



## This Class (Lecture 19):

Black Holes

## Next Class:

The Milkyway

**Last Nightlab tonight!**

**HW7 due on Friday.**

**Mar Lecture report due in discussion class on April 5<sup>th</sup>.**

**Nightlab report due in discussion class on April 12<sup>th</sup>.**

**Nightlab makeup assignment available online (see assignments), also due on April 12<sup>th</sup>.**

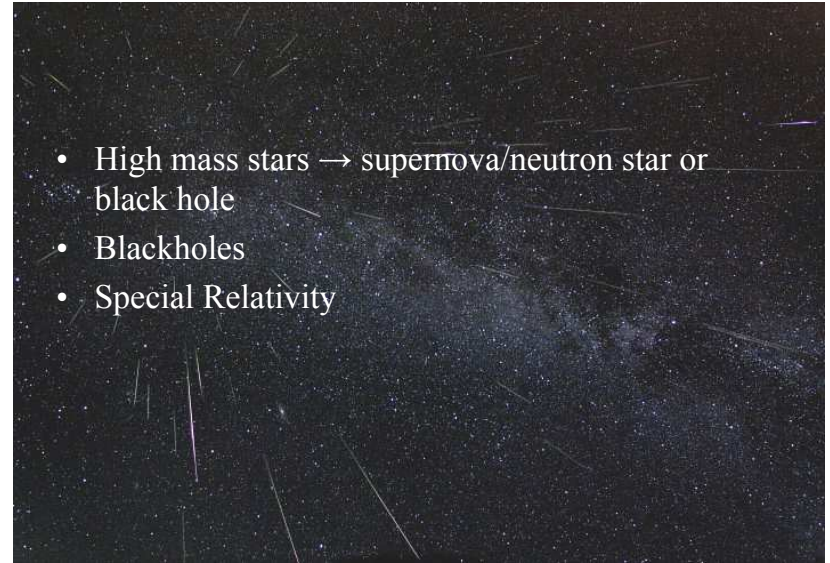
Music: *Black Hole Sun* – Soundgarden

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- High mass stars → supernova/neutron star or black hole
- Blackholes
- Special Relativity



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## Death throes



- What triggers a supernova?
  - Hydrostatic equilibrium is lost, gravity wins
  - Iron core with  $M > M_{\text{Chandra}}$
- What happens?
  - Quick core collapse overcoming electron degeneracy pressure.
  - Outer layers rebound off the core, explosion of envelope
- What are end products?
  - Enriched ejecta and compact neutron star (if core mass  $< 3$  solar masses)

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## Supernova!



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# Supernova Leftovers

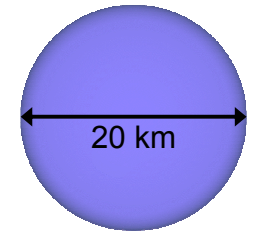
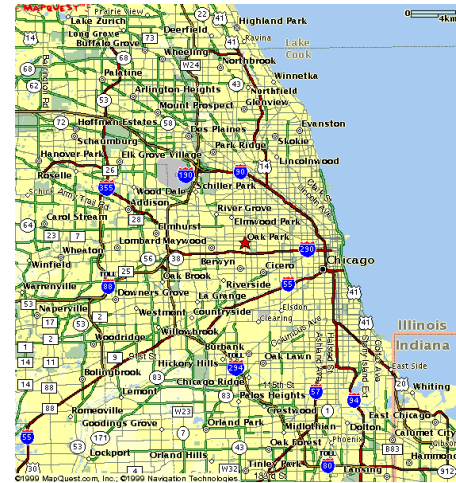


- What's left of the star's core after a massive star supernova?
- A **neutron star**
  - About 1.4 – 3 solar masses
  - Very small diameter – around 20 km!
  - Composed of a sea of neutrons
    - Supported by *neutron degeneracy pressure*!
    - Teaspoon of neutron star material on Earth would weigh almost 1 billion tons!!!!
  - Surface gravity – 200 billion times that on Earth
  - Escape velocity – half the speed of light

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# Relative Sizes of Stellar Corpses



Neutron star

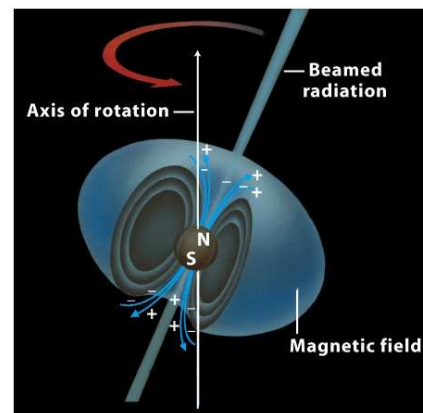
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# What are Pulsars?



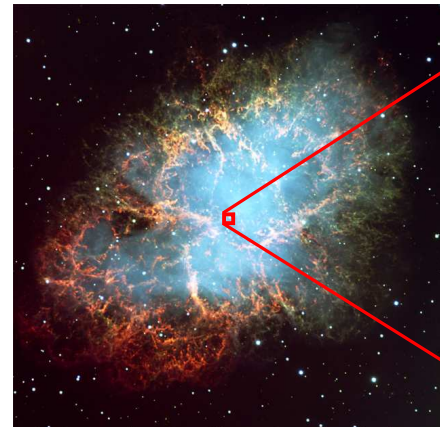
- Intense beams of radiation emanate from regions near the north and south magnetic poles of a neutron star
- These beams are produced by streams of charged particles moving in the star's intense magnetic field
- As the Pulsar gives energy to its surroundings, it slows down.
- The periods increase (few billionths of a second each day)



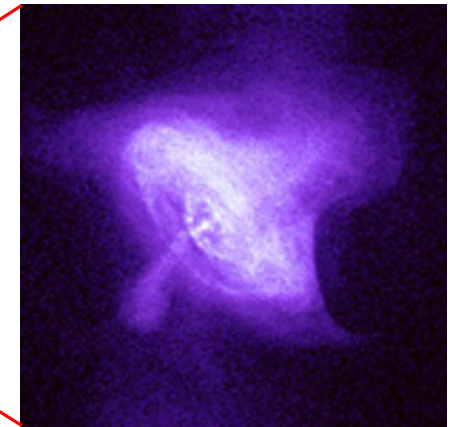
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# Crab Nebula – Remnant of the Supernova of 1054



Optical - ESO

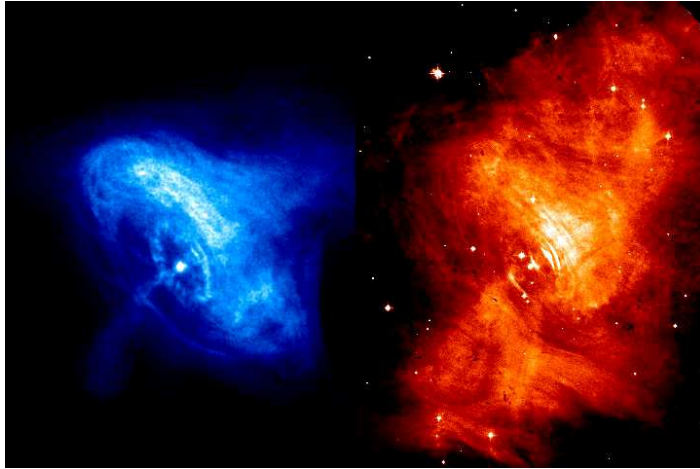


X-ray - Chandra

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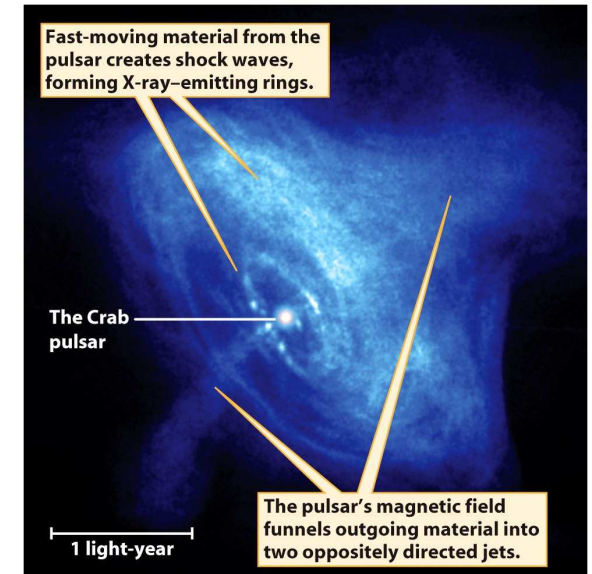
# Crab Nebula – Remnant of the Supernova of 1054



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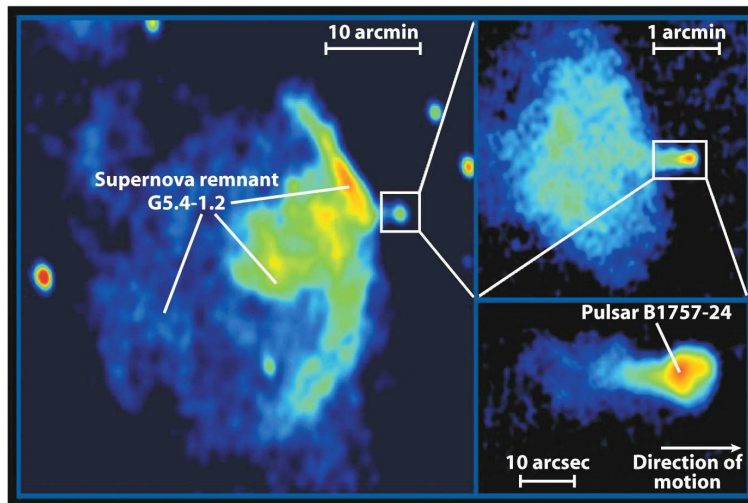
# The Crab



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# Escaping Pulsars



- Some Pulsars are ejected during the supernovae.
- Can outrun the explosion.
- This one is 600 km/s

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# When Neutron Degeneracy Isn't Enough



- Maximum neutron star mass
  - About  $3.0 M_{\odot}$
  - Original star around  $30 M_{\odot}$
- Beyond this mass, neutron degeneracy cannot stop gravity
- Nothing left to stop, so total collapse– gravity rules!
- **A black hole**
  - $v_{\text{esc}} > c$

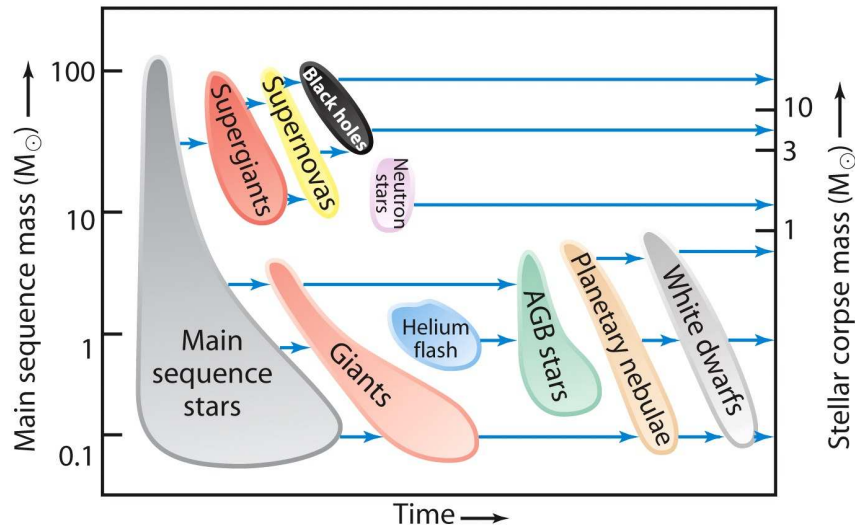


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## Stellar Evolution Recap



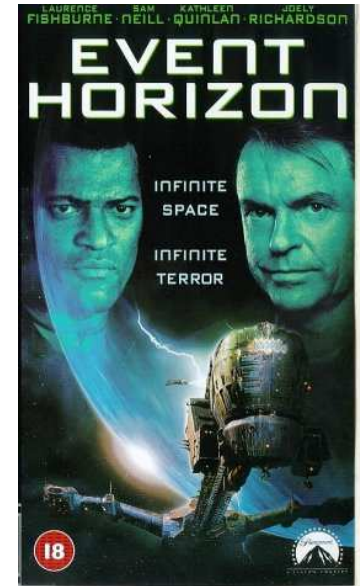
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## Black Holes



- Black holes inspire fear, awe, uncertainty, and bad science fiction
- Many people think that black holes are dangerous
  - That they suck matter in like “cosmic vacuums”
- Black holes follow the same laws of gravity as everything else



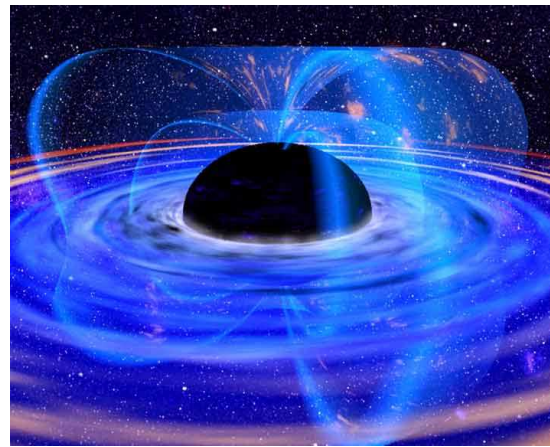
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## Physics of Black Holes



- Black holes are simple, yet strange objects
- Intense gravity due to compactness
- Newton's Laws cannot describe what happens in the presence of such an intense gravitational field
- We need Einstein's *Theory of Relativity*



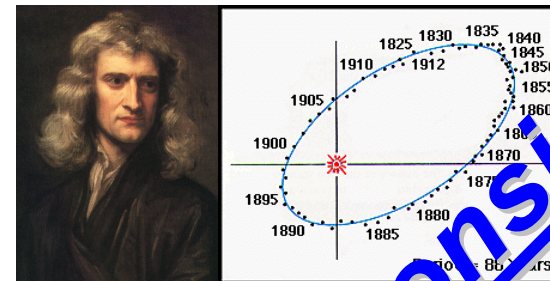
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## Two Threads of Thought in Physics up to 1900



**Mechanics**  
(Newton's Laws)



All motion is relative  
No speed is special

**Electromagnetism**  
(Maxwell's Equations)



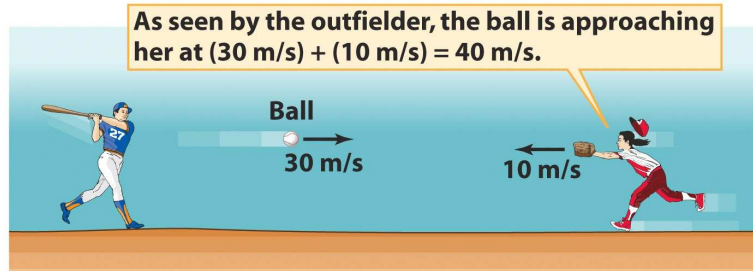
The speed of light is the  
**same** for all observers

**Inconsistent**

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## Why Galileo and Maxwell Can't Both Be Right



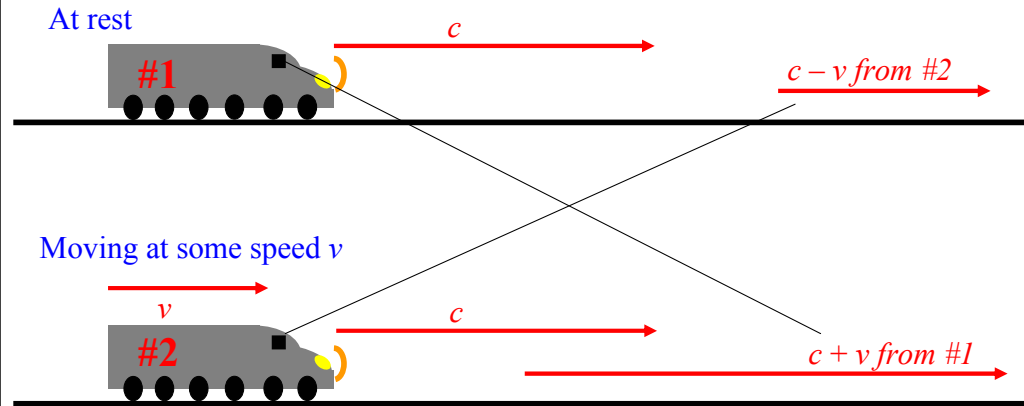
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## Why Galileo and Maxwell Can't Both Be Right



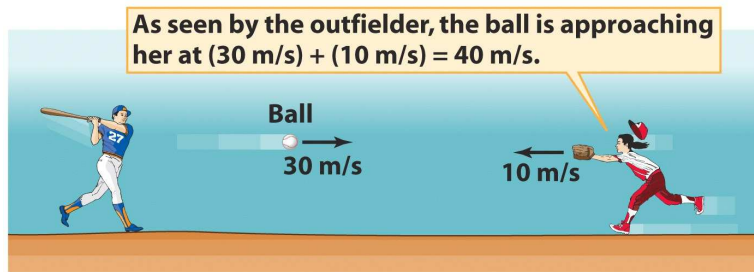
Consider two locomotives emitting light from their headlamps:



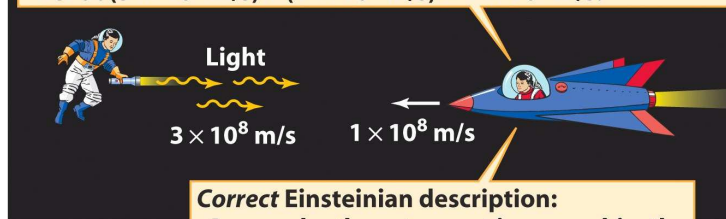
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## Why Galileo and Maxwell Can't Both Be Right



**Incorrect Newtonian description:**  
As seen by the astronaut in spaceship, the light is approaching her at  $(3 \times 10^8 \text{ m/s}) + (1 \times 10^8 \text{ m/s}) = 4 \times 10^8 \text{ m/s}$ .



**Correct Einsteinian description:**  
As seen by the astronaut in spaceship, the light is approaching her at  $3 \times 10^8 \text{ m/s}$ .

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## Why Galileo and Maxwell Can't Both Be Right



So the speed of light can't be the same for everyone if Galileo – and our intuition – are right. But Maxwell says it is constant!

Something must happen. And what must happen for Galileo and Maxwell to be both right, is that there is a modification of time and distance. Remember

$$\text{speed} = \frac{\text{dist}}{\text{time}}$$

Distance and time become relative to the observer.

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## Approaching the “c”



- Time dilation – Moving clocks run slow.
- Length contraction – Moving objects contract along direction of motion.
- Mass increase – moving clocks get more massive



<http://www.richard-seaman.com/Travel/Japan/Hiroshima/AtomicBombMuseum/IndividualArtifacts/>

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## Gamma



The factor by which all of these changes occur is called “gamma”

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

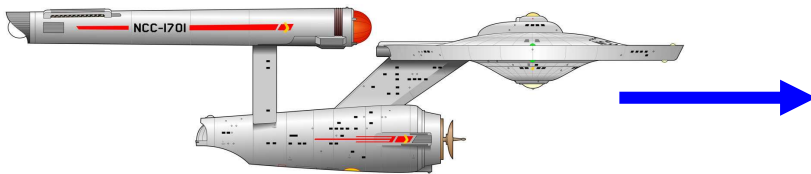
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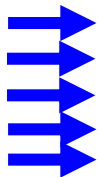
## Counterintuitive Result #1



**Moving objects appear shorter in the direction of relative motion (Lorentz contraction)**



Fraction of the speed of light	% of original length
0.00	100%
0.001	99.99995%
0.01	99.995%
0.1	99.5%
0.5	86.6%
0.9	43.6%
0.99	14.1%

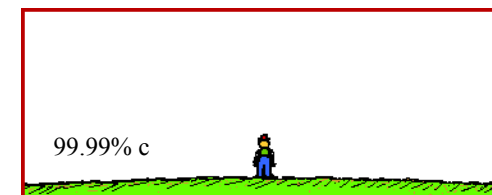
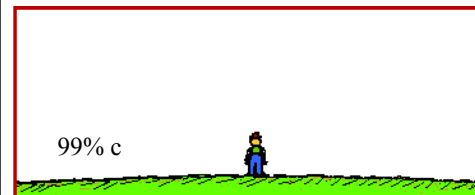
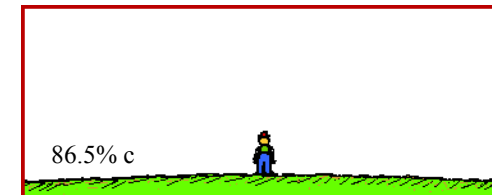
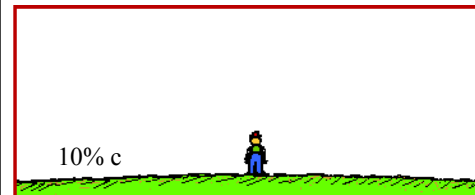


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## Counterintuitive Result #1



**Moving objects appear shorter in the direction of relative motion (Lorentz contraction)**



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<http://www.physicsclassroom.com/mmedia/specrel/lc.html>

## Length Contraction



- This effect has some benefits:
  - Outside observers will see that the length of the spaceship has shrunk.
    - This doesn't really help or harm us
  - But, from the astronaut point of view, the entire universe outside their window has shrunk in the direction of motion, making the trip shorter!
- It's all relative.

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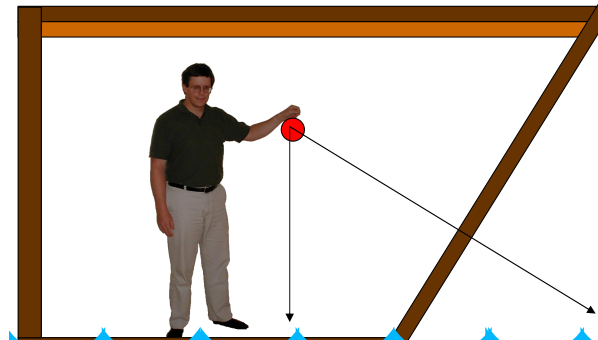
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## Galileo's ship thought experiment



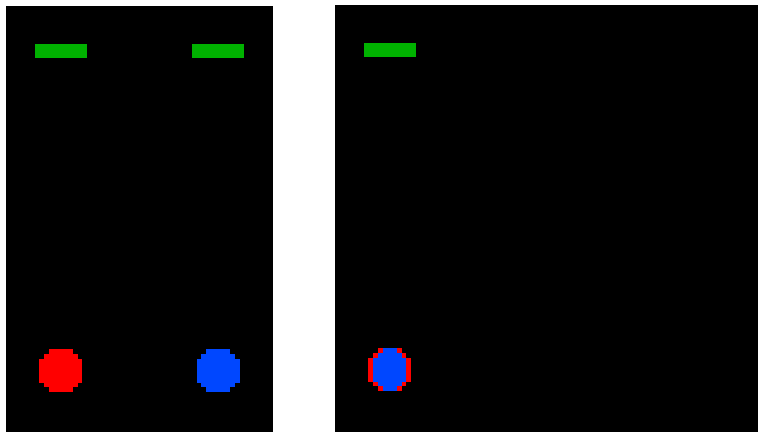
No experiment within the ship's cabin can detect the ship's motion if the ship moves in the same direction at a constant velocity. This is still true, even when considering the speed of light.

Frame of reference 1: We are moving with the ship



Frame of reference 2: We are at rest with respect to the water

## If the Ball is light?



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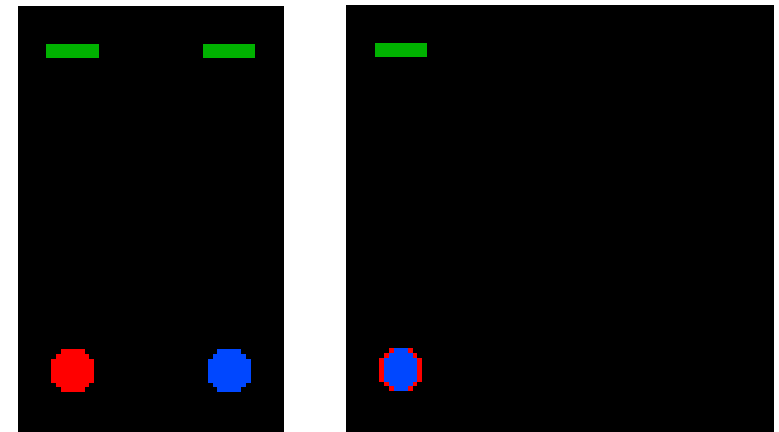
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A. Hamilton (Colorado)

## Counterintuitive Result #2



***Time appears to advance more slowly for moving objects (time dilation)***



Red & blue at rest

Blue moving to right

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## Time Dilation



- The effects of time dilation are curious but not prohibitive for space travel
  - Astronauts will age less than the Earth-bound folks waiting for the return. Can spoil the homecoming celebrations.
  - The faster you go, the bigger difference between astronaut time and Earth time
- Example: Trip to the center of the Galaxy and back. Accelerate at 1g for the first half and decelerate for second half and you can go 30,000 ly in 20 years! But more than 30,000 years has elapsed on Earth!

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## Mass Increase

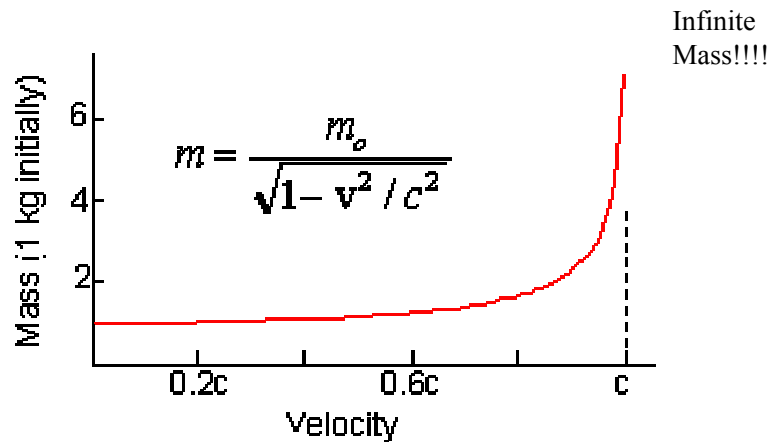


- The increase of effective mass (and kinetic energy) with velocity makes acceleration and deceleration more difficult if you intend to travel close to “c”
  - This translates to very costly starflight in terms of required energy.
  - And now the interstellar dust that you strike at relativistic speeds appears as larger mass.
  - For 99% speed of light travel, 5.5 meters of shield would erode every year.

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## The gamma factor and mass:



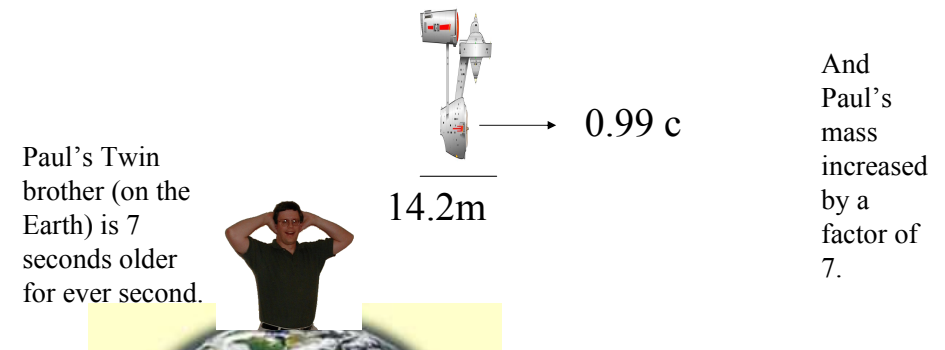
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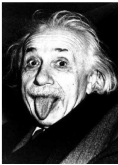
## So, what does that mean?



- If you're on a 100m spaceship going near the speed of light (.99 c), the spaceship would look 100m long, but someone on the Earth would observe the spaceship to only be 14m long.
- As you speed by the Earth your clock would tick 1 second, and an observer would tick about 7 seconds— $\gamma = 7$ .







## Special Relativity Summary



**All motion is relative, except for that of light.  
Light travels at the same speed in all frames of reference.**

**Objects moving close to the speed of light appear to shrink in the direction of travel.**

**Time appears to advance more slowly for objects moving close to the speed of light.**

**Mass of the moving object appears to rapidly increase as an objects moves close to the speed of light.**

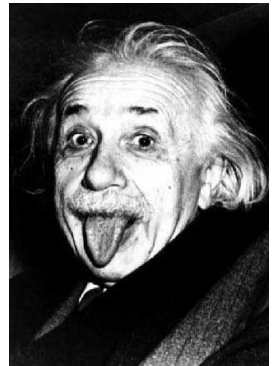
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## The Theory of General Relativity



- Recall Galileo: for free body motion is straight line, constant speed
- Important to note that ALL free bodies move this way. straight line, constant speed, INDEP of size, mass
- Q: Why?
- A: That's the way it is!
- Q: Be more specific: that's the way WHAT is?
- A: Einstein: that's the way space and time are if nothing else going on (no forces) space and time constructed so that free bodies move in straight lines at constant speed independent of nature of the object
- **That's the way space and time are**



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## Free Fall



- Recall Galileo's experiment.
- The objects in the gravity field, move independent to mass or even object.
- For Newton, the object mass cancels out of the gravity equation.



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## The Theory of General Relativity

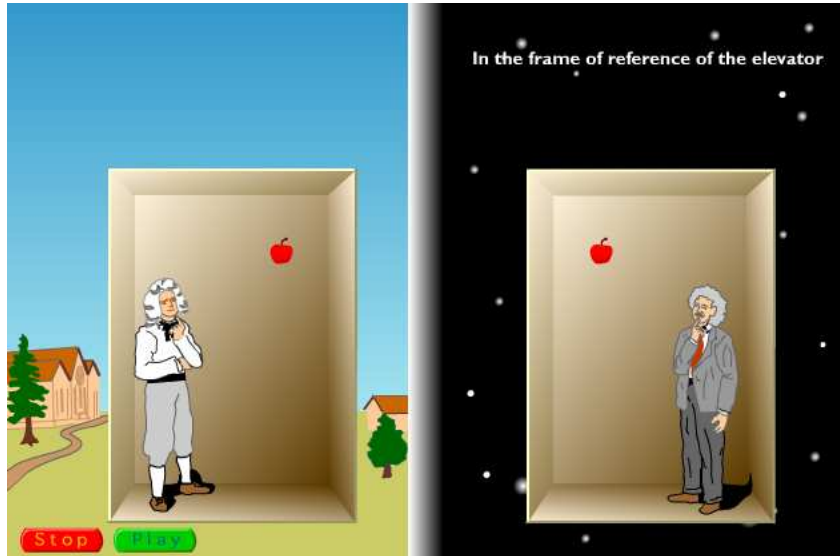


- Since objects move the same in a gravity field, INDEPENDENT of object, then gravity is not a force, but also a feature of space-time!
- Objects do their best to move in a straight line.
- **Newton:** Matter causes force (gravity)  
⇒ particles follow curved lines in “flat” (Euclidean geometry) space
- **Einstein:** Matter causes spacetime to be “curved”  
⇒ particles follow straight lines (“geodesics”) in curved space

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# The Theory of General Relativity



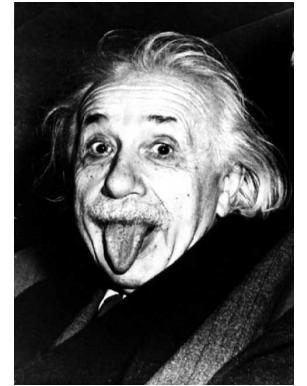
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# The Theory of Relativity



- Einstein's Theory of Relativity tells us how gravity works
  - Space and time are not distinct
  - They are bound together in 4-dimensional **spacetime**
  - Matter tells spacetime how to curve
  - Curved spacetime tells matter how to move



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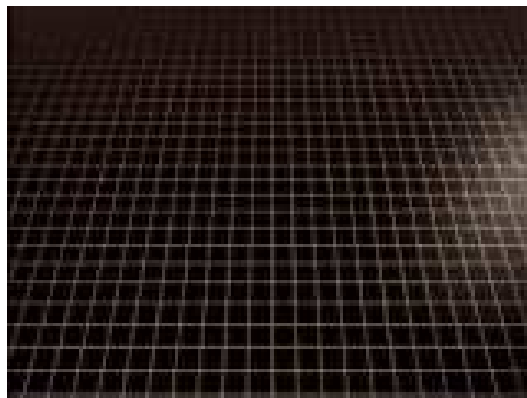
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## Curved Spacetime



- No matter = Flat Spacetime
- Massive object = Dent in Spacetime

- Everything follows curvature of spacetime including light (photons)



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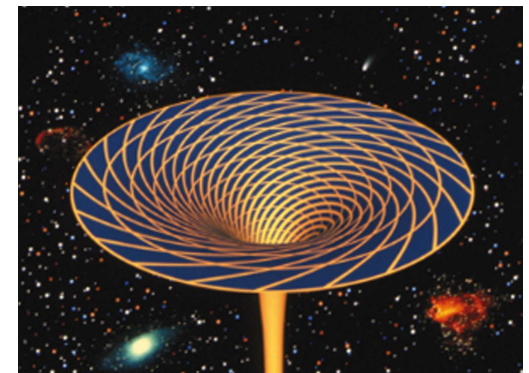
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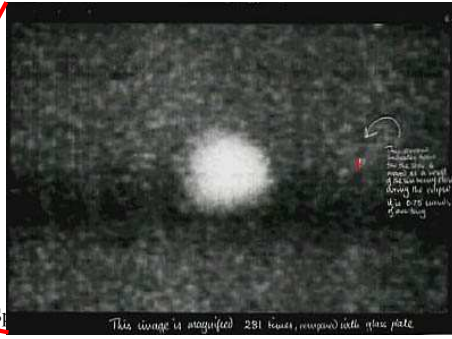
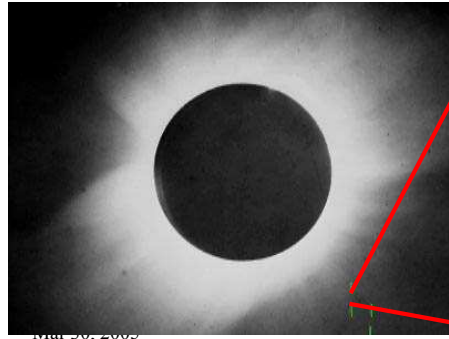
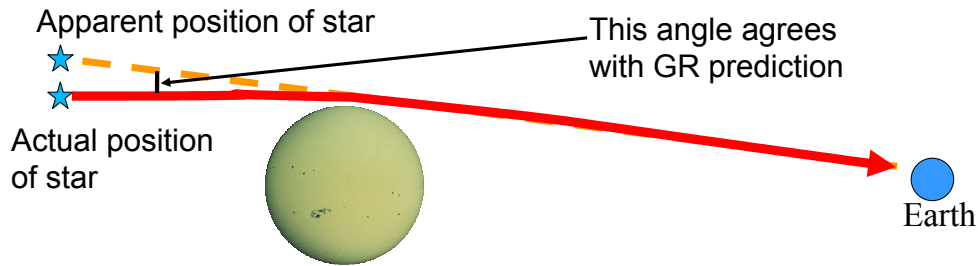
- Everything follows curvature of spacetime including light (photons)



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## Eddington and the 1919 Eclipse



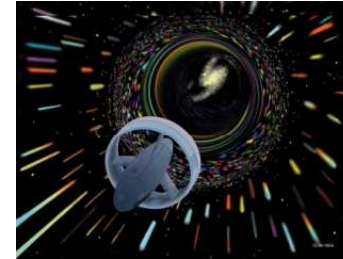
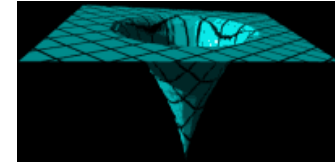
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This image is magnified 281 times, compared with glass plate.

## General relativity



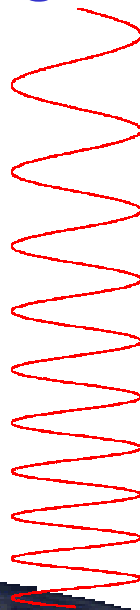
- Gravitational fields can also change space and time
  - A clock runs more slowly on Earth than it does in outer space away from any mass, e.g. planets.
- Einstein revealed that gravity is really 'warped' space-time.
- A black hole is an extreme example.



## Gravity Also Redshifts Light

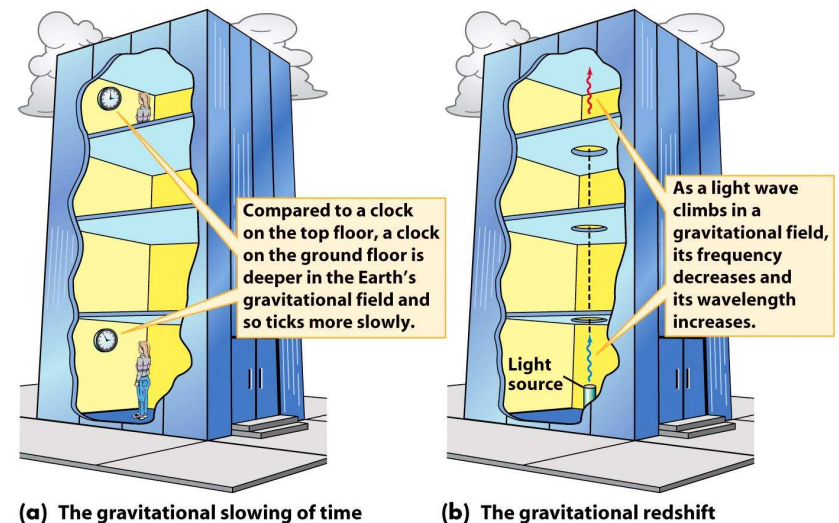


- Light loses energy as it climbs out of a gravitational field so its wavelength increases (redshift).
- As with light bending, the effect is small but measurable.



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## Gravity Redshifts Light



(a) The gravitational slowing of time

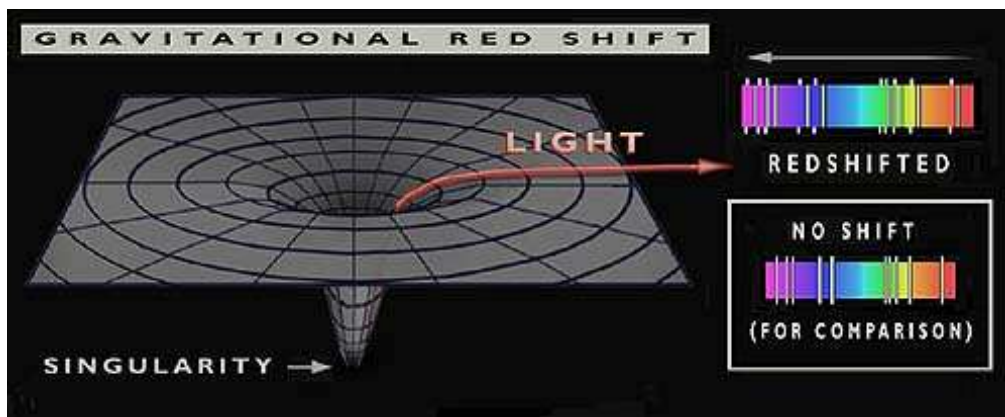
(b) The gravitational redshift

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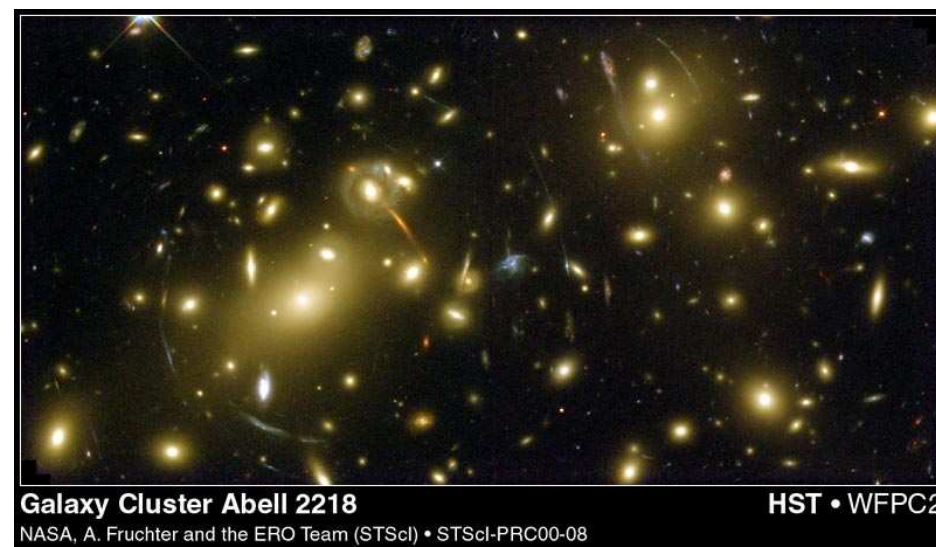
## Gravity Redshifts Light



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## Einstein Lens



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<http://antwrp.gsfc.nasa.gov/apod/ap000201.html>

## To each ones own



- From each other's view, things would appear normal, time and colors.
- That's the theme of relativity: if I only measure things nearby, I see normal things. Weirdness only happens at distance from observer or moving at great speeds relative to observer.

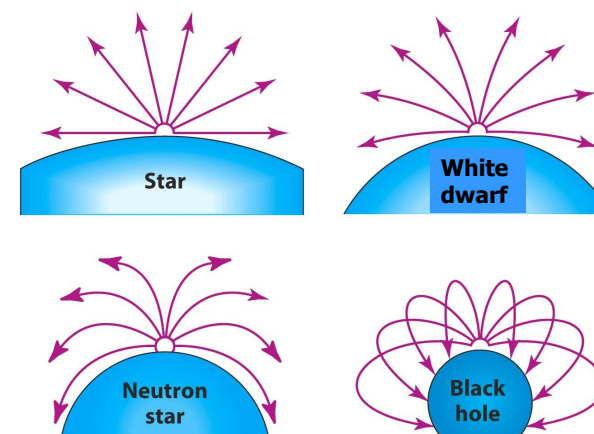
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## Back to Black Holes



- When matter gets sufficiently dense, it causes spacetime to curve so much, it closes in on itself
- Photons flying outward from such a massive object arc back inward!
- Neither light or matter can escape its gravity, it is a **black hole!**



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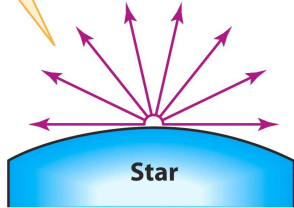
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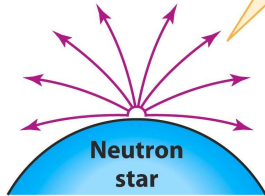
## Black Holes



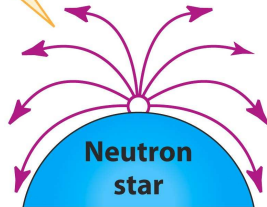
1. A supergiant star has relatively weak gravity, so emitted photons travel in essentially straight lines.



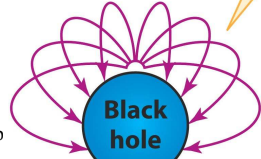
2. As the star collapses into a neutron star, the surface gravity becomes stronger and photons follow curved paths.



3. Continued collapse intensifies the surface gravity, and so photons follow paths more sharply curved.



4. When the star shrinks past a critical size, it becomes a black hole: Photons follow paths that curve back into the black hole so no light escapes.

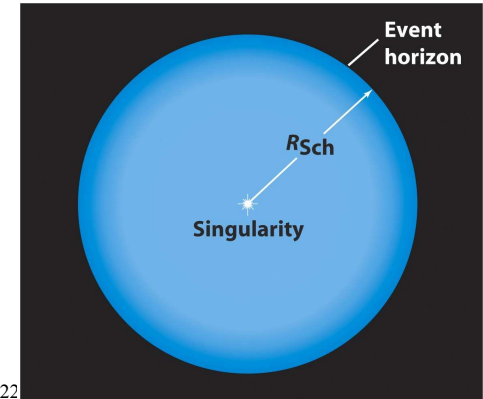


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## Black Hole



- The matter in a black hole collapses to a point – called a **singularity**
- A black hole is separated from the rest of the Universe by a boundary, the **event horizon**
- Nothing can escape from within its radius
- This radius is called the Schwarzschild radius



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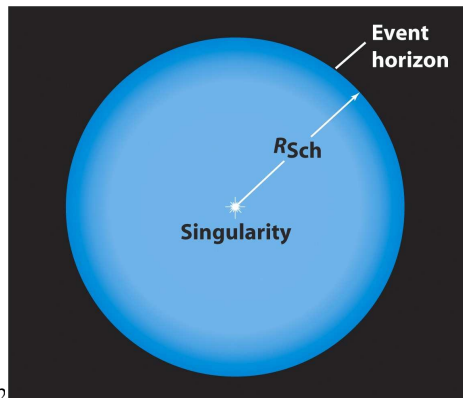
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## Black Hole



- The Schwarzschild radius is
- More massive black hole = larger the event horizon
  - $R_{Sch} = 3 (M/M_{sun}) \text{ km}$
  - If mass of an object is in a space such that  $< R_{Sch}$  then it is a BH
  - For Earth  $R_{Sch} = 1 \text{ cm}$
- The radius of no return
- Cosmic roach hotel

$$R_{Sch} = \frac{2GM}{c^2}$$



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## Thought Question



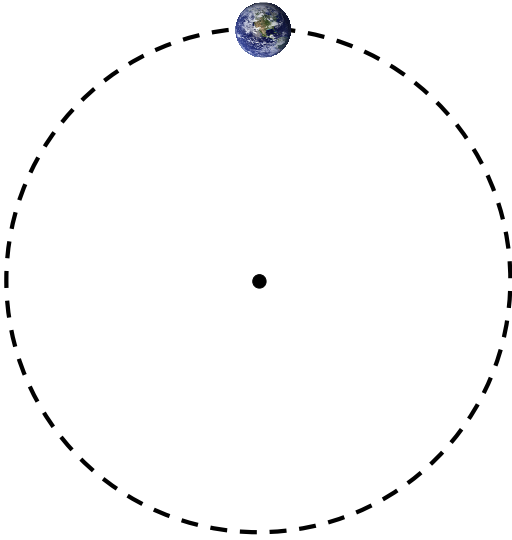
What do you think would happen to the Earth if the Sun collapsed into a black hole?

- Fall in directly
- Slowly spiral in
- Stay in its orbit
- Slowly spiral away
- Fly away in a straight line

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*Well outside of a black hole –  
It looks just like any other mass*



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