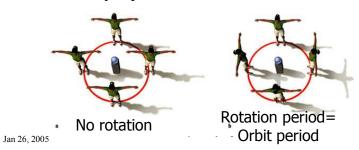
nttp://www.astro.uiuc.edu/projects/data/MoonPhases/index.html

Astronomy 122 Spring 2006

Jan 26, 2005

The Face of the Moon

- Ì
- Did you notice that we only see one face of the Moon?
- Does this mean the moon doesn't rotate?
- *No*, the Moon rotates so that the same face is always pointed at the Earth
- A lunar day equals a lunar orbit!



Dark Side of the Moon?

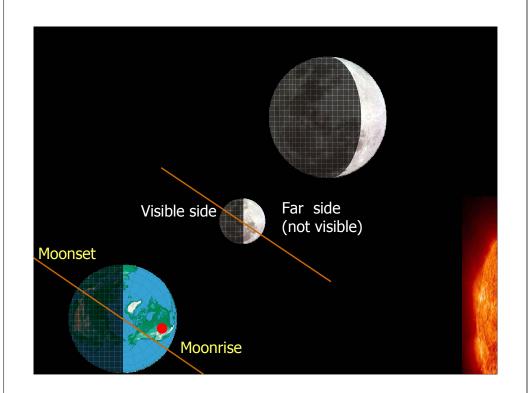


Is there really a dark side of the Moon?

NO! It is better called the Far Side of the Moon.

There is a side we don't see, but during the New Moon phase, it is well lit.

Basically the lunar day is nearly a month long.



A "Moonth"



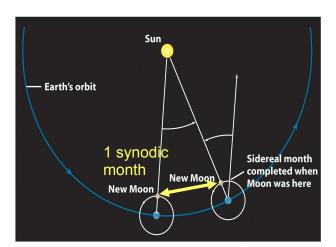
- The month is based on the time it takes the Moon to cycle through its phases
 - 29.5 days called *synodic period*
- Moon makes one full orbit of the Earth in 27.3 days
 - Called its sidereal period
- Because of the Earth's orbit about the Sun, the Moon travels more than a full orbit each synodic period
 - -360 degrees/27.3 days = 13.2 degrees/day
 - Earth rotates about 15 degrees/hour
 - So Moon rises 13.2/15 hours = 53 minutes later each day

Jan 26, 2005

Astronomy 122 Spring 2006

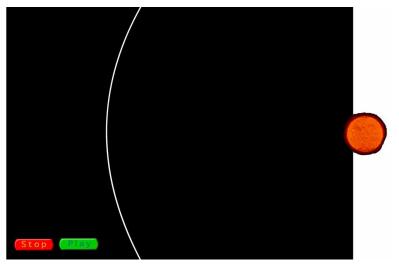
Synodic vs. Sidereal Period





Synodic vs. Sidereal Period



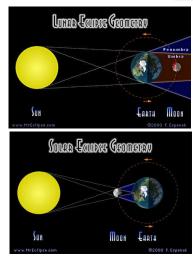


Astronomy 122 Spring 2006

Basic of Eclipses



- Lunar Eclipse
 - When the Moon passes into the Earth's shadow
 - Sun Earth Moon *full moon*
- Solar Eclipse
 - When the Earth crosses the Moon's shadow
 - Sun Moon Earth new moon
- Why don't eclipses happen every full and new moon?



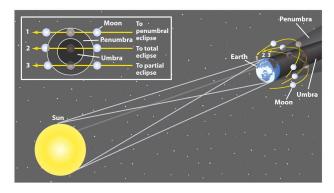
Jan 26, 2005

Astronomy 122 Spring 2006

Lunar Eclipses



- Three types
 - Total The Moon goes completely through the Earth's umbra
 - Partial The Moon only goes partly through the umbra
 - Penumbral The Moon only passes into the penumbra



Umbra? Penumbra?

Latin for "complete shadow" and "partial shadow"

Total Lunar Eclipse- Time Lapse



• Occurs when the Moon passes through Earth's umbra completely.



- Occur roughly twice a year, and last for about an hour or two.
- Can be seen by anyone experiencing night during the lunar eclipse.

Red Moon



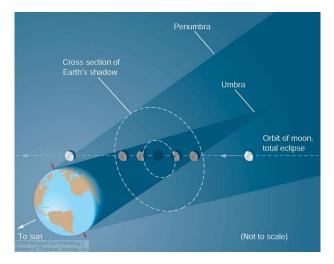
• During a total lunar eclipse, the Moon turns a blood-red/burnt orange color.



- Red color caused by sunlight diffused through the Earth's atmosphere
- Atmosphere scatters blue light away, so Moon is dimly illuminated in red
- The next total eclipse visible in Champaign-Urbana will be in 2007.

Total Lunar Eclipse



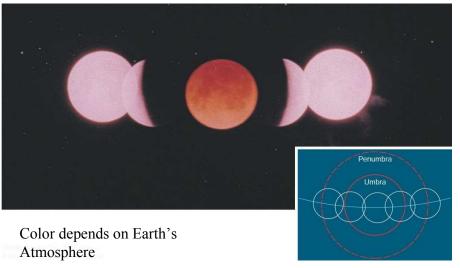


Jan 26, 2005

Astronomy 122 Spring 2006

Eclipsed





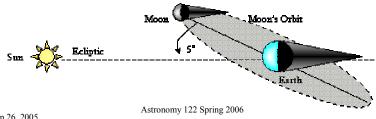
Jan 26, 2005

Astronomy 122 Spring 2006

The Moon's Orbit is Tilted!



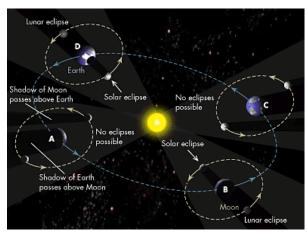
- The Moon's orbit is tilted to the ecliptic by 5°
- Just like the Earth's rotation axis is tilted to the ecliptic
- The Moon must be near the ecliptic for an eclipse to occur



Eclipse Seasons



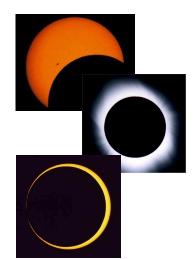
Nodes are the two points in each orbit at which the Moon crosses the Earth's orbital plane. For lunar or solar eclipses to occur the nodes must be aligned with the Earth and the Sun. Hence, eclipses can occur only twice per year and these epochs are called eclipse seasons.



Solar Eclipses



- Earth passes into the Moon's shadow
- Only occur at the new moon
- Three types
 - Partial when the moon only partially blocks the sun
 - Total when the moon completely blocks the sun
 - Annular when the moon is too small to completely block the sun



Astronomy 122 Spring 2006

Jan 26, 2005