



## This Class (Lecture 23):

Galaxies

**HW9 due on Friday.**

## Next Class:

Active Galaxies & Quasars

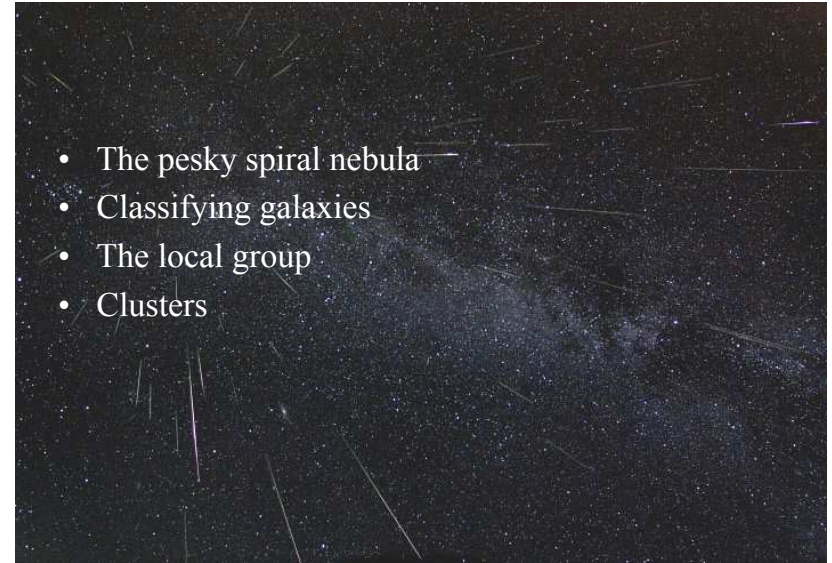
Music: *Galaxies* – Laura Viers

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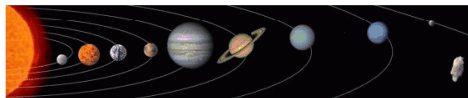


- The pesky spiral nebula
- Classifying galaxies
- The local group
- Clusters



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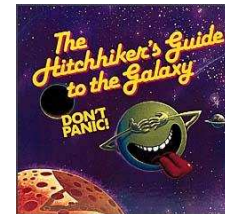


## *Astronomy: The Big Picture* *Moving from our Galaxy outward!*



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## *Space is Big!*



“Space is big. Really big. You just won't believe how vastly hugely mind-bogglingly big it is. I mean, you may think it's a long way down the road to the chemist, but that's just peanuts to space...

To be fair though, when confronted by the sheer enormity of the distances between the stars, better minds than [ours] have faltered.

The simple truth is that interstellar distances will not fit into the human imagination.”

--Douglas Adams

*The Hitchhiker's Guide to the Galaxy*

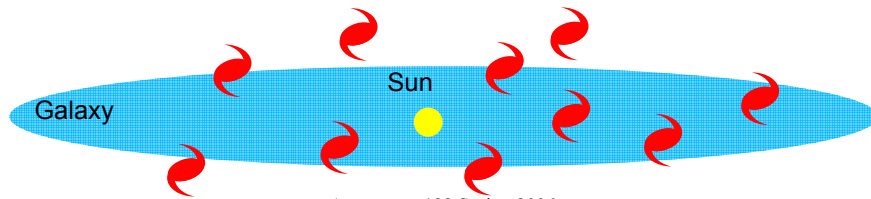
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## Those weird Spiral Nebulae?



- Dim, diffuse, “interstellar” nebulae with spiral structure were seen in the 17<sup>th</sup> century.
- Some disagreement on what they were.
  - Kant: Our galaxy is a spiral “island universe” and the other spiral nebulae are the same and far away
  - Herschel and others: Milky Way is all there is in the Universe, and the spiral nebulae are nearby. More prevalent idea.



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## “Spiral Nebulae”



- Dim, diffuse “nebulae” with spiral patterns
- Spiral structures catalogued mid-1800s by Lord Rosse (Ireland)



Rosse's M51 sketch



“Leviathan”  
1.8 m telescope

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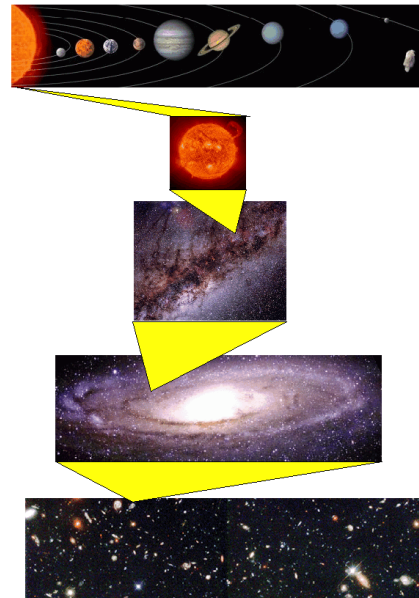
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## One of



We are:

- 1 planet out of 9 in our solar system.
- 1 stellar system of 100 billion stars in our Milky Way
- 1 galaxy of the 100 billion galaxies in the observable Universe.



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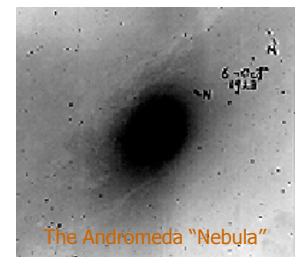
## Edwin Hubble



- In 1923, Hubble resolved M31, the Andromeda “Nebula”, into stars
- If these stars were like the stars in our Galaxy, then M31 must be far away!
- Estimated the distance to M31 to be 300,000 parsecs (modern estimate is 700,000)
- Andromeda is an “island universe” like our own Galaxy.



Hubble at Mt. Wilson  
Observatory



The Andromeda “Nebula”

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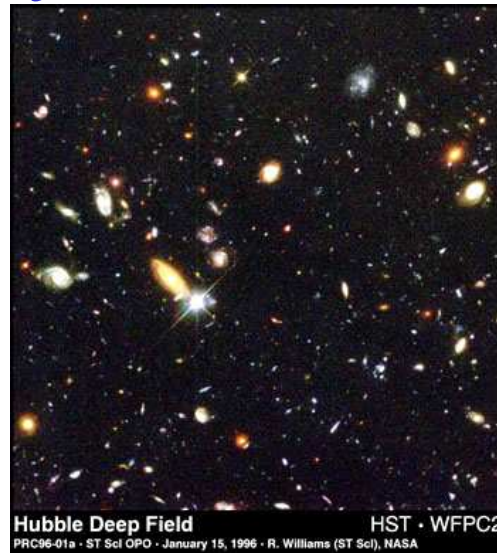
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# Galaxies – Fundamental “Ecosystems” of the Universe



- Galaxies “fill” universe.
- Typical separation  
 **$\sim 10^6$  pc or 1 Mpc!**
- Most distance is 1000's of Mpc away
- Galaxies are huge masses of stars
- Range in size from large (MW-like) to small “Dwarf”
  - 1 billion to 100's of billions of stars



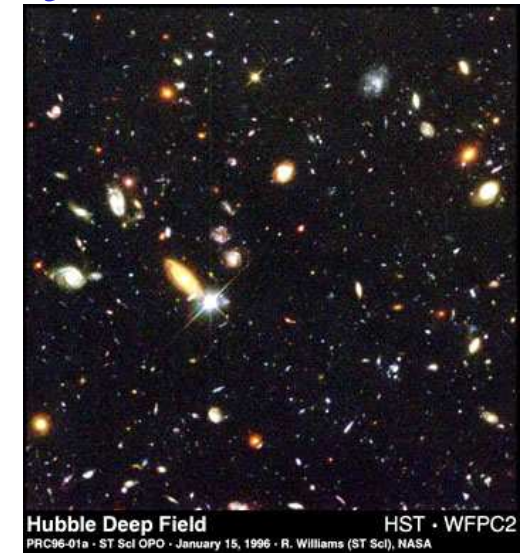
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# Galaxies – Fundamental “Ecosystems” of the Universe



- Galaxies are the cosmic engines that turn gas and recycles the gas the stars eject back
- In between no star formation occurs – “nothing happens” in intergalactic space.

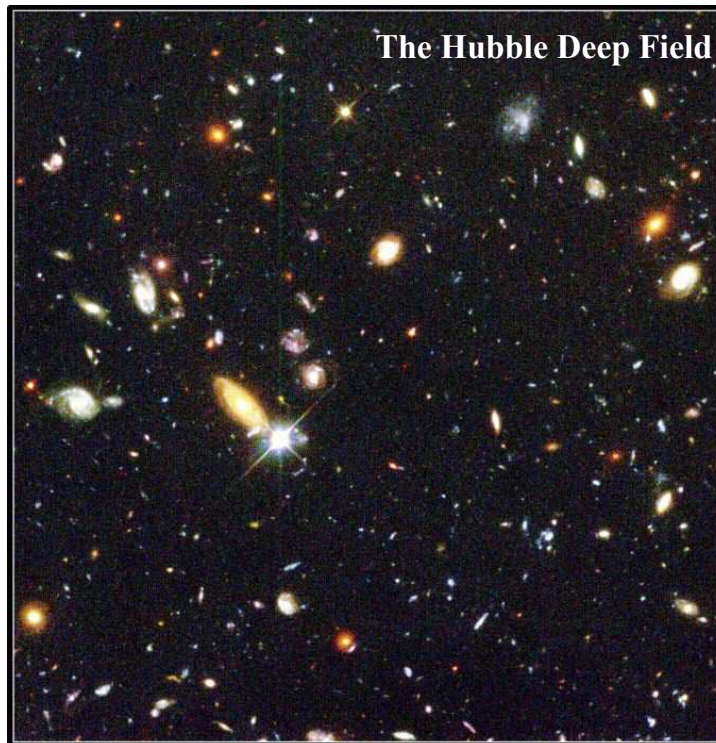


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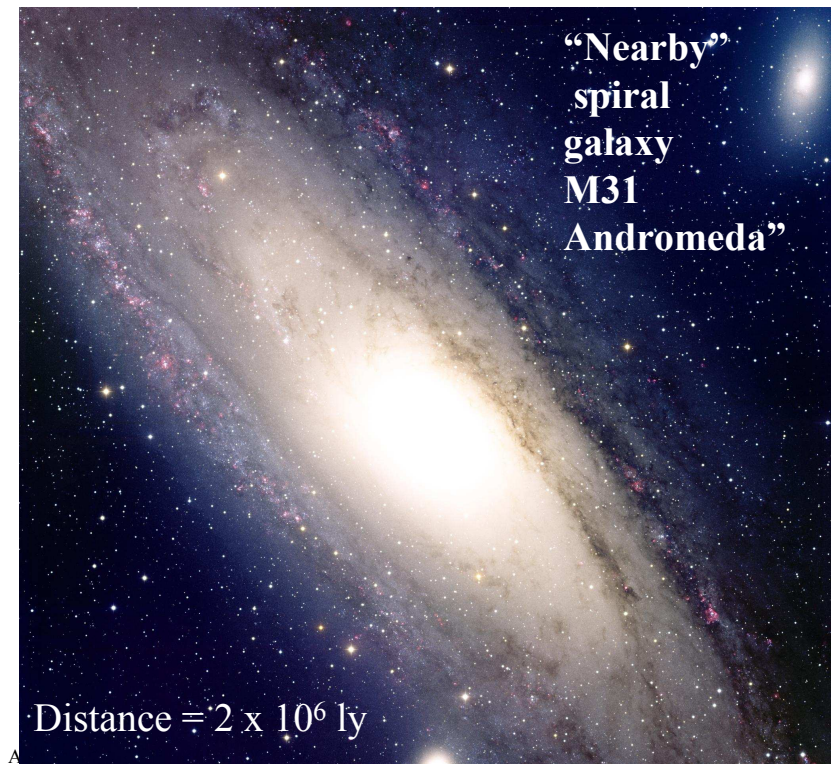
## Distant galaxies:

- The deepest optical image of a patch of sky
- Like looking back in time ...
- Galaxies as they were, 1 to 10 billion years ago.

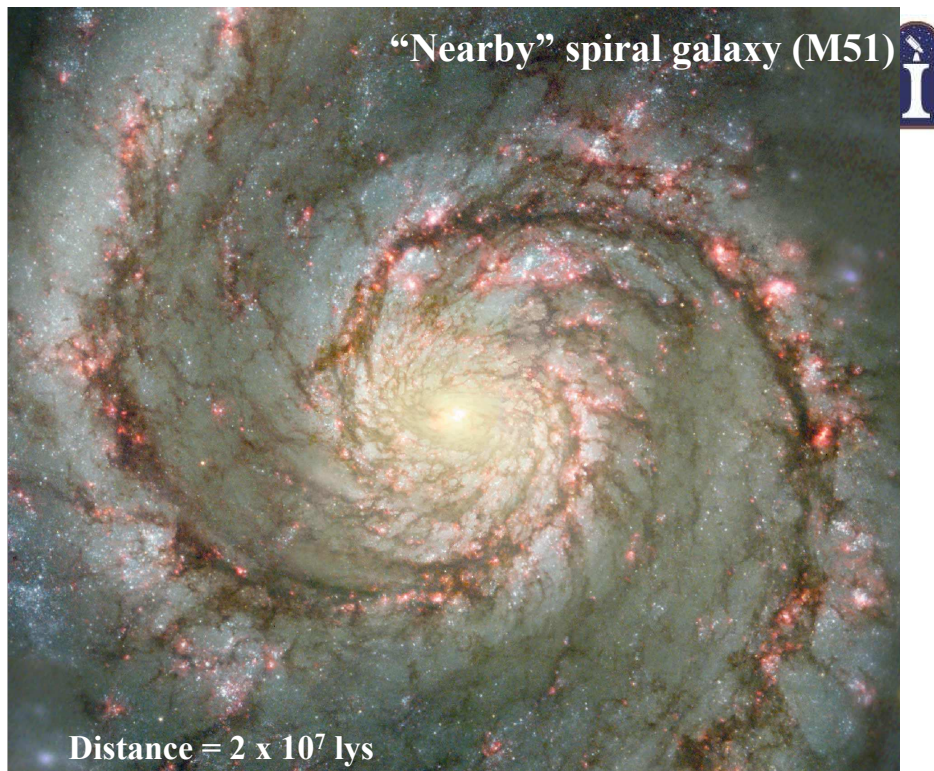


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## “Nearby” spiral galaxy M31 Andromeda”



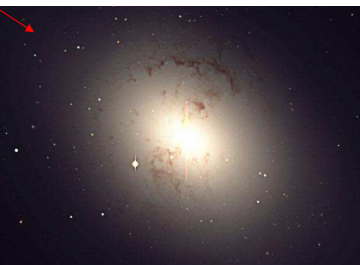




## Galaxies are the Fundamental “Ecosystems” of the Universe

### Three Main Types of Galaxies:

- Spirals (77%)
- Ellipticals (20%)
- Irregulars (3%)



## *Thought Question*

What does the color of the galaxy tell us about the age of the stellar population?

# Classes of Galaxies



- **Spirals (S)**
  - Basic structure: disk and bulge
  - Medium to large galaxies
  - The disk has the young blue stars, while the bulge has older red stars
- **Ellipticals (E)**
  - Pure bulge, no disk component
  - Large range in sizes
  - All older red stars
- **Irregulars (Ir)**
  - Well... odd, irregular structure
  - Smaller galaxies
  - Mostly young blue stars



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# Galaxy Types: Overview



	Spirals	Ellipticals	Irregulars
<b>Mass (<math>M_{\odot}</math>)</b>	$10^9 - 10^{12}$	$10^5 - 10^{13}$	$10^8 - 10^{10}$
<b>Luminosity (<math>L_{\odot}</math>)</b>	$10^8 - 10^{10}$	$10^5 - 10^{11}$	$10^7 - 10^9$
<b>Diameter (kpc)</b>	5 - 200	1 - 200	1 - 10
<b>Color</b>	<b>Disk:</b> bluish-white <b>Bulge:</b> reddish - yellow	Reddish-yellow	Bluish-white

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## Spiral Galaxies



- Spirals are classified on the amount of bulge component (and how tightly the arms are wound)
- These are designated as Sa, Sb, Sc, in order of decreasing bulge



More bulge and tightly wound



Sb



Sc

More disk and loosely wound

More disk means more ongoing star formation!

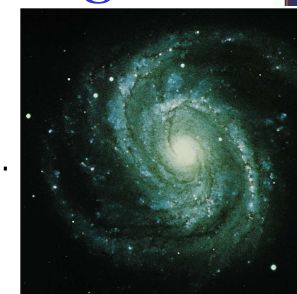
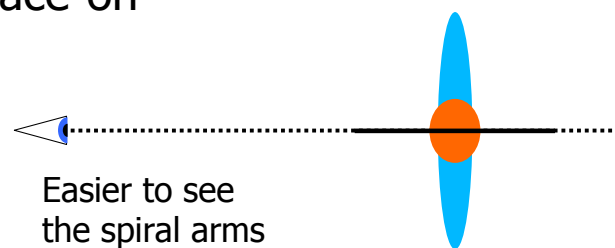
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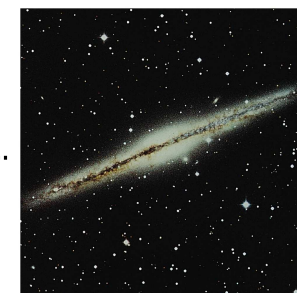
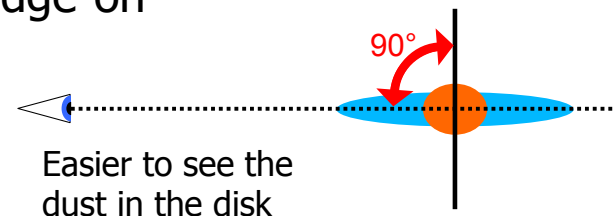
## Effect of Viewing Angle



Face-on



Edge-on



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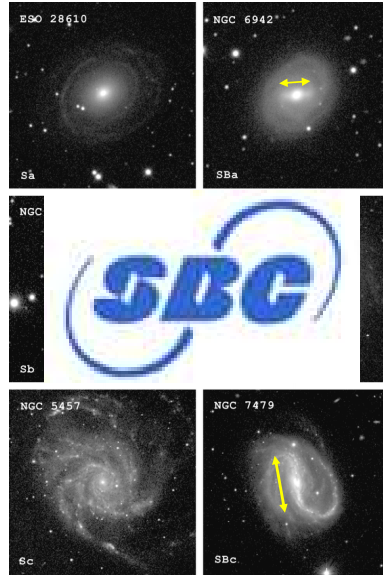
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## Barred Spirals



- About 20% of spiral are *barred spirals*
- The spiral arms branch off from a straight bar of stars that passes through the central bulge
- They are designated with an “SB” rather than the usual “S” for spiral galaxies
- The classes of barred spirals are SBa, SBb, and SBc



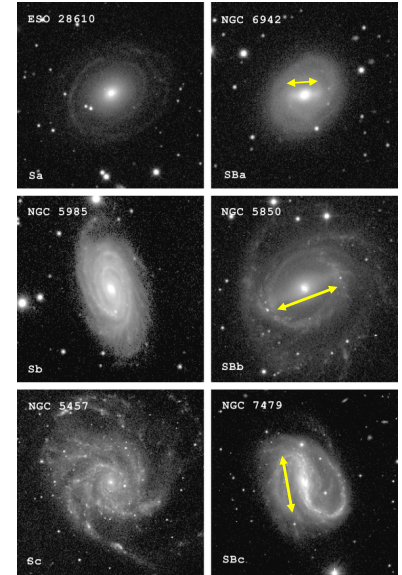
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## Why do we see Spiral Arms?



- They are easily seen as the arms contain numerous bright O and B stars that illuminate dust in the arms
- However, stars overall are evenly distributed throughout the disk



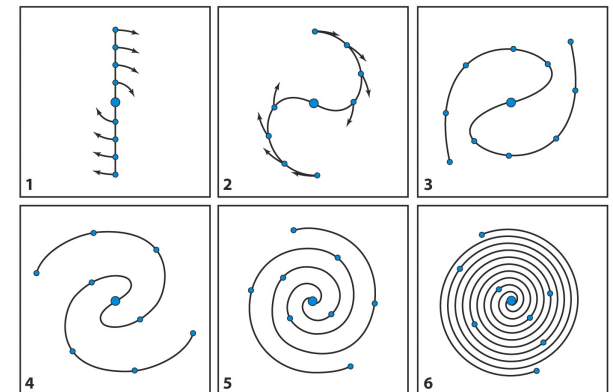
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## The Winding Problem



- If the arms are stationary, they should wind up and disappear
- This is not observed
- Spiral arms are **not** a permanent collection of stars
- Star, gas, and dust pass *through* the spiral arms



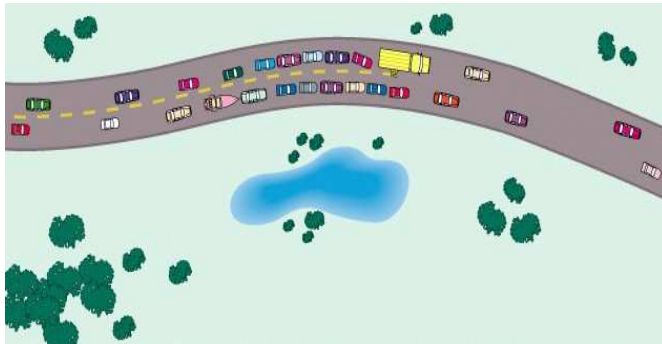
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## Density Waves



- Spiral arms are caused by waves in the gas and dust
  - Make the gas clump up
  - Like an interstellar traffic jam
- Increased density of gas and dust sparks formation of new O- and B-type stars that light up the spiral arm



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## Elliptical Galaxies



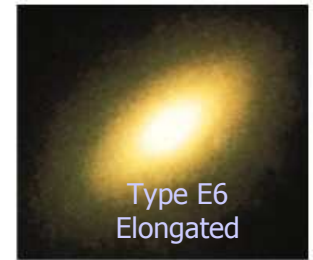
- Like a spiral galaxy's bulge
  - Mostly old, redder stars, little gas and dust
  - No disk organization, stars on random orbits
- Classified by how elliptical they appear
  - E0 (spherical) to E7 (elongated)



E0 M105



E3 NGC 4365



E6 NGC 3377

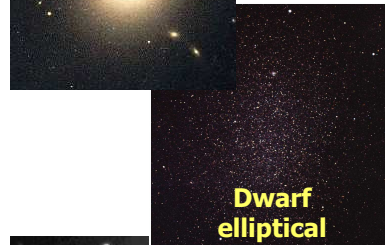
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## Varieties of Elliptical Galaxies



- Ellipticals come in a great range of masses
- The largest are *giant ellipticals*
  - Up to 100+ times more massive than the Milky Way
- The smallest are the *dwarf ellipticals*
  - 10,000 to a million times less massive than the Milky Way
  - Some only a few times larger than a globular cluster!
- Of course, there are also “garden variety” ellipticals
  - About 100 times smaller than to equal in size to the Milky Way



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## Irregular Galaxies



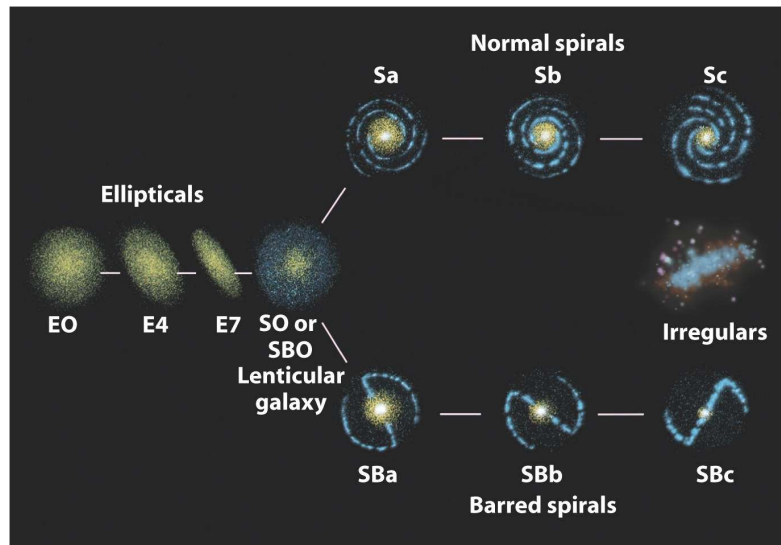
- Chaotic systems of stars
- Prominent examples: The Magellanic Clouds
  - Two of the Milky Way's satellite galaxies
- Generally smaller galaxies
  - Thousands to tens of times smaller than the Milky Way
- Chaotic systems of stars
  - No disk, no elliptical structure
- Dominated by young, blue stars



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# Hubble's "Tuning Fork" Classification Scheme



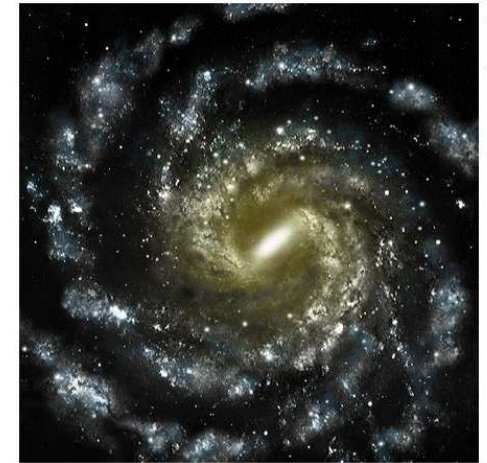
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# What Type of Galaxy is the Milky Way?



- The Milky Way is a spiral galaxy
  - Probably type Sb
- It might be a barred spiral!
  - Type SBb?



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# Measure other Galaxies Rotation Curves



- What do you think we find?
  - No other galaxies have dark matter
  - Some other galaxies have dark matter
  - All other galaxies have dark matter

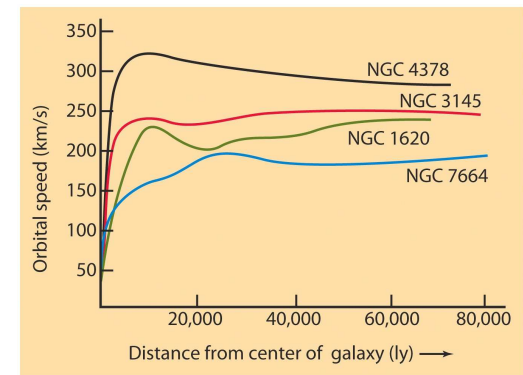
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# Masses of Galaxies



- As with the Milky Way, we measure the speed of a galaxy's rotation
- Like the Milky Way, other galaxies have a *flat rotation curve*
- Indicates a halo of **dark matter**
- We aren't special that way either.**

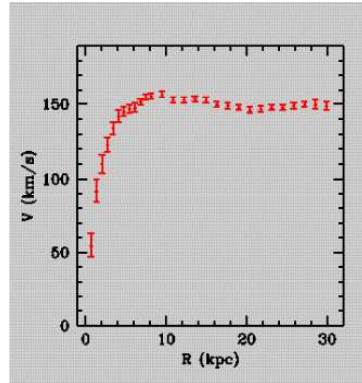
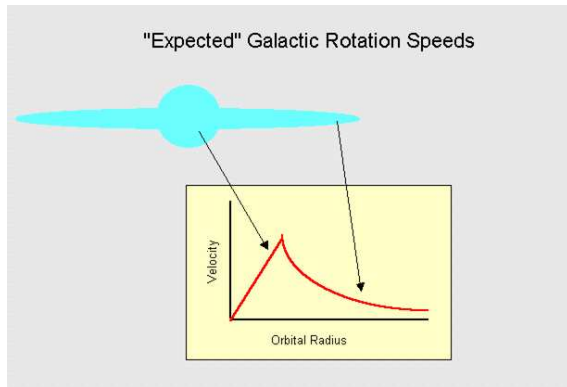


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# Galaxy Rotation Curve



Where's the matter?

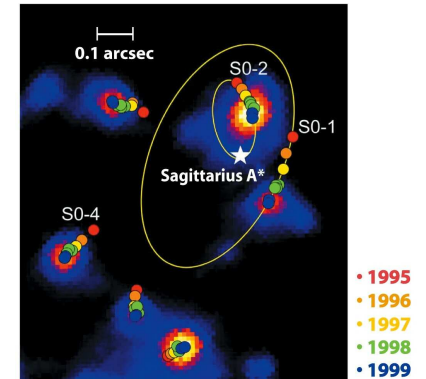
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# Galactic Nuclei



- Rapid orbits of radio sources around the Milky Way's center indicate a 2.5 million solar mass black hole at its nucleus!
- Do other galaxies show evidence for such supermassive black holes as well?



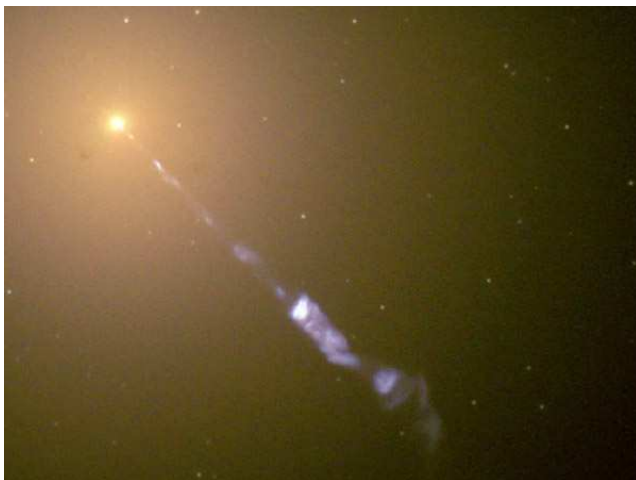
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# M87's Central Monster



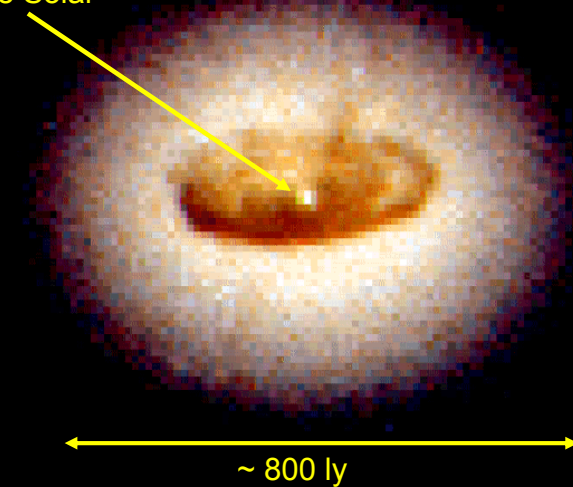
- Jet of M87
- 5000 light-year blowtorch!
- Probably from the disk of the black hole at the center
- 3 billion solar masses!



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1.2 billion solar masses within region the size of the Solar System



Core of Galaxy NGC4261

HST · WFPC2

PRC95-47 · ST ScI OPO · December 4, 1995  
H. Ford and L. Ferrarese (JHU), NASA





## The Local Group



- Our Galaxy is in a poor, irregular cluster
- Called the **Local Group**
- Dominated by two large spirals
  - The Milky Way
  - The Andromeda Galaxy (M31)
- About 40 smaller galaxies
  - Some satellites of the big two
  - M33 (small spiral)
  - Lots of dwarfs ellipticals and irregulars



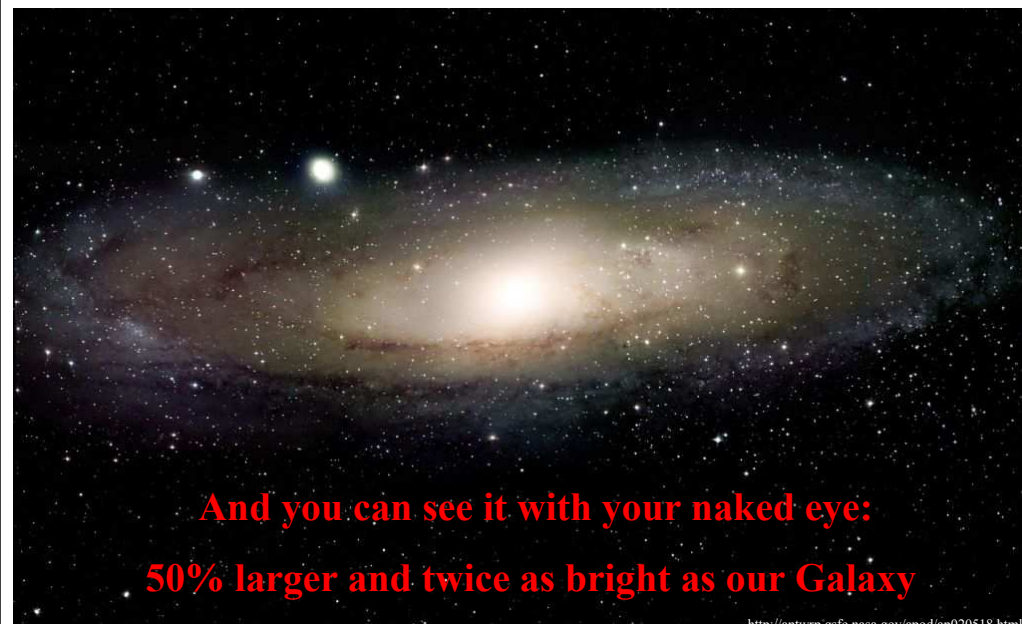
Triangulum (M33)

Local Group dwarf galaxies

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## The Andromeda Galaxy



<http://antwrp.esfc.nasa.gov/apod/ap020518.html>

## Fate of the Milky Way: It's coming right for us!



- What will happen to the Milkyway?
  - It will continue to grow as it cannibalizes the smaller orbiting galaxies.
  - The Andromeda galaxy is on a collision course.
  - Eventually (billions of years) we will end up a combined galaxy.
  - Probably look like an elliptical galaxy.



<http://www.seds.org/messier/small/m87.gif>

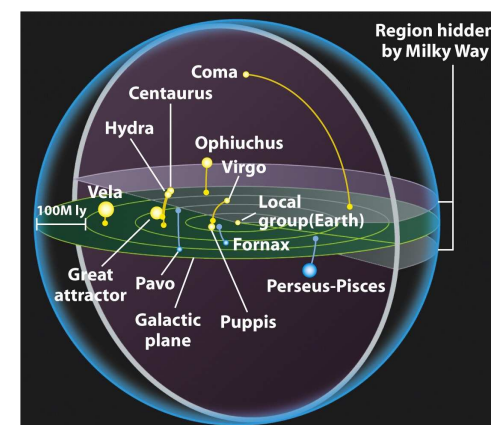
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## Galaxies Are Not Alone



- Galaxies are **not** scattered randomly throughout the Universe
- Galaxies are found in **clusters**
- Like clusters of stars, clusters of galaxies come in a wide variety
  - Poor or rich?
    - Dozens or thousands of members?
  - Regular or irregular?
    - Is the cluster concentrated towards the center?

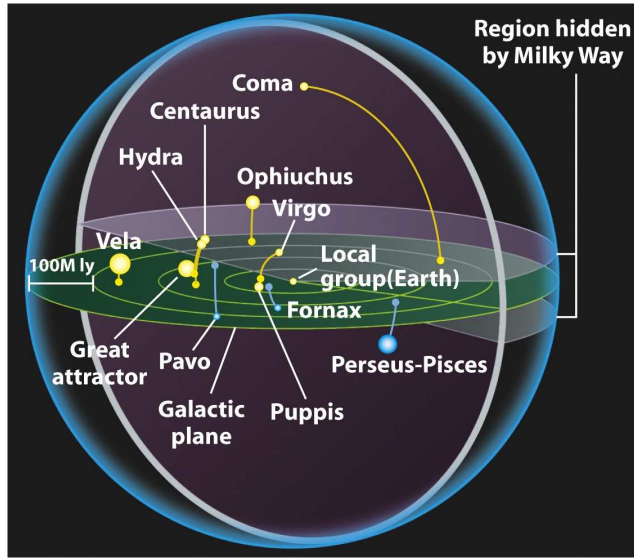


800 Mly sphere, centered on Earth

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## Nearby Clusters



800 Mly sphere, centered on Earth

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## Virgo Cluster

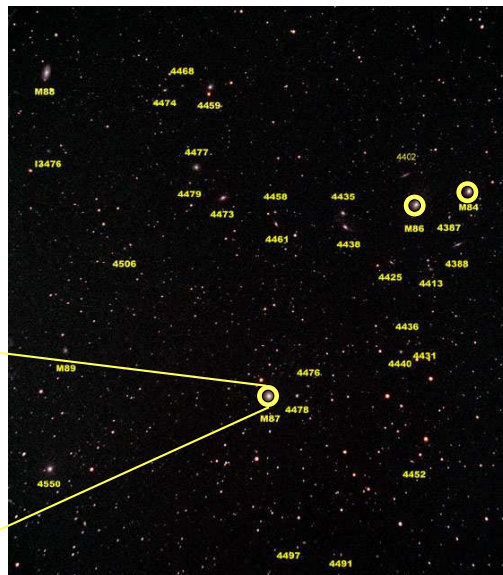


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## The Virgo Cluster



- More than 1,000 galaxies
- 16 Mpc away from the Milky Way
- About 4 Mpc across
- Dominated by three giant elliptical galaxies
- Our cluster is headed right for it.



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## Fornax Cluster



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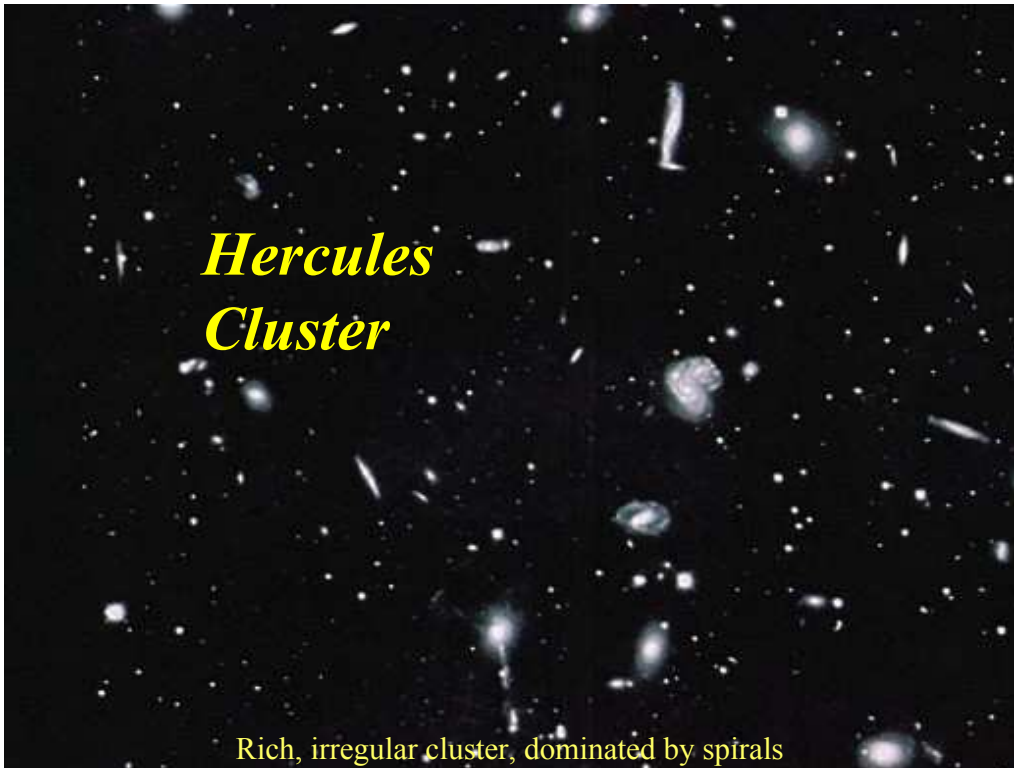
## *The Coma Cluster*

- Rich, regular cluster
- 90 Mpc = 300 million lyrs.
- Over 2000 galaxies.
- Dominates by two ellipticals



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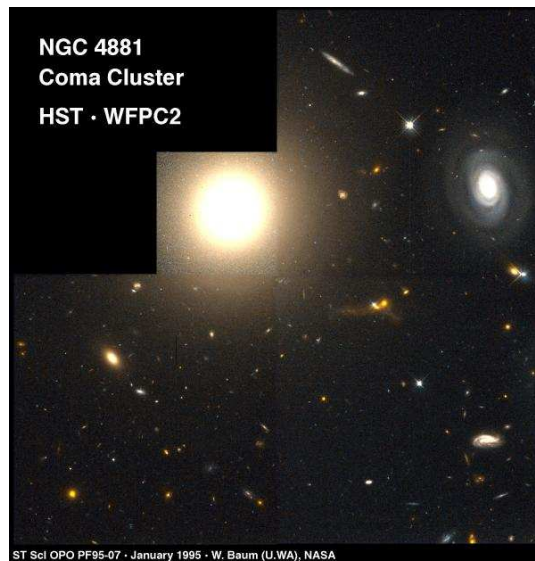
## *Hercules Cluster*

Rich, irregular cluster, dominated by spirals

## *Coma Cluster*



- A zoom near one of the ellipticals
- Contains many spirals



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