



- Homework #1 is due Friday at 11:50am!
- Planetarium shows are getting full.
- Solar Observing starts next Monday!
- Nighttime observing starts in < 2 weeks.

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## Outline



- Dance of the Planets– Planetary motion
- Tycho Brahe and his observations
- Johannes Kepler and his interpretation
  - 3 laws
    - Orbits are on ellipses
    - An orbit sweeps out equal area in equal time
    - An orbital period is related to the semimajor axis of the orbit.

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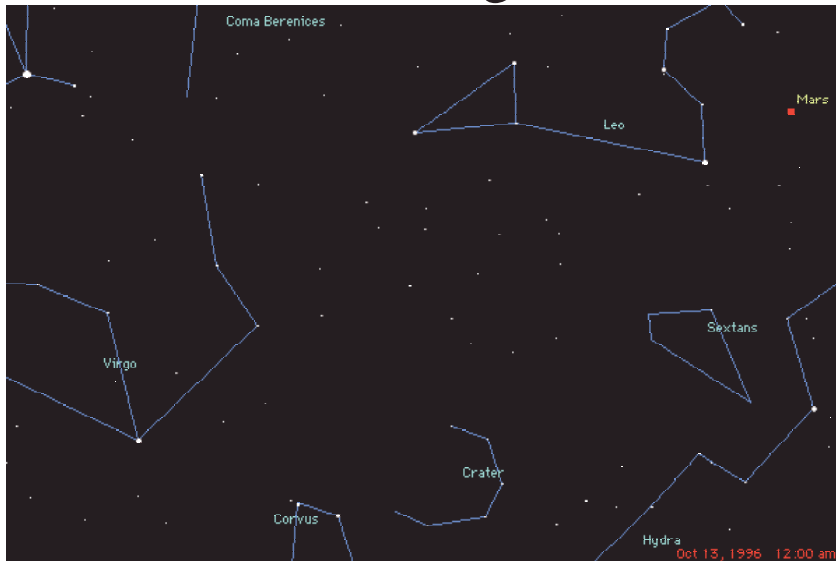
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## Mars Retrograde



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## Tycho Brahe (1580)



- Spent his life producing a catalog of carefully observed stars and planets using “state-of-the-art” observatory
- No telescopes!
- Yes, had a metal nose, but did not die from burst bladder



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# Uraniborg



Accurate measurements to about 1 minute of arc (1/15 the diameter of the moon)



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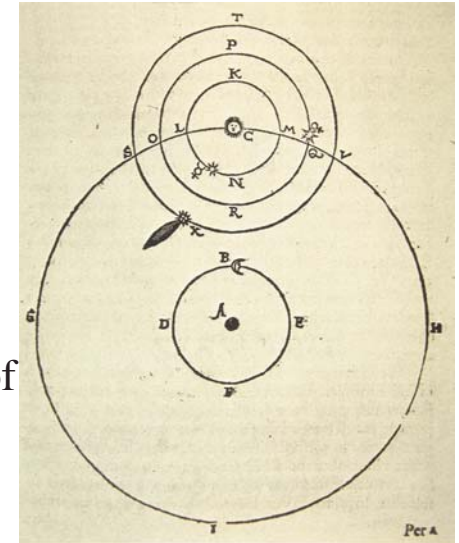


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# Tycho's Model



- Data did not fit with geocentric view
- Developed a new model that tried to keep the geocentric viable– but too complicated
- Thought that the Laws of nature demanded a geocentric cosmology



[http://es.rice.edu/ES/humsoc/Galileo/People/tycho\\_brahe.html](http://es.rice.edu/ES/humsoc/Galileo/People/tycho_brahe.html)

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# Johannes Kepler (1600)



- Tycho's assistant in Prague
- After Tycho's death, succeeded Tycho's position and had access to the excellent data
- How to fit the heliocentric model to data of Mars?



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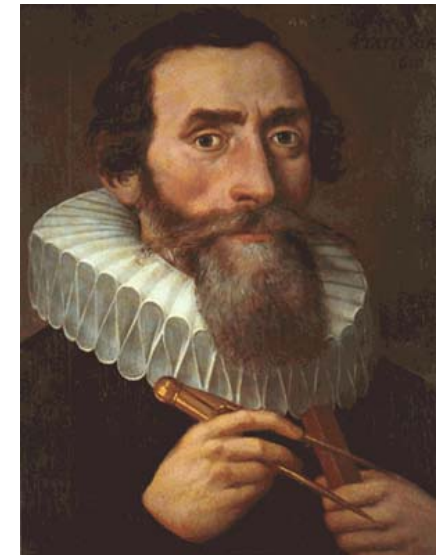
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# Johannes Kepler (1600)



There was a problem. The data was so good that it could not be fit with the heliocentric model if only circles were used.

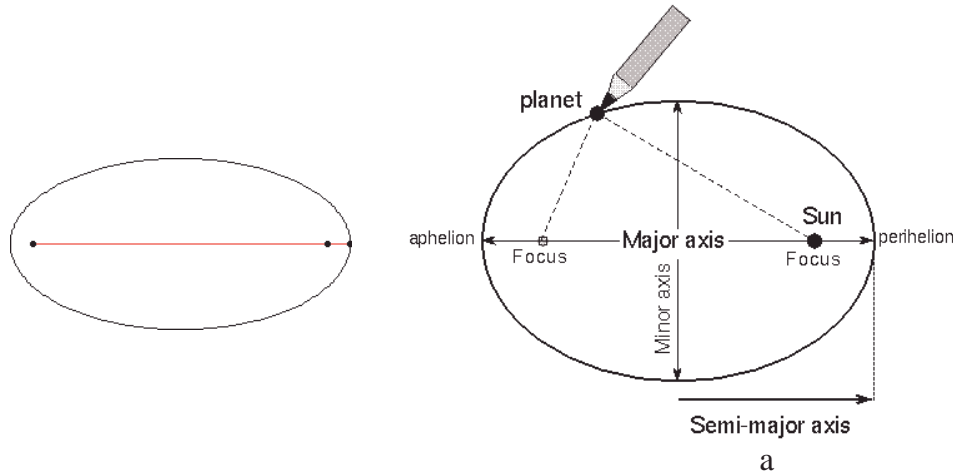
Then, he began to work with the ellipse.



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# Kepler's 1<sup>st</sup> Law: Orbits of planets are ellipses with the Sun at one focus



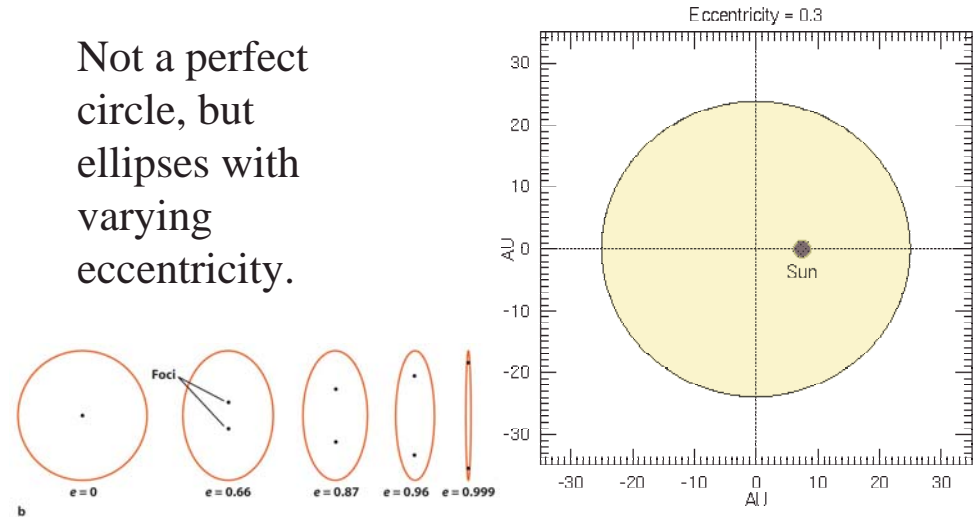
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# Orbits of planets are ellipses with the Sun at one focus



Not a perfect  
circle, but  
ellipses with  
varying  
eccentricity.



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## Implications

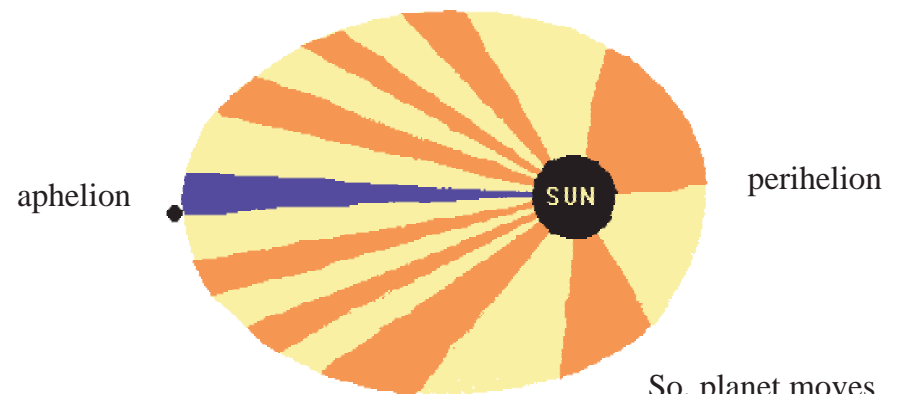


New Twist— even the Sun isn't at the center of  
the solar system now. How does that change  
our view of the Universe and our place in it?

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## Kepler's 2<sup>nd</sup> law: The Line that connects the planet to the Sun sweeps out equal areas in equal time



So, planet moves  
faster at perihelion.

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## Kepler's 2<sup>nd</sup> law:

The Line that connects the planet to the Sun sweeps out equal areas in equal time

Example: Note the inequality of the seasons: spring & summer have 93 days, autumn has 90 days, and winter has 89 days

Earth moves faster during autumn and winter (when it's closer).

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## Kepler's 1<sup>st</sup> and 2<sup>nd</sup> Law

<http://csep10.phys.utk.edu/astr162/lect/binaries/visual/kepleroidframe.html>

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## Astronomical Unit

Now we need to define the often used term Astronomical Unit or AU. This is simply the average distance of the Earth to the Sun, which is also about the Earth's Semi-Major axis. It is equal to  $1.5 \times 10^8$  km. Then, we can say that Jupiter for example is 5.2 AU from the Sun, or 5.2 times the distance away as the Earth. Just an easier unit.

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## Kepler's 3<sup>rd</sup> Law:

The squares of the orbital sidereal periods of the planets about the Sun are proportional to the cubes of the orbital semimajor axes

Planet	P (yr)	a (AU)	P <sup>2</sup>	a <sup>3</sup>
Mercury	0.24	0.39	0.06	0.06
Venus	0.61	0.72	0.37	0.37
Earth	1.00	1.00	1.00	1.00
Mars	1.88	1.52	3.5	3.5
Jupiter	11.86	5.20	141	141
Saturn	29.46	9.54	868	868

$$P^2 = a^3$$

$$P \times P = a \times a \times a$$

Where P is in years and a is in AU.

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## Kepler's 3<sup>rd</sup> Law



Works for Satellites, Moons, Comets,  
Asteroids, Binary Stars... (a caveat)

Halley's Comet returns every 76 years. What  
is its semimajor axis?

$$P^2 = a^3 \text{ or } a^3 = 76^2 = 5776$$
$$\text{so } a = (5776)^{1/3} = 18 \text{ AU}$$

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## Kepler's Laws



The farther away from the Sun, the longer it  
takes for the planet to orbit AND the slower  
it's average orbit speed.

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## Kepler's Laws



<http://csep10.phys.utk.edu/astr162/lect/binaries/visual/kepleroldframe.html>

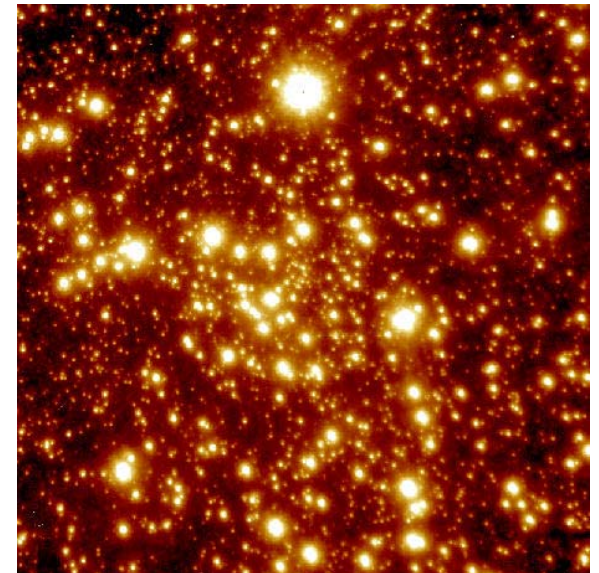
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## Kepler's Laws: A Black Hole



A group in  
Munich is using  
Kepler's Laws to  
determine the  
mass of the black  
hole in the center  
of our galaxy  
using deep near-  
infrared  
observations.



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[http://www.mpe.mpg.de/www\\_ir/GC/intro.html](http://www.mpe.mpg.de/www_ir/GC/intro.html)

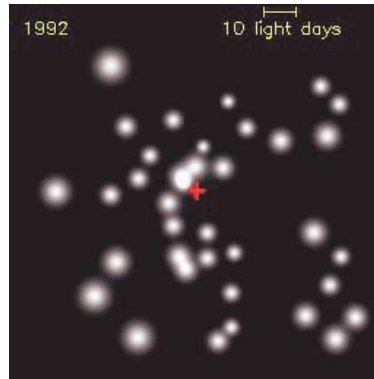


# Application of Kepler's Laws



By using the deepest images of stars toward the Galactic Center, they have been able to detect a full orbit of a star with a period of 15.2 years and Semimajor axis of 950 AU.

That means that the black hole is about 2.6 million solar masses!!!



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# Galileo (1610)



First to systematically use the telescope (but did not invent it).

- Moon has mountains and valleys
- Milky Way consists of faint stars
- Saturn is elongated
- Venus shows phases
- Jupiter has moons (now called Galilean moons)

Wow! Big stuff. The moons of Jupiter did not orbit the Earth!



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# Galilean Moons



Geocentric Cosmology was still preferred model of the Universe and Galileo was declared a Heretic and spent years under house arrest.



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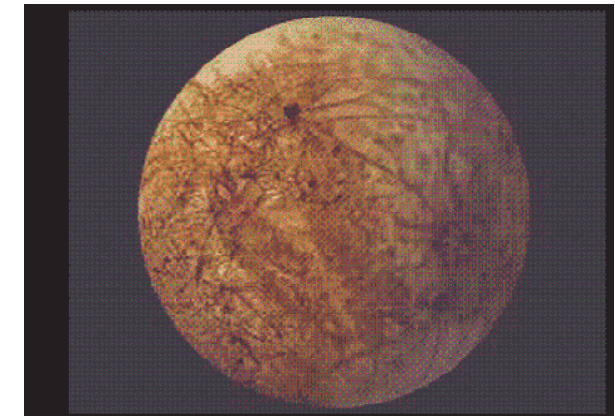
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<http://photojournal.jpl.nasa.gov/catalog/PIA01299>

# Europa



What was the problem? Galileo's observations directly challenged the Geocentric view that was held by the church. And there was still no why.

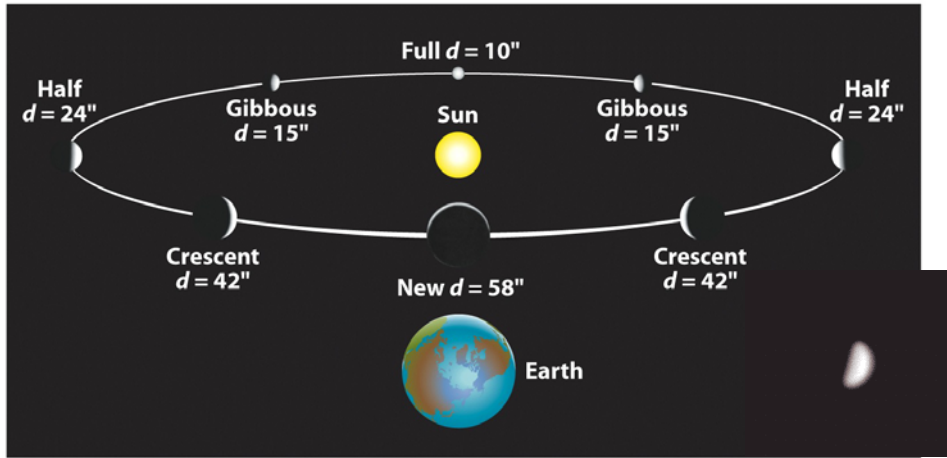


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<http://antwrp.gsfc.nasa.gov/apod/ap010116.html>

## The Phases of Venus



Could not be explained with the Geocentric model

<http://www.calvin.edu/academic/phys/observatory/images/venus/venusb.html>

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## Phases of Venus



Compare the Heliocentric to Geocentric models to explain the phases of Venus.

<http://www.astro.ubc.ca/~scharein/a310/SolSysEx/phases/Phases.html>

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## Galileo (1610)



- Disproved Ptolemaic system
- Rome bullied him into recanting (cleared in 1992)
- Now we understand the motions and the fact that the solar system **MUST** be Heliocentric, but now we need a reason why?

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## Kepler's Laws



Kepler discovered these patterns in nature by using the data that Tycho collected, **BUT** the world had to wait until someone could understand the Natural Law that predicts Kepler's Laws.

The real problem: On Earth we're used to things that move but always come quickly to a rest. Why didn't the planets stop?

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# Isaac Newton



- Gave us a reason why-- GRAVITY.
- Developed fundamental laws of nature.
- Kepler's 3<sup>rd</sup> law now became a way to probe the structure of the Universe!

