

# Astronomy 122



This Class (Lecture 24):

Large Scale Structure

***HW10 due on Friday.***

Next Class:

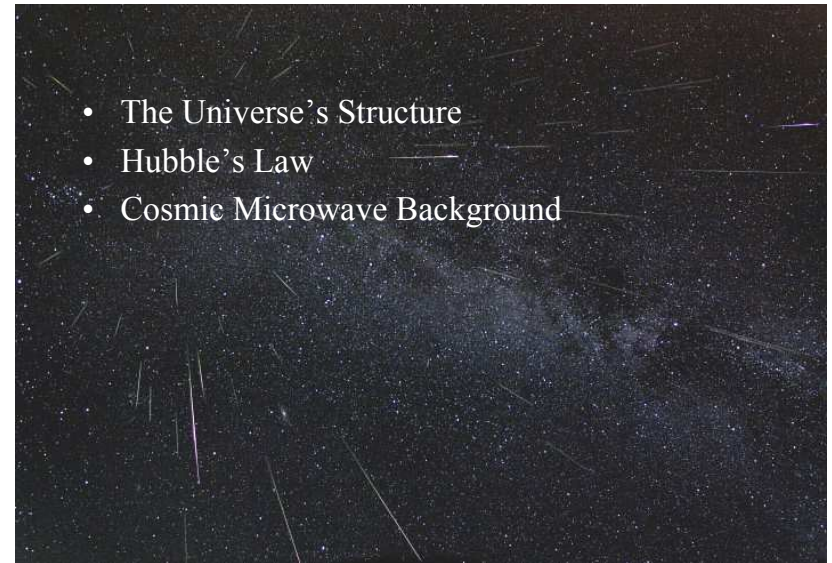
Active Galaxies & Quasars

Music: *Rocket Man* – Elton John

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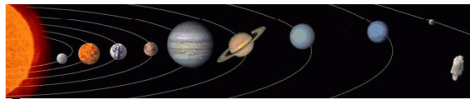
# Outline



- The Universe's Structure
- Hubble's Law
- Cosmic Microwave Background

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***Astronomy:  
The Big Picture***  
*Moving from our Galaxy outward!*



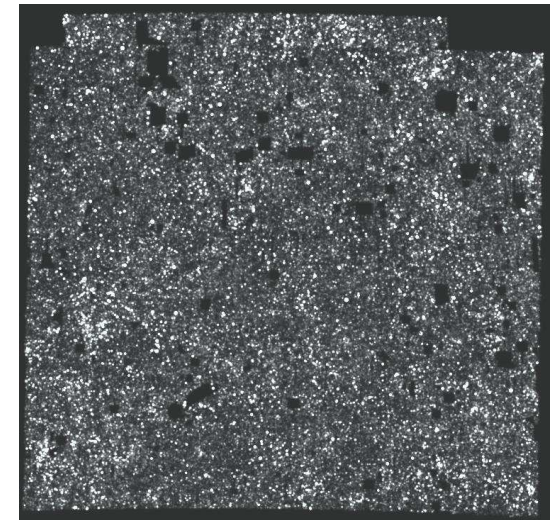
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# The Universe



- On small scales—  
clumpy
- On large scale—  
smooth
  - 4 deg x 4 deg
  - Each point is a  
Galaxy
  - About 710,000



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<http://www-int.stsci.edu/~postman/deeprange.html>

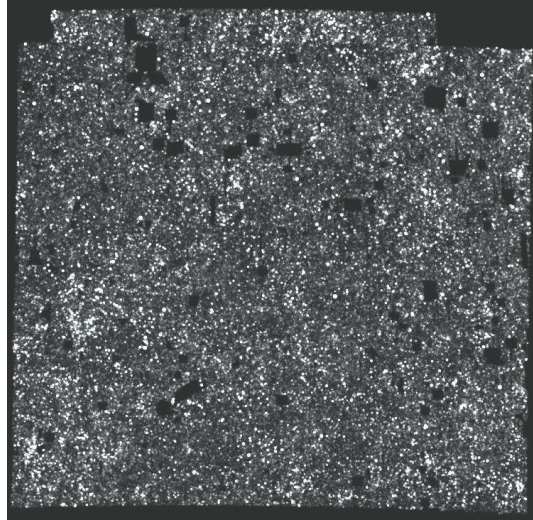
# The Universe



## The Universe is

1. Homogeneous (gal's uniformly fill space)
2. Isotropic (looks same in all directions)

- These are the starting points for our Cosmological journey



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# Measuring the Distance



- Basic idea:
  - Find the intrinsic brightness of an object
  - Compare to the observed brightness
  - Use brightness difference to determine distance
- No method is good for all distances
  - Different techniques have been developed for different distances
  - To find distance to Andromeda galaxy, Hubble used Cepheid variables.

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# How are Galaxies Moving?



It's 1928 and Edwin Hubble is measuring how galaxies move by measuring the velocity WRT us.

What does he find?

- a) More galaxies receding than approaching.
- b) More galaxies approaching than receding.
- c) About equal numbers of each.

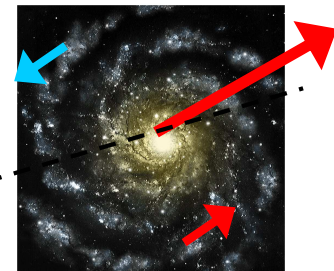
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# Redshift of Galaxies



- Most galaxies are moving away from us!
- The farther away, the faster they are moving away.
- What does this mean?
- Key to understanding the Universe!



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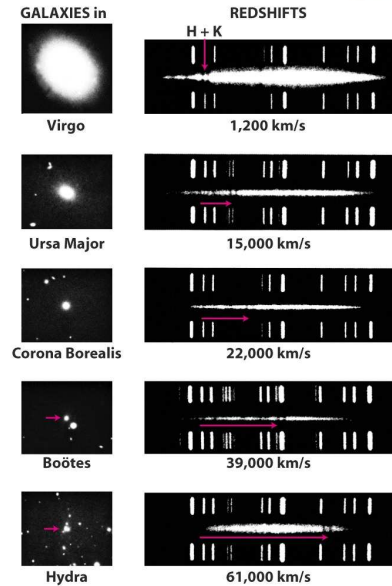
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## Redshifted Galaxies



- Hubble observed that the spectrum lines of most galaxies are **redshifted**
- Redshift:  

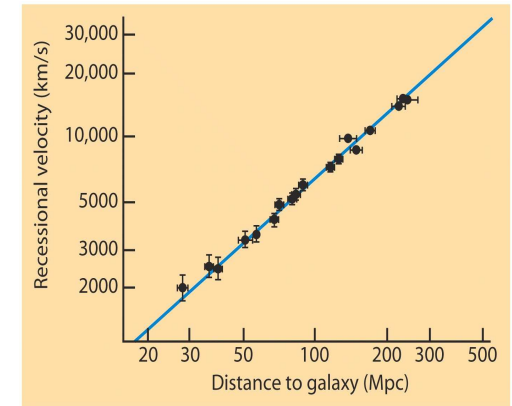
$$z = (\lambda_{obs} - \lambda) / \lambda$$
- At low redshifts,  $z \ll 1$  &  $v = cz$
- Exceptions are the closest galaxies*



## The Hubble Law



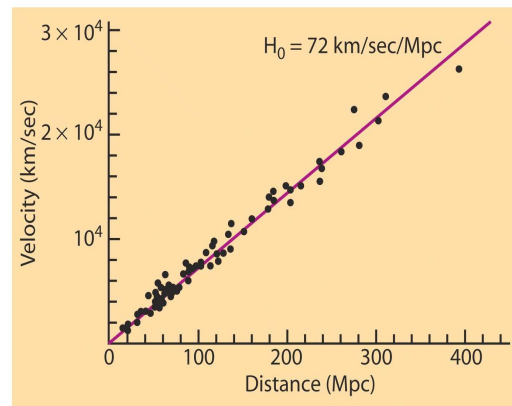
- Calculate velocity from Doppler and distance from Cepheid variables
- $V = H_0 \times D$ 
  - Where  $v$  is velocity,  $d$  is distance, and  $H_0$  is the Hubble constant
- Current best value is  $H_0 = 72 \text{ km/s/Mpc}$



## Redshift and Distance



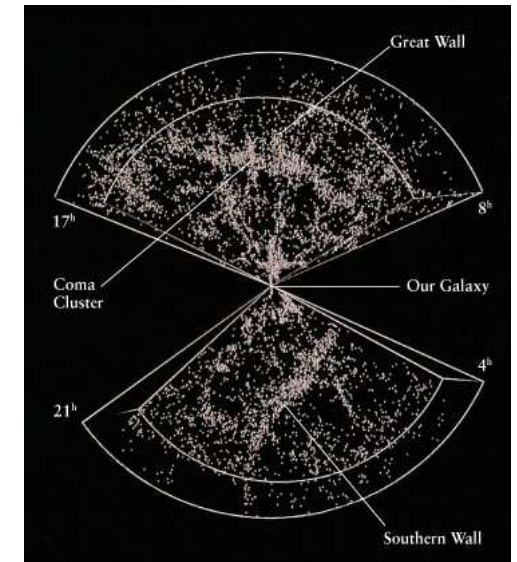
- The Hubble Law gives us a new way of finding distances
- Remember,
  - $v = cz$
  - $v = H_0 d$
- Therefore,  $d = cz / H_0$ !
- We can now measure distances to extremely distant galaxies!



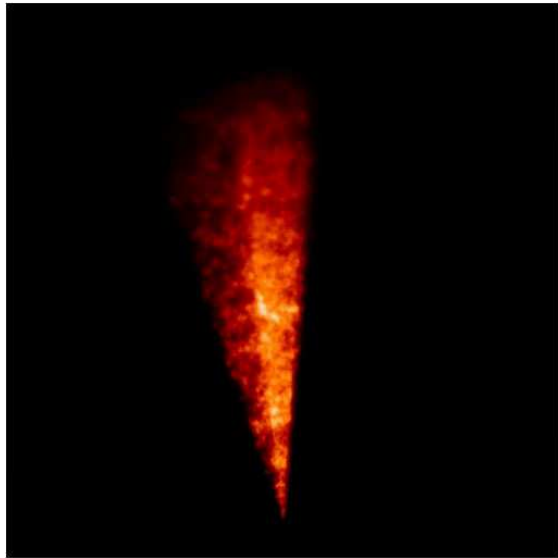
## Structure of the Universe



Clusters of galaxies are grouped together in **superclusters**



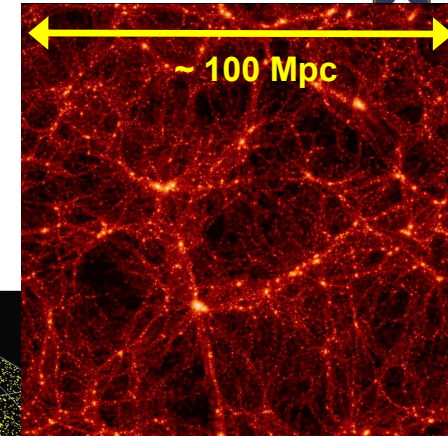
# Structure of the Universe



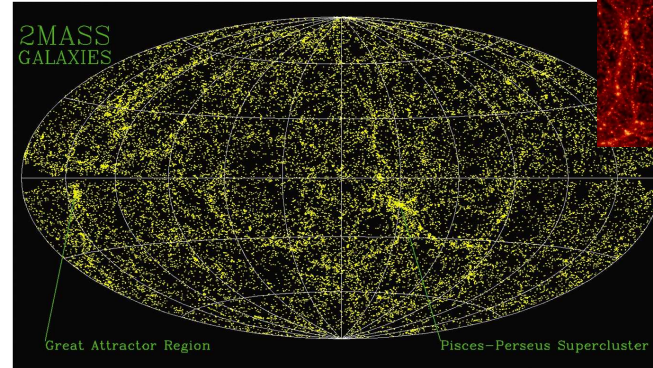
# Structure of Universe



- Superclusters are distributed in Universe.
- Filamentary and sheet structure.
- Voids of nothing between them.

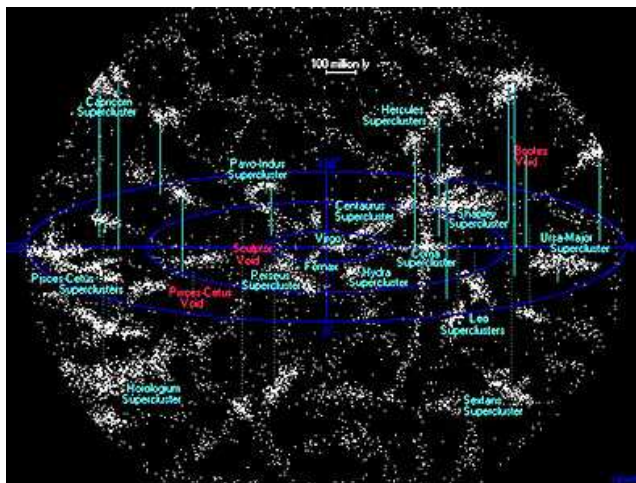


Computer simulation (A. Jenkins)



The arrangement of walls, filaments, and voids resembles soap bubbles

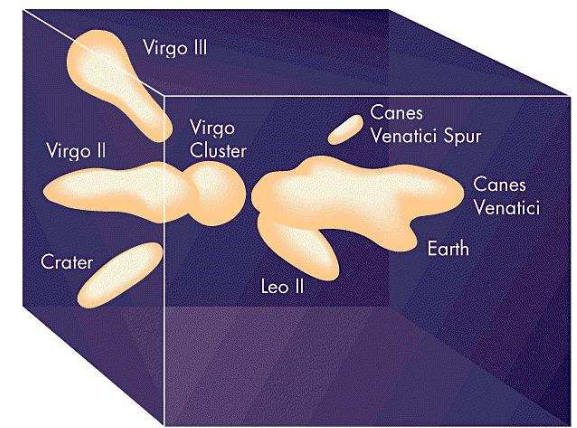
# “Sudsy” Universe



# The Local Supercluster



- Our Local Group is part of a supercluster centered on the Virgo Cluster
- The local supercluster is still expanding
- We are getting farther from the galaxies in the local supercluster





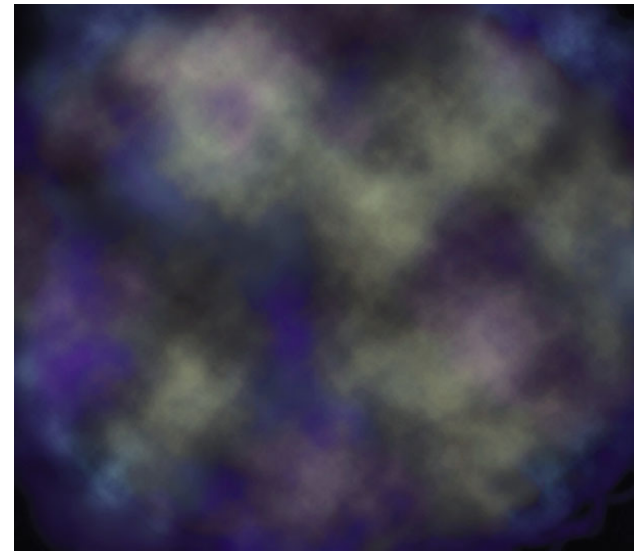
# How did galaxies form?

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# Galaxy Birth



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## Basic Assumptions

- Matter originally filled all of space almost uniformly
- Gravity of denser regions pulled in surrounding matter
- Probably condensed around regions of dark matter



# Galaxy Birth



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- Denser regions contracted, forming *protogalactic clouds*
- H and He gases in these clouds formed the first stars



# Galaxy Birth



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- Supernova explosions from first stars kept much of the gas from forming stars
- Leftover gas settled into spinning disk
- *Conservation of spin*

# But why do some galaxies end up looking so different?



NGC 4414



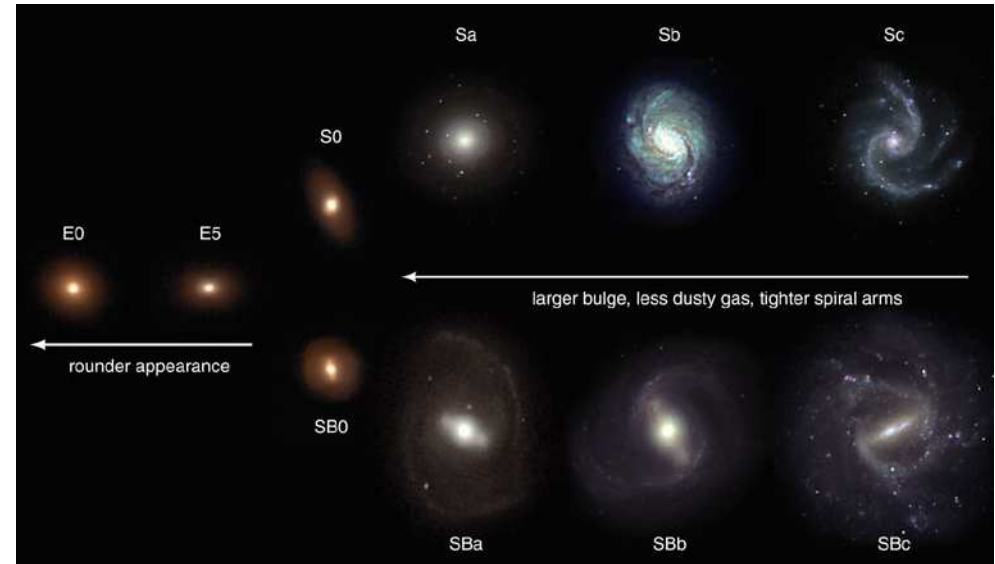
M87



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# Why don't all galaxies have similar disks?



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# Conditions in the Protogalactic Cloud?

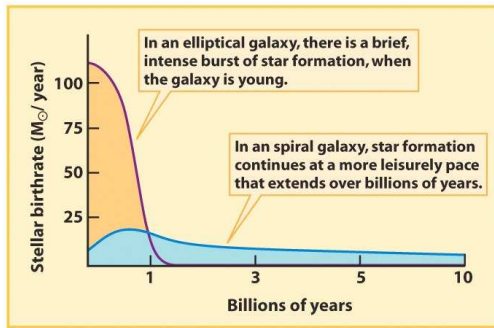


1. Stars form gradually within a protogalaxy.    2. Gas not involved in star formation collapses to form a disk.    3. A spiral galaxy results.

(a) Formation of a spiral galaxy

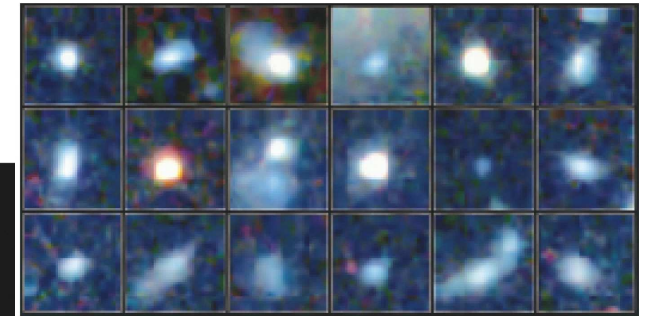
1. Stars form rapidly within a protogalaxy.    2. Gas is quickly consumed to make stars.    3. A elliptical galaxy results.

(b) Formation of an elliptical galaxy

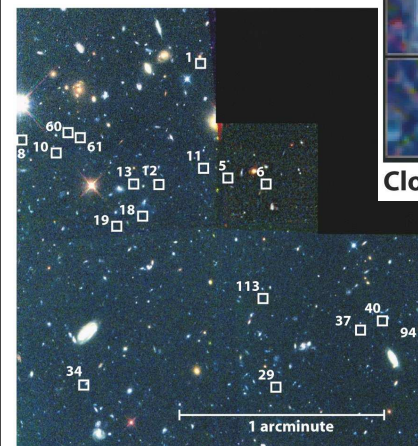


(c) The stellar birthrate in galaxies

# Looking Back in Time



Closeup images of the numbered objects in (a)



A portion of the constellation Hercules

- Older galaxies are around 11 billion yrs.
- They are small and blue.
- Add a number together would make a modern galaxy.

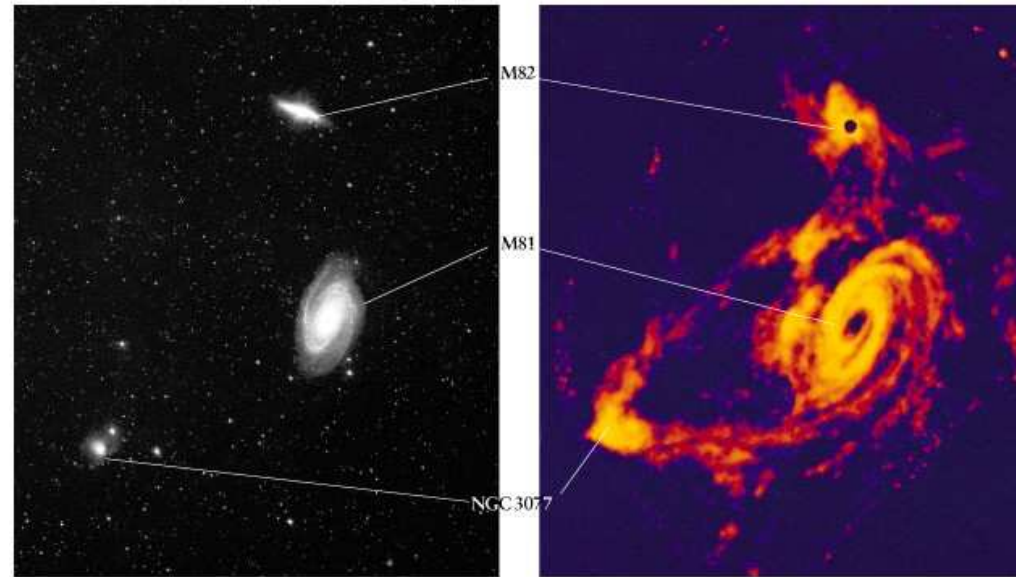
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*Any  
connection  
between  
these  
galaxies?*



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## *Collisions*



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## *Galaxies Collide*



NGC 2207 &  
IC 2163

NGC 7676  
“The Mice”



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## *Collisions*



- They do not involve colliding stars– but rather gravitational fields
- Might form hot intergalactic gas
- Could initiate rapid star formation - called *Starburst Galaxies*
- Collision causes stars to be scattered into “tails”
- Causes galaxy mergers called “galactic cannibalism”

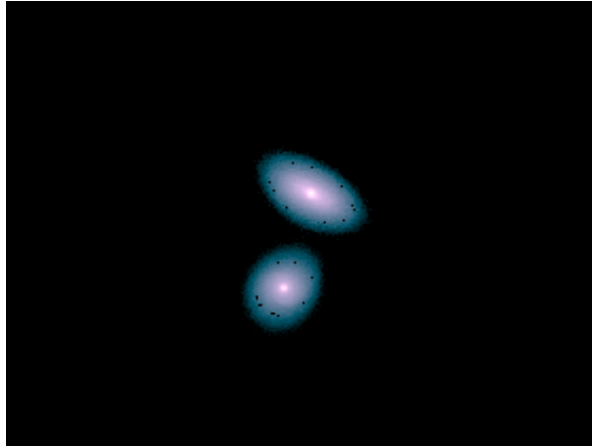
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# Galaxy Collisions

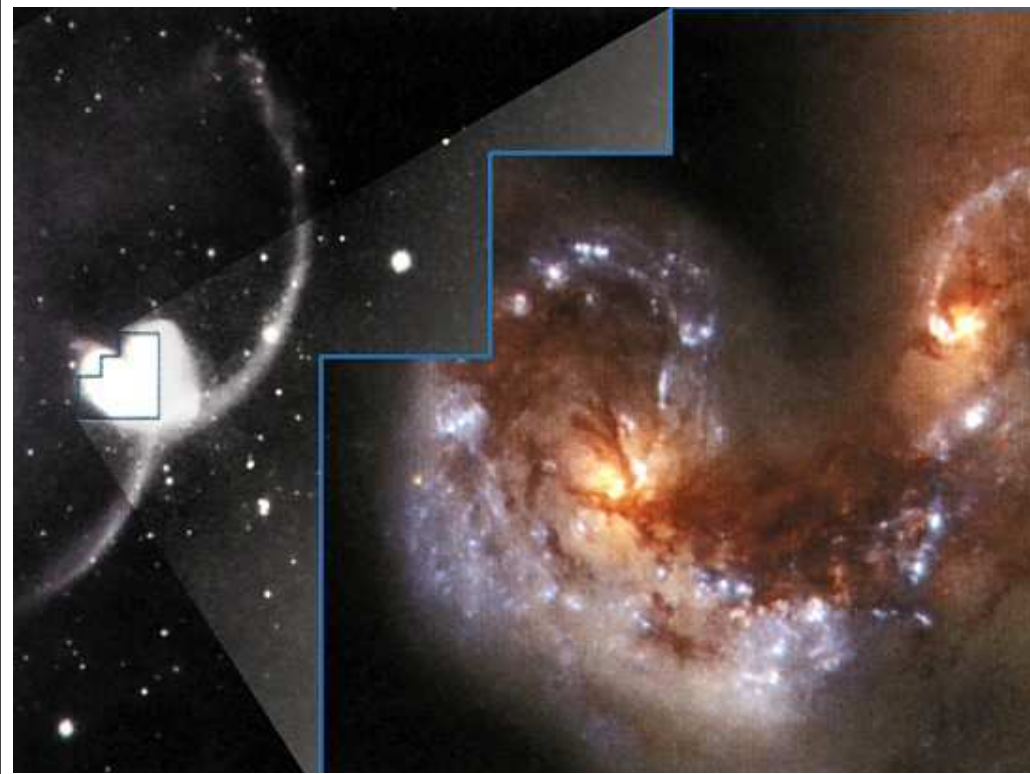


Computer simulation of two galaxies colliding by John Dubinski and Lars Hernquist



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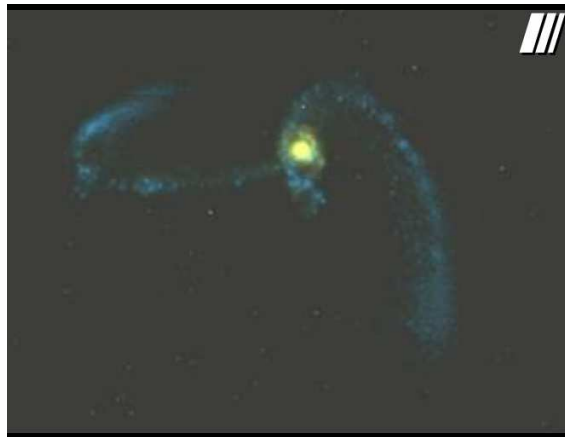
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## Galaxies Collide



When spiral galaxies collide, their bulges merge, while the disks are turned into *tidal tails*



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## Galaxies Collide



NGC 2207 & IC 2163

NGC 7676  
"The Mice"



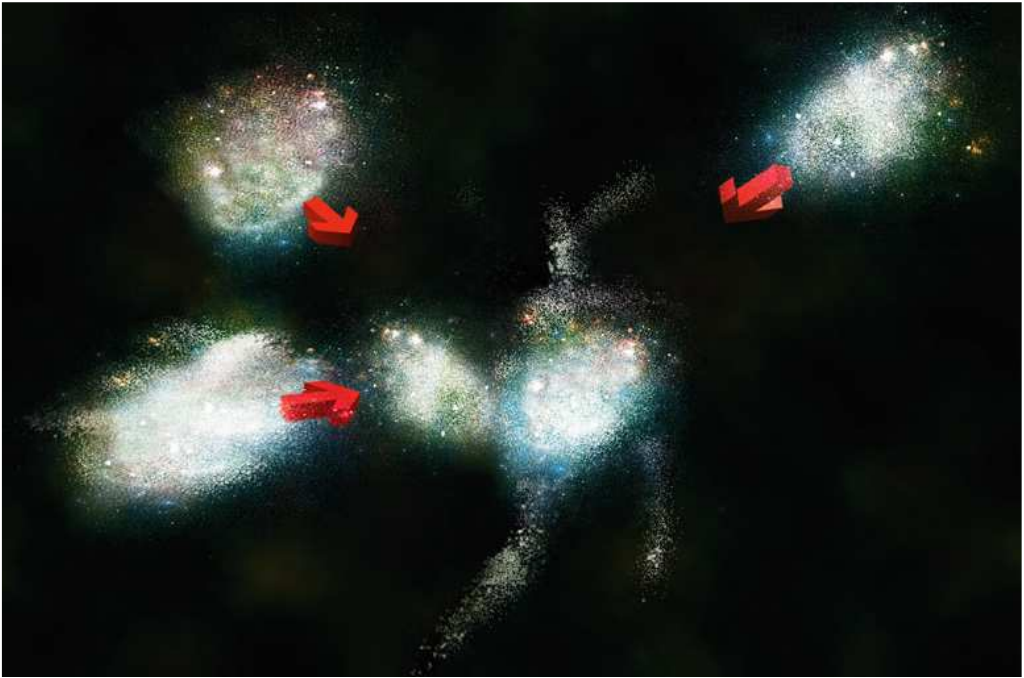
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# Multi-galaxy Collisions



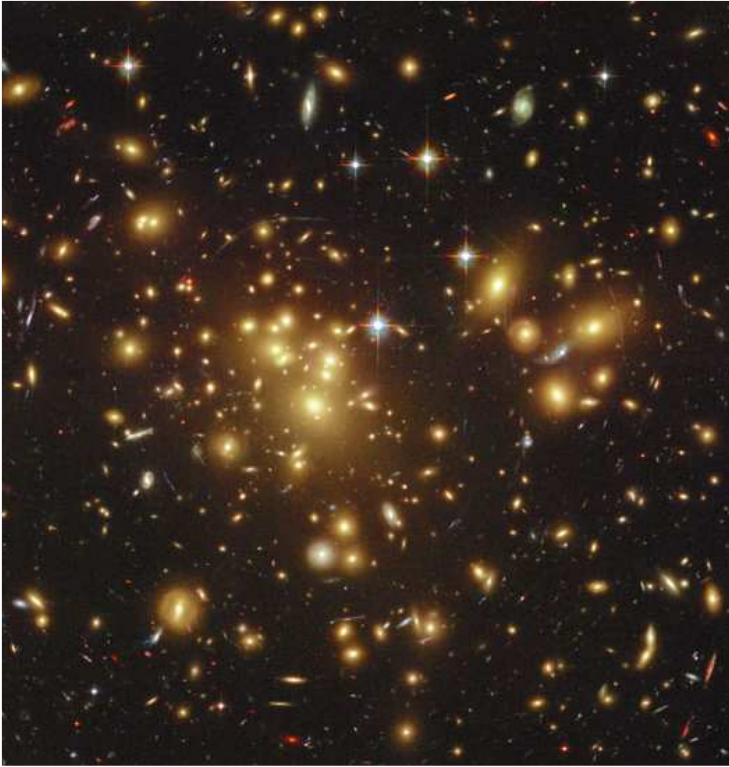
Collisions are also a factor in galaxy evolution!

Modeling such collisions on a computer shows that spiral galaxies can merge to make a giant elliptical

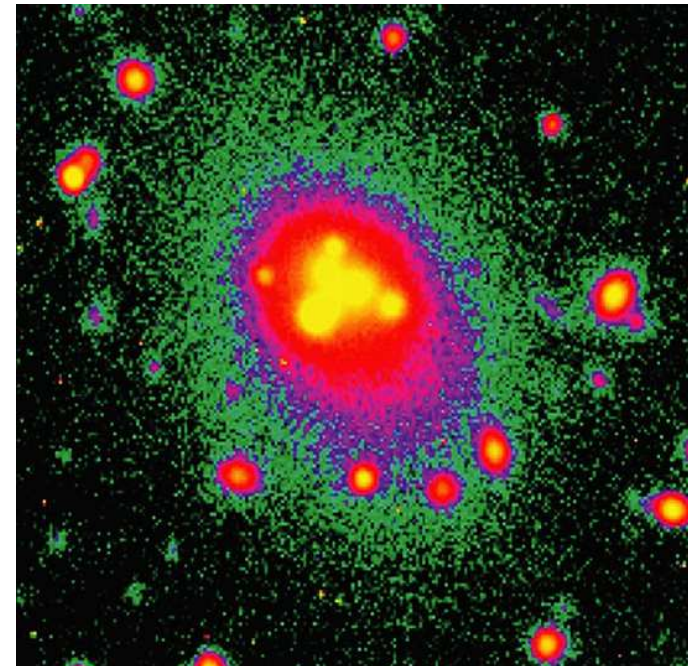


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Collisions may explain why elliptical galaxies tend to be found where galaxies are closer together



Giant elliptical galaxies at the centers of clusters seem to have consumed a number of smaller galaxies

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## Thought Question



When galaxies collide, what happens to the stars?  
Do they collide as well?

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## Why do galaxies differ?



- Angular momentum may determine size of disk
- Density of protogalactic cloud may determine how fast a galaxy forms
- Collisions shape galaxies early on
  - Mergers of small objects make halo & bulge
  - Mergers of larger objects make elliptical galaxies
- Relatively undisturbed galaxies can still have disks

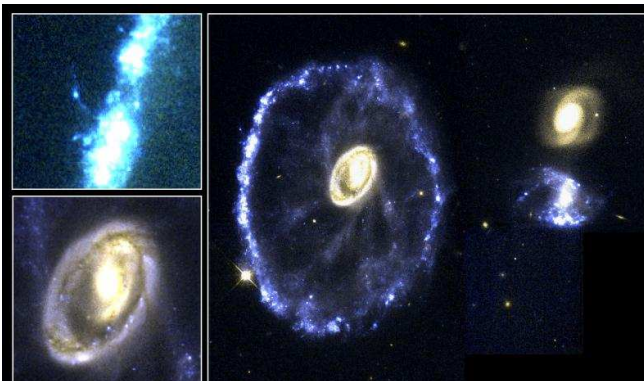
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## Starburst Galaxies



- Galaxies with enhanced rates of star formation
- Usually forming massive stars for a short period (few Myr).
- Probably due to collisions



Cartwheel Galaxy

HST · WFPC2

PR95-02 · ST ScI OPO · January 1995 · K. Borne (ST ScI), NASA

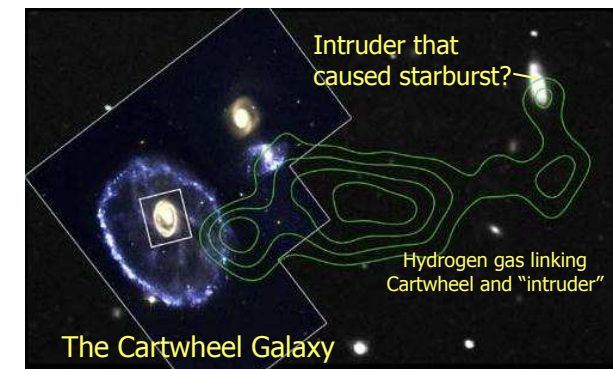
12/23/94 zgf

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## Starburst Galaxies



- Appear very blue, lots of young stars
- Gas is compressed by interactions, triggers galaxy-wide star formation



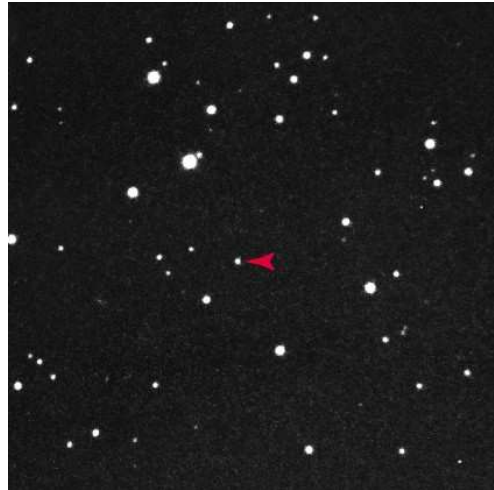
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# A Very Strange Star !?!



- Looked like a blue star, but had very odd spectrum lines
- Turned out it was simply greatly redshifted  $\Rightarrow z = 0.16$
- That's 2 billion light years away!
- It must be 100 times brighter than the entire Milky Way!
- **Not** a star



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# Quasars...

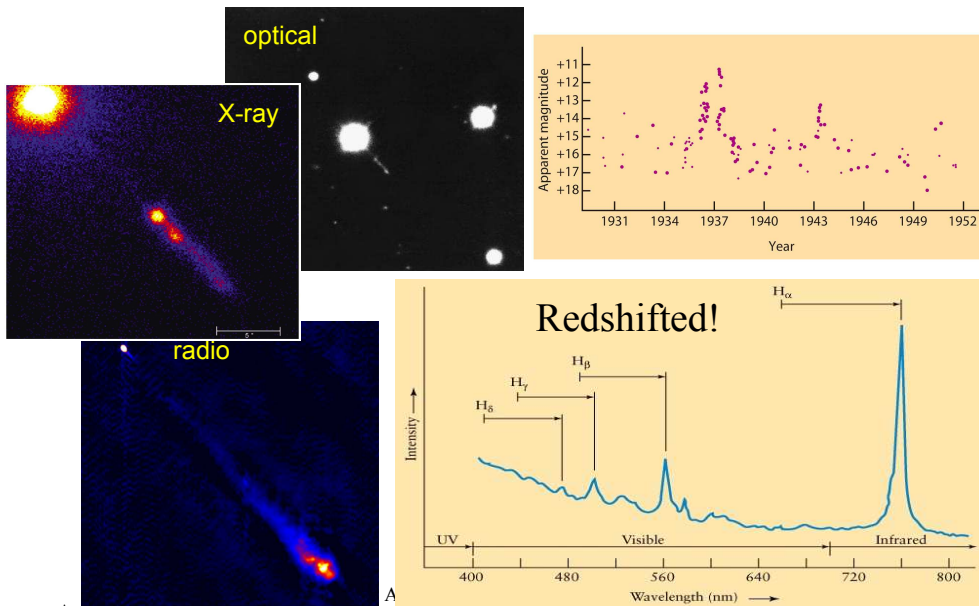


- These objects have a spectrum much like a dim star
  - But highly redshifted
  - Appear to moving away from us very fast!
- Dubbed **quasars** (quasi-stellar radio sources)
- Hubble's Law tells us that they are at "astronomical" distances
  - Up to 13 billion light years away!
- Great distances - must be very bright
  - Some 1 million times the brightness of our Galaxy!
- Highly variable
  - Must be small - about the size of our Solar System

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# The First Quasar Discovered: 3C 273

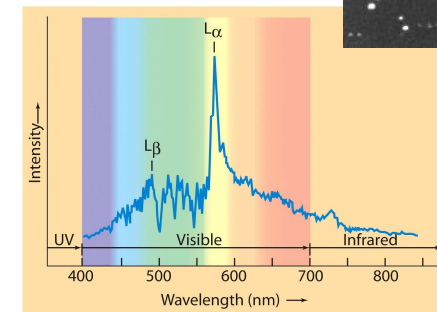
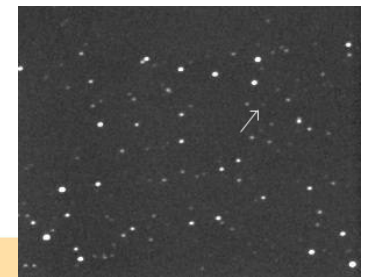


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# PKS 2000-330



- Redshifted so much that UV emission can be seen in the optical
- This quasar appears to be moving away from us at 92% of the speed of light!



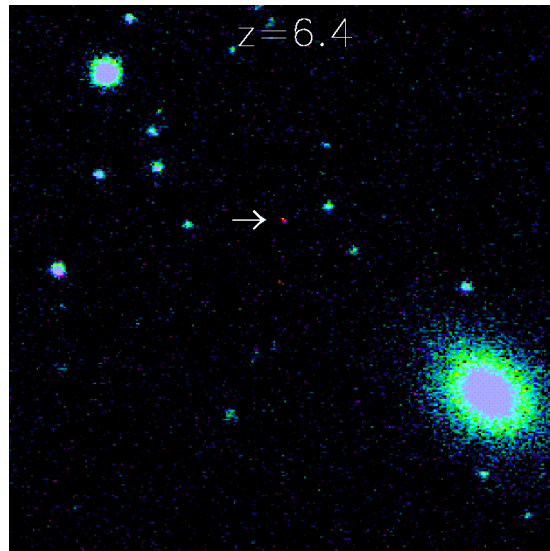
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## The Furthest Known Object



- Although there are now sources that may beat it, this galaxy is at  $z=6.4$ !
- That means only 800 Myrs after the Big Bang!
- Probably only current record holder.
- Seems to be moving at 95% the speed of light!



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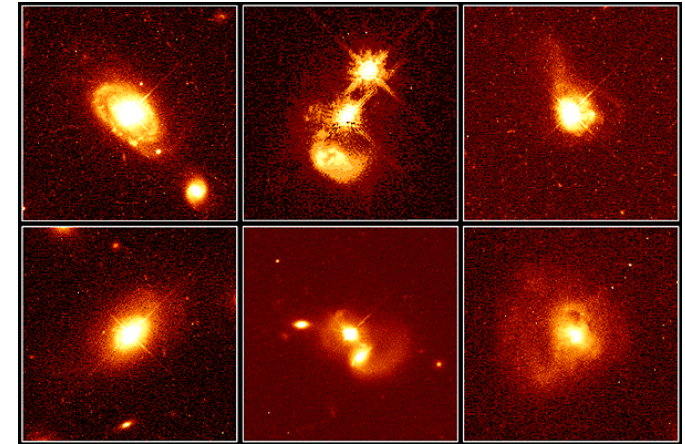
<http://www.sdss.org/news/releases/20030109.quasar.html>

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## Quasar Host Galaxies



- Quasars live in distant galaxies
- They are *galactic nuclei*!

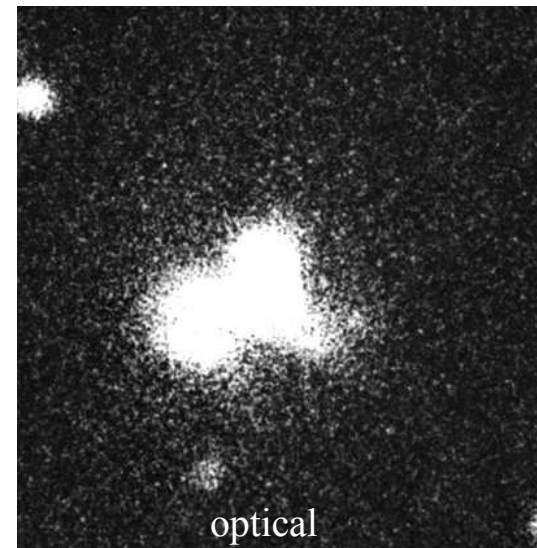


**Quasar Host Galaxies** HST • WFPC2  
PRC96-35a • ST ScI OPO • November 19, 1996  
J. Bahcall (Institute for Advanced Study), M. Disney (University of Wales) and NASA

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## Are there quasars in the nearby Universe?



- Looks like a star
- But bright in the radio
- And it's moving away from us fast!
- Moving away at 14,000 km/s.
- That's about 5% the speed of light!
- 635 million light years away! Or 194 Mpc.
- Similar to a quasar??

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# Active Galaxies



- There are no quasars in the nearby Universe *now*
- But there are some very energetic galaxies (about 1% of all galaxies)
  - Very bright, star-like nuclei
  - Often, energetic outflows of matter from the nucleus
- Called **active galaxies**
- Types of active galaxies
  - **Seyfert galaxies**
  - **Blazars**
  - **Radio galaxies**
- Like quasars, but not as energetic



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# Seyfert Galaxies



- Look like normal spiral galaxies, but with incredibly bright nuclei
- Potentially as bright as a trillion Suns!
- Brightness varies tremendously
- Over a few weeks it's brightness can change by the ENTIRE brightness of the Milky Way

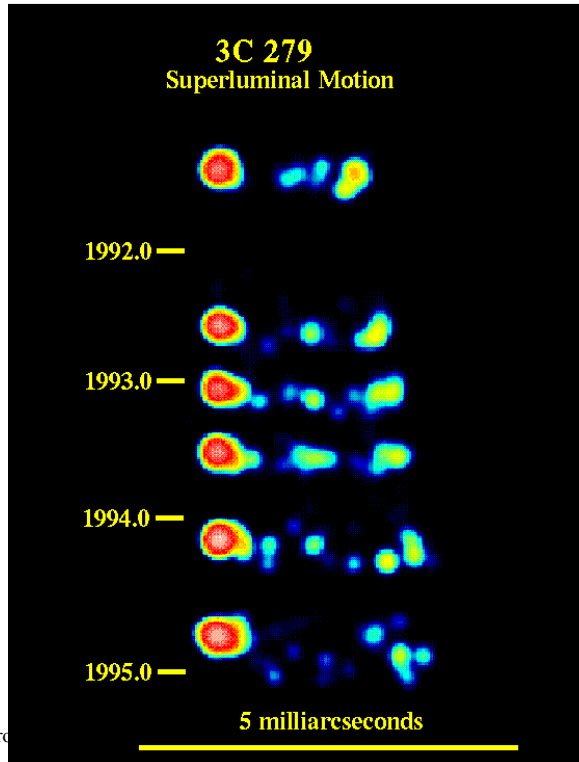
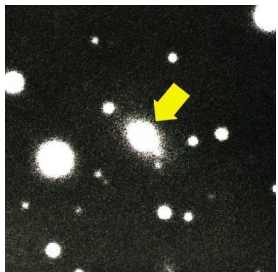


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# Blazars

- Bright nuclei with almost completely featureless spectrum.



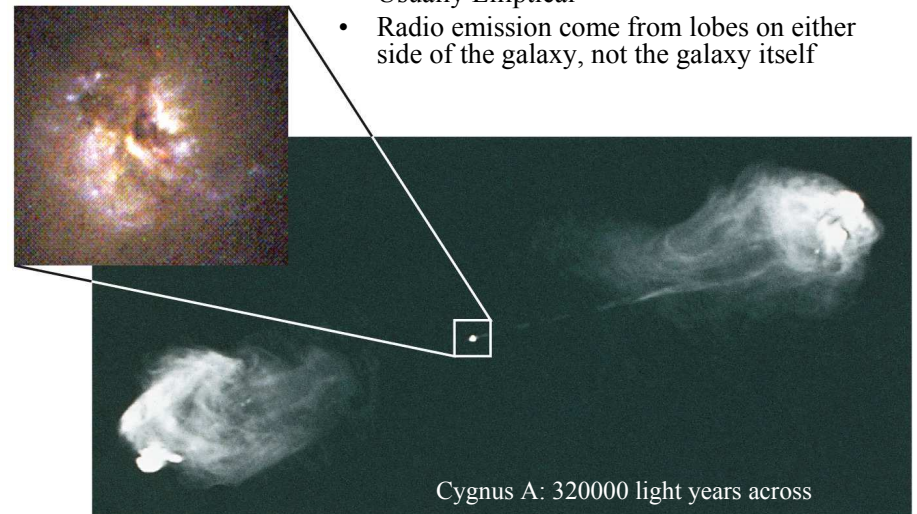
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# Radio Galaxies

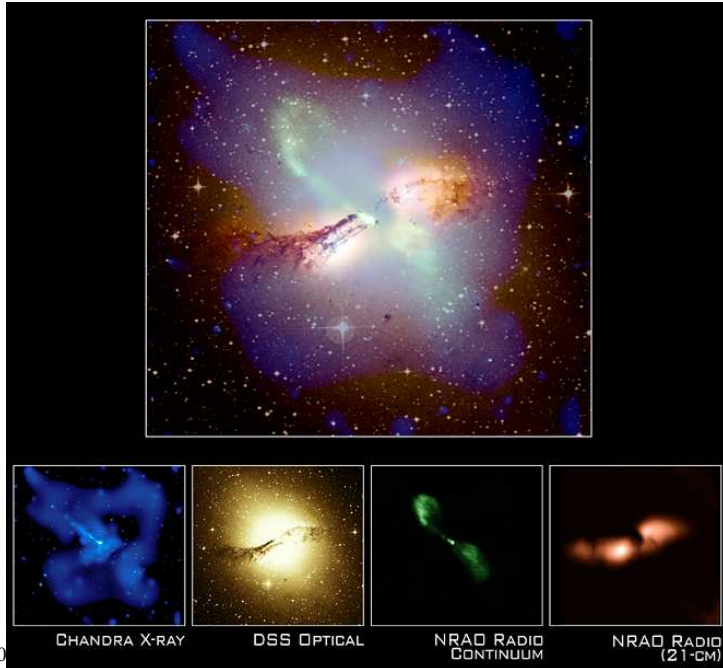


- Galaxies that emit large amounts of radio waves
- Usually Elliptical
- Radio emission come from lobes on either side of the galaxy, not the galaxy itself



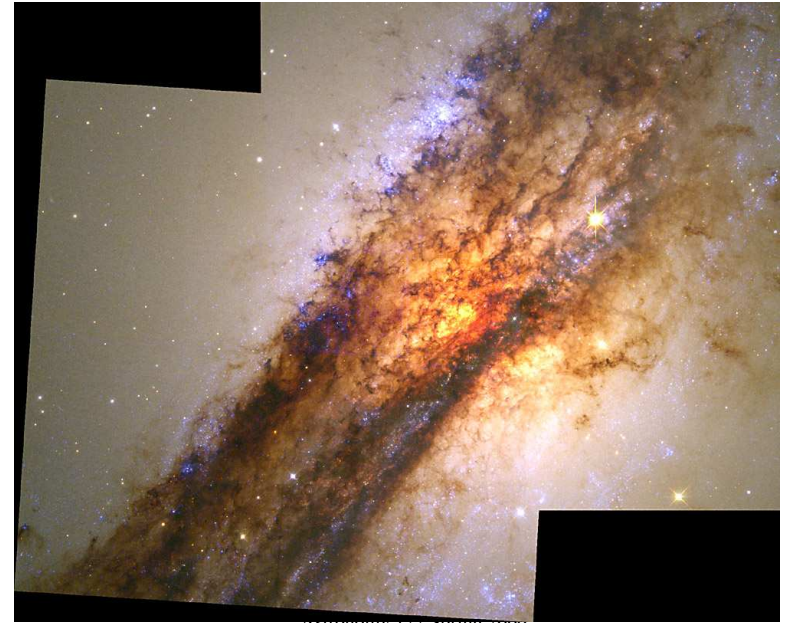
Cygnus A: 320000 light years across

# Radio Galaxies: Centaurus A



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# Centaurus A



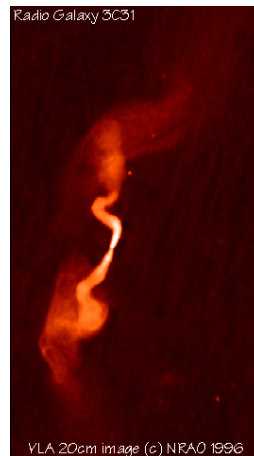
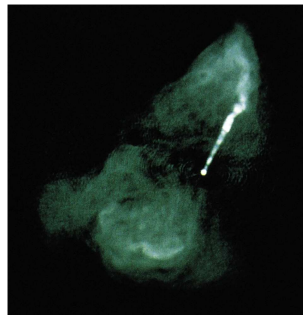
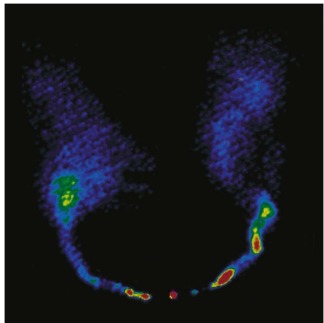
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# Radio Galaxies



- There are varying types of radio galaxies
- Called *radio loud* as they can be 10 million times as bright as the Milky Way at radio wavelengths

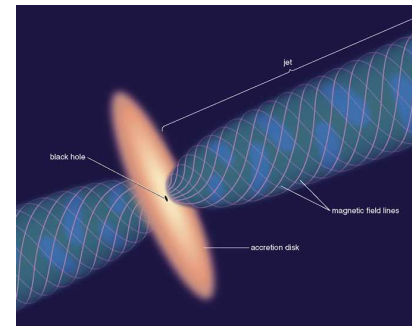
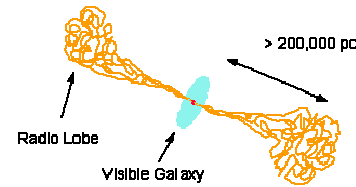


VLA 20cm image (c) NRAO 1996

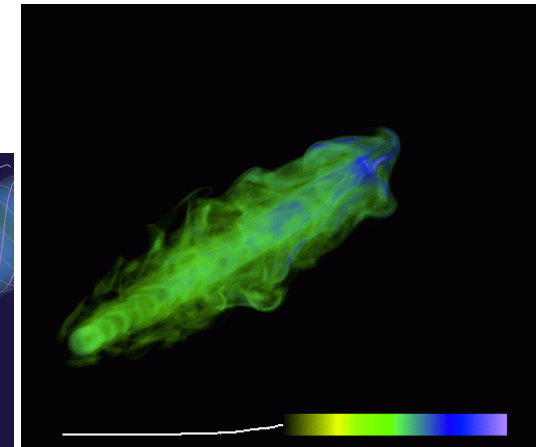
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# Radio Galaxy Jets



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# What is the power source for quasars and other active galaxies?

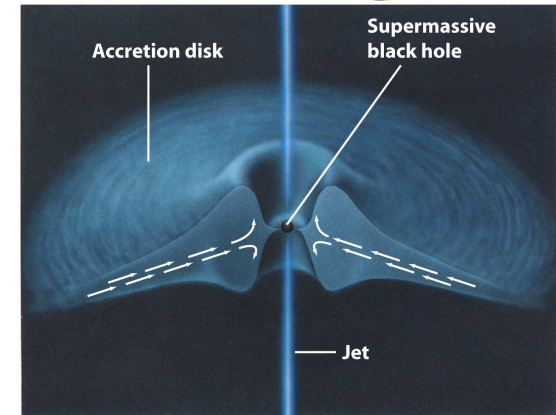
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# Driving Active Galaxies: The Monster Within



- A scary blue monster?
- Probably not
- More likely - the energy source is a supermassive black hole
- Accretion disk emits tremendous amounts of energy as it falls onto the black hole



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# The Central Engine



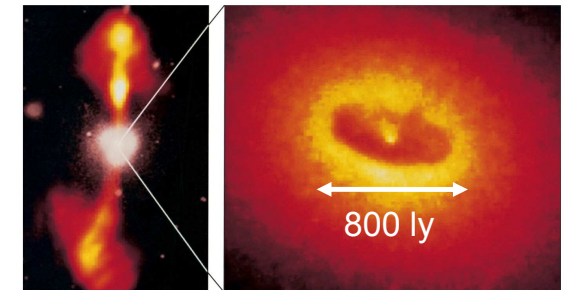
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# Supermassive Black Holes



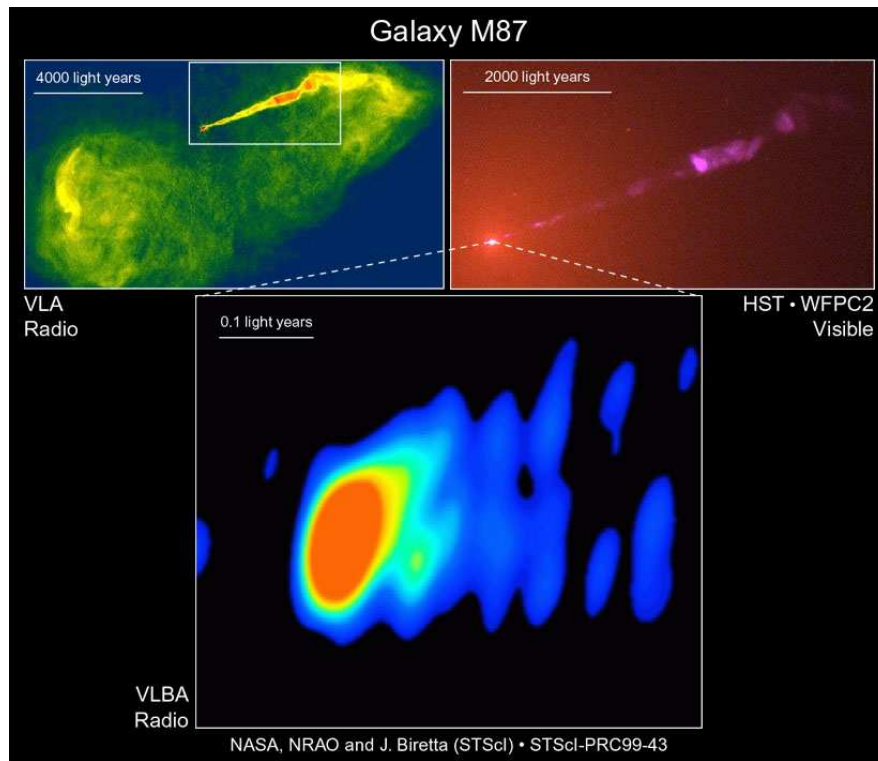
- Energy source for active galaxies
- Only thing compact enough and energetic enough
- Black holes > 1 billion solar masses
- Compression of material falling into black hole heats it up and forces some into jet



NGC 4261 in the Virgo Cluster

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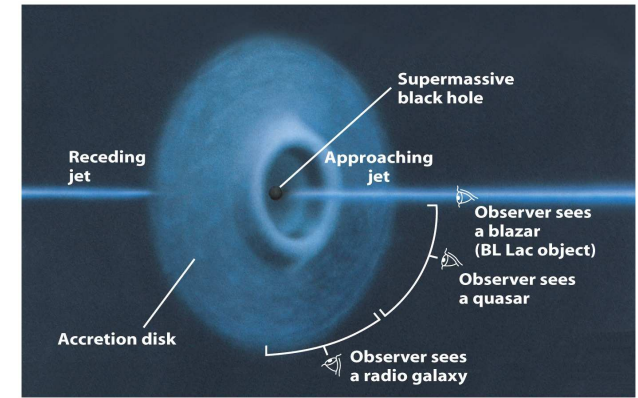
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## A Unified Model



- Active galaxies and quasars have the same energy source (supermassive black holes)
- Orientation matters!



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## Quasars and Active Galaxies



- Supermassive black holes probably exist in most if not all galaxies' cores
- In the past, active galaxies were more common than now
- Very powerful active galaxies were at one time quasars?
- As the Universe evolved, the quasars calmed down
  - Turned off?
  - Became today's active galaxies?

**TABLE 16-1** Galaxy and Quasar Luminosities

Object	Luminosity (watts)
Sun	$4 \times 10^{26}$
Milky Way Galaxy	$10^{37}$
Seyfert galaxies	$10^{36} - 10^{38}$
Radio galaxies	$10^{36} - 10^{38}$
Quasars	$10^{38} - 10^{42}$

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## Think-Pair-Share



If our Galaxy's supermassive black hole were fed, would it turn into an active galaxy?

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